YFS MC Approach to Precision Theory for Collider Physics: Origin, Development, and Outlook

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In collaboration with S. Jadach¹, W. Placzek, M. Skrzypek, Z.A. Was, S.A. Yost, and A. Siodmok



¹Deceased.

In Memory of Prof. Stanislaw Jadach

 Sadly, Prof. Stanislaw Jadach passed away suddenly on Feb. 26, 2023 CERN COURIER: May - June Issue, 2023:

A leading light in radiative corrections

physicist, died on 36 Pebruary at the age of 25. His foundational contributions to the physics programmes at LSP and the LHC, and for the approach Purple Circular Collider at CRRN, have significantly beloed to advance the field of elementary porticle physics and its future aspirations.

Burn In Council, Poland, Indach graduated in committee and solvered in shape of the property of the committee of the co University. There, he also defended his doctorate. received his habilitation degree and worked until 1992. During this period, while: partly under marrial loss to Reland, Indach took trips to Leiden, Faris, Lundon, Stanford and Knoxyllie, and formed collaborations on precision theory calculations based on Monte Carlo every-generatormethods in 1933 he moved to the institute - the physics programmes at LUP and the LUC. of Maclesc Physics Polish Academy of Sciences

1904, he worked until his death. Along series of publications and computer pro- action processes. grammes for re-current percurbative transferd

Most of the analysis of LEP data was based instore developed a new constrained Markovian. Benede Wood Broker Univerdry.



(HAS) where, receiving the ritrle of professor in —earlustedy on the novel calculations provided by Jadach and his colleagues. The most important repoble for experiment, in taky ladach solved. UAPs, ladach and co-workers intelligently conthat publish in a dirate-author expert, inspired bined securing first-order calculations for the BADCOR conference. by the classic work of Vennie. Practically and production and decay processes to achieve the Sturn featuring answers indistricts in method for tracessary of the control accuracy, bypassing mentor. Modest, gentle and sensitive, be did any number of photons. It was widely believed. The need for full first-order calculations for the not judge or impose. He never refused requests that soft-phases approximations were restricted. Bur-fermion process, which were unfeatible as: and always had time for others. His professional to many photons with very low energies and the time Contrary to what was deemed possible, innovining was impressive, we knew almost that it was impossible to relate, consistently, the Jadach and his colleagues achieved calculative everything about QRD, and there were few distributions of one or two energetic phonons. Gonzalez simultaneously take into account (\$0) other topics in which he was not at least knowto those of any turniber of soft photons, ladach radiative corrections and the complete agin- edgeable, etc entition beyond physics was and his colleagues solved this problem in their gots correlation effects in the production and equally extensive. He is already profus notice and papers in polic for differential cross sections. Geomofree tax lectors, Health and success in dearly reliated. and laser in 1999 at the level of spin amplitudes. The 1970s in novel simulations of strong inter-

Among other novel require, he and his collabo- Institute of Muclear Physics and

usebackward evolution and predefined parton. distributions, and proposed anew method, using a "physical" factorisation scheme, for combining a hard process at neut-to leading order with a partor caucade, much simpler and more efficient than alternative methods.

Jadach was already updating his LEP-era calculations and software towards the increased precision of RCC-ee, and is the co-editor and co-surface of a major paper delineating the need for new theoretical calculations to meet the protosefoolider's physics needs. He co-organised and participated in many physics workshops ar CRR and in the preparation of comprehensive reports, starting with the famous 1989.

LSP THE ROW REPORTS. Jadach, a member of the Polish Academy of Arts and Sciences (PAAS), received the most prestigious awards in physics in Poland, the Marte Skindowska-Curie Prize (PAS), the Prior to UK all calculations of radiative curpersons were based on first- and, interpartially. Shabba scattering, the production of legion and prize of the Minister of Science and Higher second-order reality. This limited the theoret- quark pains, and the production and decay of . Mucadam for lifetime orient fic achievements. leaf precision to the VM level, which was unac - W and I honor pairs. For the W-pair results at He was along on-intransr and permanentmenther of the innernational advisory board of the

Stanielaw (Stassek) was a wonderful man and

Wheter Pleasek logistic sign University, After LRP, Indach turned to LRC physics. Made Skutypekond Shignless Was





OUTLINE

- Introduction
- Recapitulation of YFS Exact Amplitude-Based Resummation
- Approach to Precision Collider Physics: LHC, FCC, CPEC, CPPC, ILC, CLIC
- Improving the Collinear Limit in YFS Theory
- Outlook





- How did we get started on our YFS journey?
 - 1986 MarkII Radiatvie Corrections Meeting Organized by G. Feldman at SLAC:
 - Preparation for 'Precision Z Physics' at SLC: Did not turn-on until 1989, MKII Observed ~ 750 Z's
 - Staszek and I met in this Meeting.
- There was a No-Go Belief: Jackson-Scharre Naive Exponentiation-Based Methods – Nothing Better
- We started discussing whether the approach of Yennie, Frautschi, and Suura could do better –
 - It worked at the level of the amplitudes:
 - Could a MC realize all that?
 - Was renormalization group improvement alive?



- Discussion aided by my participation in the 27th Cracow School of Theoretical Physics
 - Long walks in the mountains
 - Staszek had already written MPI-PAE-PTH-87-6: "Yennie-Frautschi-Suura soft photons in Monte Carlo event generators"
 - We presented RG Improvement at the School.
- Proof of Principle:
 - "Exponentiation of Soft Photons in the Monte Carlo: The Case of Bonneau and Martin," University of Tennessee preprint UTHEP-88-0101, and SLAC-PUB-4543, Phys. Rev. D38, 2897 (1988)
 - "Multiphoton Monte Carlo for Bhabha Scattering at Low Angles," University of Tennessee preprint UTHEP-88-11-01, 1988, Phys. Rev. D40, 3582 (1989)"



- This was followed by "YFS2-The Monte Carlo for Fermion Pair Production at LEP/SLC with the Initial State Radiation of Two Hard and Multiple Soft Photons", CPC 56 (1990) 351
- ⇒ KORALZ 3.8, BHLUMI 2.01
- "Final State Multiple Photon Effects in Fermion Pair Production at SLC/LEP," UTHEP-91-0903, Phys. Lett. B274 (1992) 470.
- ⇒ KORALZ 4.0, BHLUMI 4.04, ...





- More applications followed:
 - BHWIDE, BHLUMI 2.30, YFSWW3, KORALW, KORALW&YFSWW3
- CEEX: Proof of Principle
 - "Coherent Exclusive Exponentiation CEEX: The Case of the Resonant e+e- Collision," CERN-TH-98-253, UTHEP-98-0801; Phys. Lett. B449, 97 (1999)
- ⇒ KKMC, KKMC 4.22, KKMC-ee, KKMC-hh, KKMC-ee (C++), ...
- Applications: SLC, LEP1 and LEP2, BaBar, BELLE, BES, Φ-Factory, LHC
- Applications: TESLA, ILC, CLIC, FCC, SSC-RESTART, CEPC, CPPC, ...





- The Future of Precision Theory: Dictated by Future Accelerators – FCC, CLIC, ILC, CEPC, CPPC, ...
- \bullet Using FCC as an example, factors of improvement from \sim 5 to \sim 100 are needed from Theory
- Resummation is a key to such improvements in many cases:
 - Today, we discuss amplitude-based resummation following the YFS MC methodology made possible by Staszek's seminal contributions.
- YFS → 'no limit to precision'
- See 1989 CERN Yellow Book article by Berends et al.



 The Future of Precision Theory: Dictated by Future Accelerators – FCC, CLIC, ILC, CEPC, CPPC, ···

Gianotti: 1/10/23 -- von der Leyen: 7/16/25, EU MFF Moonshots -- FCC, clean aviation, ...

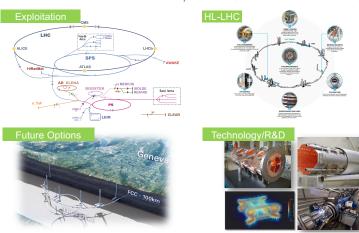


Figure: Future of CERN.



 The Future of Precision Theory: Dictated by Future Accelerators – FCC, CLIC, ILC, CEPC, CPPC, ...

Theory

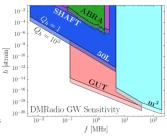
Some physics highlights:

Gianotti: 1/10/23

- Higher-order calculations of background processes for LHC, HL-LHC and future colliders
- Axion physics and, in particular, studies for using axion haloscopes to detect high-frequency gravitational waves through oscillating electromagnetic signals sourced by spacetime distortions (arXiv: 2202.00695)
- String Theory: Exploring the swampland and how its conjectures can reveal information on the energy scales of nature (arXiv: 2205.12293)
- Bounds on the energy growth of gravitational amplitudes (arXiv: 2202.08280)

Other activities:

- Full restart of scientific activities and visitor programmes after Covid.
- □ TH served as a focal point for the physics community to discuss ecofriendly practices for organising scientific events and business travel. These issues were discussed in a dedicated Theory Institute, named "Sustainable HEP"

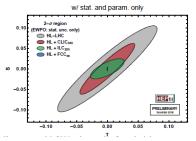




 The Future of Precision Theory: Dictated by Future Accelerators – FCC, CLIC, ILC, CEPC, CPPC, ... Grojean: 2/13/24

Tera-Z EW precision measurements.

- ▶ The target is to reduce syst. uncertainties to the level of stat. uncertainties. (exploit the large samples and innovative control analyses)
- ▶ Exquisite √s precision (100keV@Z, 300keV@WW) reduces beam uncertainties (EPOL)
- → ~50 times better precision than LEP/LSD on EW precision observables



Indirect sensitivity to 70TeV-scale sector connected to EW/Higgs

CG - 20 / 30

(For the impact of the theory uncertainties on the EW fit, see bonus slides)

Feb. 13, 2024





- YFS methods are exact in the infrared but treat the collinear logs perturbatively in the $\bar{\beta}_n$ residuals
- DGLAP-based collinear factorization treats the collinear logs to all orders but has a non-exact IR limit
- In this talk, we present some new results for precision collider physics based on the usual YFS methods.
- We then investigate improving the collinear limit of YFS theory.
- A Key Point: Exact Amplitude-Based Resummation Realized on Evt-by-Evt Basis – Enhanced Precision for a Given Level of Exactness: LO, NLO, NNLO, NNNLO,, essential for future precision physics as exemplified by CERN <=> Computer Algebraic Methods essential!



Recapitulation of Exact Amplitude-Based Resummation Theory

$$d\bar{\sigma}_{res} = e^{\text{SUM}_{\mathbb{R}}(\text{QCED})} \sum_{n,m=0}^{\infty} \frac{1}{n!m!} \int \prod_{j_1=1}^{n} \frac{d^3 k_{j_1}}{k_{j_1}}$$

$$\prod_{j_2=1}^{m} \frac{d^3 k'_{j_2}}{k'_{j_2}} \int \frac{d^4 y}{(2\pi)^4} e^{iy \cdot (p_1 + q_1 - p_2 - q_2 - \sum k_{j_1} - \sum k'_{j_2}) + D_{\text{QCED}}}$$

$$\tilde{\beta}_{n,m}(k_1, \dots, k_n; k'_1, \dots, k'_m) \frac{d^3 p_2}{p_2^0} \frac{d^3 q_2}{q_2^0}, \tag{1}$$

where *new* (YFS-style) *non-Abelian* residuals $\bar{\beta}_{n,m}(k_1,\ldots,k_n;k'_1,\ldots,k'_m)$ have n hard gluons and m hard photons.

$$\begin{split} \tilde{\tilde{\beta}}_{n,m} : \textit{FCCee} &- \textit{need exact } O\left(\frac{\alpha}{\pi}, \frac{\alpha}{\pi} L, \left(\frac{\alpha}{\pi}\right)^2, \left(\frac{\alpha}{\pi}\right)^2 L, \left(\frac{\alpha}{\pi}\right)^2 L^2, \left(\frac{\alpha}{\pi}\right)^3 L^2, \left(\frac{\alpha}{\pi}\right)^3 L^3, \left(\frac{\alpha}{\pi}\right)^4 L^4\right) \\ &\Rightarrow \textit{Computer Algebraic Methods} \iff \textit{Evaluation of Feynman Diagrams} \end{split}$$

BAYLOR



Review of Exact Amplitude-Based Resummation Theory

Here,

$$SUM_{IR}(QCED) = 2\alpha_s \Re B_{QCED}^{nls} + 2\alpha_s \tilde{B}_{QCED}^{nls}$$

$$D_{QCED} = \int \frac{d^3k}{k^0} \left(e^{-iky} - \theta (K_{max} - k^0) \right) \tilde{S}_{QCED}^{nls}$$
 (2)

where K_{max} is "dummy" and

$$\begin{array}{lll} B_{QCED}^{nls} & \equiv & B_{QCD}^{nls} + \frac{\alpha}{\alpha_s} B_{QED}^{nls}, \\ \tilde{B}_{QCED}^{nls} & \equiv & \tilde{B}_{QCD}^{nls} + \frac{\alpha}{\alpha_s} \tilde{B}_{QED}^{nls}, \\ \tilde{S}_{QCED}^{nls} & \equiv & \tilde{S}_{QCD}^{nls} + \tilde{S}_{QED}^{nls}. \end{array} \tag{3}$$

"nls" ≡ DGLAP-CS synthesization.

Shower/ME Matching: $\tilde{\bar{\beta}}_{n,m} \rightarrow \tilde{\bar{\beta}}_{n,m}$ See Ann. of Phys. **323** (2008) 2147 and references therein for more details.

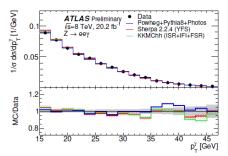


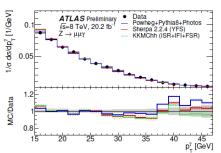


• (HL-)LHC:

 $\mathcal{K}\mathcal{K}$ MChh: Exact $\mathcal{O}(\alpha^2 L)$ CEEX EW corrections matched to a Herwig parton shower (built-in) or to any other shower via Les Houches files

(see also Liu *et al.*, to appear). \Rightarrow Recent ATLAS results on $Z\gamma$ production (Eur.Phys.J C **84** (2024) 195)





HL-LHC => Factor of ~ 10 smaller statistical errors => Test?



- (HL-)LHC:

 \(\mathcal{K} \) MChh: NISR shows effect of QED contamination in non-QED PDFs is below the errors on the PDFs:
- NISR –

$$\begin{split} \sigma(s) &= \frac{3}{4}\pi\sigma_{0}(s)\sum_{q=u,d,s,c,b}\int d\hat{x}\ dzdr\ \int dx_{q}dx_{\bar{q}}\ \delta(\hat{x}-x_{q}x_{\bar{q}}z)\\ &\times f_{q}^{h_{1}}(s\hat{x},x_{q})f_{\bar{q}}^{h_{2}}(s\hat{x},x_{\bar{q}})\ \rho_{I}^{(0)}\big(\gamma_{Iq}(s\hat{x}/m_{q}^{2}),z\big)\ \rho_{I}^{(2)}\big(-\gamma_{Iq}(Q_{0}^{2}/m_{q}^{2}),r\big)\\ &\times \sigma_{q\bar{q}}^{Born}(s\hat{x}z)\ \langle W_{MC}\rangle, \end{split} \tag{4}$$





• (HL-)LHC:

KKMChh: NISR shows effect of QED contamination in non-QED PDFs is below the errors on the PDFs:

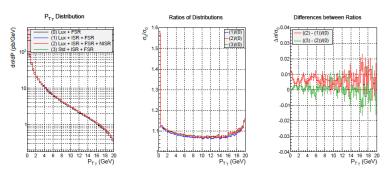


Figure 3(arXiv:2211.17177): The distribution for P_{TY} of the photon for which it is greatest for events with at least one photon and each lepton having $p_{TT} > 25$ GeV, $\eta_t < 2.5$ calculated with (0) FSR only (black), (1) FSR + ISR (blue), and (2) FSR + ISR with NISR (red) for NNPDF3.1-LuxQED NLO PDFs. For comparison, (3) shows FSR + ISR with ordinary NNPDF3.1 NLO PDFs (green). The center graph shows ISR on/off ratios (1)/(0) (blue),(2)/(0) (red) and (3)/(0) (green). The right-hand graph shows the fractional differences ((1) – (2))/(0) in red and ((2) – (3))/(0) in green.



FCC-ee:

BHLUMI and the Luminosity Theory Error – Current Purview

(M. Skrzypek et al., 2023 FCC Workshop, Krakow & MTTD2023; W. Placzek, here)

Bhabha cross sect. depends on detector acceptance angles

$$\sigma_{Bh} \simeq 4\pi lpha^2 \left(rac{1}{t_{ ext{min}}} - rac{1}{t_{ ext{max}}}
ight) = 4\pi lpha^2 \left(rac{t_{ ext{max}} - t_{ ext{min}}}{ar{t}^2}
ight), \quad ar{t} = \sqrt{t_{ ext{min}}t_{ ext{max}}}$$

 \bar{t} is the characteristic scale of the process

 \overline{t}/s is the suppression factor between s- and t-channel contributions

Machine	$\theta_{\min} \div \theta_{\max}$ [mrad]	\sqrt{s} [GeV]	\bar{t}/s	\sqrt{t} [GeV]
LEP	28÷50	M_Z	3.5×10^{-4}	1.70
FCCee	64÷86	M_Z	13.7×10^{-4}	3.37
FCCee	64÷86	240	13.7×10^{-4}	8.9
FCCee	64÷86	350	13.7×10^{-4}	13.0
ILC	31÷77	500	6.0×10^{-4}	12.2
ILC	31÷77	1000	6.0×10^{-4}	24.4
CLIC	39÷134	3000	13.0×10^{-4}	108





FCC-ee:

BHLUMI and the Luminosity Theory Error - Current Purview

Lumi at $FCCee_{M_Z}$ – Forecast study

Forecast study for FCCee _{M7}					
Type of correction / Error	Published [1]	Strict	Redone		
(a) Photonic $\mathcal{O}(L_e^2 \alpha^3)$	0.10×10^{-4}	0.10×10^{-4}	0.10×10^{-4}		
(b) Photonic $\mathcal{O}(L_{\theta}^4 \alpha^4)$	0.06×10^{-4}	0.06×10^{-4}	0.06×10^{-4}		
(b') Photonic $\mathcal{O}(\alpha^2 L^0)$		0.17×10^{-4}	0.17×10^{-4}		
(c) Vacuum polariz.	0.6×10^{-4}	0.6×10^{-4}	0.6×10^{-4}		
(d) Light pairs	0.5×10^{-4}	0.4×10^{-4}	0.27×10^{-4}		
(e) Z and s -channel γ exch.	0.1×10^{-4}	0.1×10^{-4}	0.1×10^{-4}		
(f) Up-down interference	0.1×10^{-4}	0.08×10^{-4}	0.08×10^{-4}		
Total	1.0×10^{-4}	0.76×10^{-4}	0.70×10^{-4}		



• FCC-ee:

BHLUMI and the Luminosity Theory Error – Current Purview

Lumi forecast at ILC and CLIC GeV

Forecast					
Type of correction / Error	ILC ₅₀₀	ILC ₁₀₀₀	CLIC ₃₀₀₀		
(a) Photonic $\mathcal{O}(L_e^2 \alpha^3)$	0.13×10^{-4}	0.15×10^{-4}	0.20×10^{-4}		
(b) Photonic $\mathcal{O}(L_e^4 \alpha^4)$	0.27×10^{-4}	0.37×10^{-4}	0.63×10^{-4}		
(c) Vacuum polariz.	1.1×10^{-4}	1.1×10^{-4}	1.2×10^{-4}		
(d) Light pairs	0.4×10^{-4}	0.5×10^{-4}	0.7×10^{-4}		
(e) Z and s -channel γ exch.	$1.0 \times 10^{-4(*)}$	2.4×10^{-4}	16×10^{-4}		
(f) Up-down interference	$< 0.1 \times 10^{-4}$	$< 0.1 \times 10^{-4}$	0.1×10^{-4}		
Total	1.6×10^{-4}	2.7×10^{-4}	16×10^{-4}		

Note: Lattice methods with Jegerlehner's results allow, in principle, (c) -> (c)/6

$$\Delta\alpha_{had}(t) = \Delta\alpha_{had}(-Q_0^2)|_{lat} + [\Delta\alpha_{had}(t) - \Delta\alpha_{had}(-Q_0^2)]|_{pQCDAdler}$$



Approach to Quantum Gravity

 Cosmological Constant Result Still Obtains: (Phys. Dark Univ. 2 (2013) 97)

$$\begin{split} \rho_{\Lambda}(t_0) &\cong \frac{-M_{pl}^4 (1 + c_{2,eff} k_{tr}^2 / (360\pi M_{pl}^2))^2}{64} \sum_j \frac{(-1)^F n_j}{\rho_j^2} \times \frac{t_{tr}^2}{t_{eq}^2} \times (\frac{t_{eq}^{2/3}}{t_0^{2/3}})^3 \\ &\cong \frac{-M_{pl}^2 (1.0362)^2 (-9.194 \times 10^{-3})}{64} \frac{(25)^2}{t_0^2} \cong (2.4 \times 10^{-3} eV)^4. \end{split}$$

$$t_0 \cong 13.7 \times 10^9 \text{ yrs}$$

 $c_{2,eff} \cong 2.56 \times 10^4$, cosmological index of the ST \Rightarrow Constraints: BHs, etc., in progress.



• Basic Formula for CEEX/EEX realization of the YFS resummation of $e^+e^-\to f\bar f + \eta \ , \ f=\ell,q, \ \ell=e,\mu,\tau,\nu_e,\nu_\mu,\nu_\tau, \ q=u,d,s,c,b,t :$

$$\sigma = \frac{1}{\text{flux}} \sum_{n=0}^{\infty} \int d\text{LIPS}_{n+2} \, \rho_A^{(n)}(\{p\}, \{k\}), \tag{5}$$

•

$$\rho_{\mathsf{CEEX}}^{(n)}(\{p\},\{k\}) = \frac{1}{n!} e^{\mathsf{Y}(\Omega;\{p\})} \bar{\Theta}(\Omega) \frac{1}{4} \sum_{\mathsf{helicities}\ \{\lambda\},\{\mu\}} \left| \mathcal{M} \begin{pmatrix} \{p\}\{k\} \\ \{\lambda\}\{\mu\} \end{pmatrix} \right|^2.$$
(6)

By definition, $\Theta(\Omega, k)=1$ for $k\in\Omega$ and $\Theta(\Omega, k)=0$ for $k\not\in\Omega$, with $\bar{\Theta}(\Omega; k)=1-\Theta(\Omega, k)$ and

$$\bar{\Theta}(\Omega) = \prod_{i=1}^n \bar{\Theta}(\Omega, k_i).$$





• For Ω defined with the condition $k^0 < E_{min}$, the YFS infrared exponent reads

$$Y(\Omega; p_{a},...,p_{d}) = Q_{e}^{2} Y_{\Omega}(p_{a}, p_{b}) + Q_{f}^{2} Y_{\Omega}(p_{c}, p_{d})$$

$$+ Q_{e} Q_{f} Y_{\Omega}(p_{a}, p_{c}) + Q_{e} Q_{f} Y_{\Omega}(p_{b}, p_{d})$$

$$- Q_{e} Q_{f} Y_{\Omega}(p_{a}, p_{d}) - Q_{e} Q_{f} Y_{\Omega}(p_{b}, p_{c}).$$
(7)

Here

$$Y_{\Omega}(p,q) \equiv 2\alpha \tilde{B}(\Omega,p,q) + 2\alpha \Re B(p,q)$$

$$\equiv -2\alpha \frac{1}{8\pi^{2}} \int \frac{d^{3}k}{k^{0}} \Theta(\Omega;k) \left(\frac{p}{kp} - \frac{q}{kq}\right)^{2}$$

$$+2\alpha \Re \int \frac{d^{4}k}{k^{2}} \frac{i}{(2\pi)^{3}} \left(\frac{2p-k}{2kp-k^{2}} - \frac{2q+k}{2kq+k^{2}}\right)^{2}.$$
(8)

• Fundamental Idea of YFS: isolate and resum to all orders in α the infrared singularities so that these singularities are canceled to all such orders between real and virtual corrections.

What collinear singularities are also resummed in the YFS resummation algebra?



 Focusing on the s-channel and s'-channel contributions, we have

$$Y_{e}(\Omega_{I}; p_{1}, p_{2}) = \gamma_{e} \ln \frac{2E_{min}}{\sqrt{2p_{1}p_{2}}} + \frac{1}{4}\gamma_{e} + Q_{e}^{2} \frac{\alpha}{\pi} \left(-\frac{1}{2} + \frac{\pi^{2}}{3} \right),$$

$$Y_{f}(\Omega_{F}; q_{1}, q_{2}) = \gamma_{f} \ln \frac{2E_{min}}{\sqrt{2q_{1}q_{2}}} + \frac{1}{4}\gamma_{f} + Q_{f}^{2} \frac{\alpha}{\pi} \left(-\frac{1}{2} + \frac{\pi^{2}}{3} \right),$$
(9)

where

$$\gamma_{e} = 2Q_{e}^{2} \frac{\alpha}{\pi} \left(\ln \frac{2p_{1}p_{2}}{m_{e}^{2}} - 1 \right), \quad \gamma_{f} = 2Q_{f}^{2} \frac{\alpha}{\pi} \left(\ln \frac{2q_{1}q_{2}}{m_{f}^{2}} - 1 \right), \tag{10}$$

- \Rightarrow The YFS exponent resums the collinear big log term $\frac{1}{2}Q^2\frac{\alpha}{\pi}L$ to the infinite order in both the ISR and FSR contributions.
- Can this be improved to the result of Gribov and Lipatov to exponentiate $\frac{3}{2}\frac{\alpha}{\pi}L$ via the QED form-factor?



The YFS form factor derivation illustrated in Fig. 4

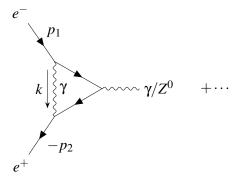


Figure: Virtual corrections which generate the YFS infrared function *B*. Self-energy contributions are not shown.



the amplitude factor

$$\mathcal{M}_{\mu} = \frac{\int d^{4}k}{(2\pi)^{4}} \frac{-i}{k^{2} + i\epsilon} \bar{v}(\rho_{2})(-iQ_{e}e)\gamma^{\alpha} \frac{i}{-\dot{\rho}_{2} - \not{k} - m + i\epsilon} (-ie)\gamma_{\mu}(v_{A} - a_{A}\gamma_{5})$$

$$\frac{i}{\dot{\rho}_{1} - \not{k} - m + i\epsilon} (-iQ_{e}e)\gamma_{\alpha}u(\rho_{1})$$
(11)

where $A = \gamma$ or Z.





ullet Scalarising the fermion propagator denominators \Rightarrow

$$\mathcal{M}_{\mu} = -ie \frac{\int d^{4}k(-iQ_{e}^{2}e^{2})}{(2\pi)^{4}} \frac{1}{k^{2}+i\epsilon} \bar{v}(p_{2}) \gamma^{\alpha} \frac{-p_{2}-k+m}{k^{2}+2kp_{2}+i\epsilon} \gamma_{\mu}(v_{A} - a_{A}\gamma_{5})$$

$$\frac{p_{1}-k-m}{k^{2}-2kp_{1}+i\epsilon} \gamma_{\alpha} u(p_{1}).$$
(12)

Using the equations of motion

$$(p_1 - k - m)\gamma_{\alpha}u(p_1) = \{(2p_1 - k)_{\alpha} - \frac{1}{2}[k, \gamma_{\alpha}]\}u(p_1),$$
 (a)

$$\bar{v}(\rho_2)\gamma^{\alpha}(-\rho_2-k+m) = \bar{v}(\rho_2)\{-(2\rho_2+k)^{\alpha} + \frac{1}{2}[k,\gamma^{\alpha}]\}, \qquad (b).$$
(13)





• \Rightarrow Contribution to $2Q_e^2\alpha B(p_1,p_2)$ corresponding to the cross-term in the virtual IR function on the RHS of eq.(8):

$$\begin{split} 2Q_{e}^{2}\alpha B(p_{1},p_{2})|_{\text{cross-term}} &= \int d^{4}k \frac{(iQ_{e}^{2}e^{2})}{8\pi^{4}} \frac{1}{k^{2}+i\epsilon} \\ &\frac{(2p_{1}-k)(2p_{2}+k)}{(k^{2}-2kp_{1}+i\epsilon)(k^{2}+2kp_{2}+i\epsilon)}. \end{split} \tag{14}$$

This term, together with the two squared terms in $2\alpha Q_e^2 B(p_1,p_2)$, leads to the exponentiation of $\frac{1}{2}Q_e^2\frac{\alpha}{\pi}L$.



- The two commutator terms on the RHS of eq.(13), usually dropped, can be analyzed further: possible IR finite collinearly enhanced improvement of the YFS virtual IR function B.
- Isolate the collinear parts of k via the change of variables

$$k = c_1 p_1 + c_2 p_2 + k_{\perp} \tag{15}$$

where $p_1 k_{\perp} = 0 = p_2 k_{\perp}$, \Rightarrow we have the relations

$$c_{1} = \frac{p_{1}p_{2}}{(p_{1}p_{2})^{2} - m^{4}} p_{2}k - \frac{m^{2}}{(p_{1}p_{2})^{2} - m^{4}} p_{1}k \xrightarrow{c_{L}} \frac{p_{2}k}{p_{1}p_{2}}$$

$$c_{2} = \frac{p_{1}p_{2}}{(p_{1}p_{2})^{2} - m^{4}} p_{1}k - \frac{m^{2}}{(p_{1}p_{2})^{2} - m^{4}} p_{2}k \xrightarrow{c_{L}} \frac{p_{1}k}{p_{1}p_{2}},$$

$$(16)$$

CL denotes the collinear limit $\equiv O(m^2/s)$ dropped.





• \Rightarrow $(2p_1 - k)^{\alpha}$ in eq.(13(a)) combines with the commutator term in eq.(13(b)) to produce

$$\bar{v}(p_{2})\{(2p_{1}-k)_{\alpha}\frac{1}{2}[k,\gamma^{\alpha}]\}\gamma_{\mu}(v_{A}-a_{A}\gamma_{5})u(p_{1})
= \bar{v}(p_{2})[k,p_{1}]\gamma_{\mu}(v_{A}-a_{A}\gamma_{5})u(p_{1})
\xrightarrow{CL} \bar{v}(p_{2})[c_{2}p_{2},p_{1}]\gamma_{\mu}(v_{A}-a_{A}\gamma_{5})u(p_{1})
\xrightarrow{CL} \bar{v}(p_{2})(-2c_{2}p_{1}p_{2})\gamma_{\mu}(v_{A}-a_{A}\gamma_{5})u(p_{1})
\xrightarrow{CL} \bar{v}(p_{2})(-2p_{1}k)\gamma_{\mu}(v_{A}-a_{A}\gamma_{5})u(p_{1}).$$
(17)

• Similarly, $-(2p_2 + k)^{\alpha}$ in eq.(13 (b)) combines with the commutator term in eq.(13(a)) to produce

$$\begin{split} \bar{v}(\rho_{2})\gamma_{\mu}(v_{A} - a_{A}\gamma_{5})\{-(2p_{2} + k)^{\alpha}(-\frac{1}{2}[k,\gamma_{\alpha}])\}u(p_{1}) \\ &= \bar{v}(p_{2})\gamma_{\mu}(v_{A} - a_{A}\gamma_{5})[k,\not p_{2}]u(p_{1}) \\ &\xrightarrow{CL} \bar{v}(p_{2})\gamma_{\mu}(v_{A} - a_{A}\gamma_{5})[c_{1}\not p_{1},\not p_{2}]u(p_{1}) \\ &\xrightarrow{CL} \bar{v}(p_{2})\gamma_{\mu}(v_{A} - a_{A}\gamma_{5})(2c_{1}p_{1}p_{2})u(p_{1}) \\ &\xrightarrow{CL} \bar{v}(p_{2})\gamma_{\mu}(v_{A} - a_{A}\gamma_{5})(2p_{2}k)u(p_{1}). \end{split}$$



• \Rightarrow Shift of the factor $(2p_1 - k)(2p_2 + k)$ on the RHS of eq.(14) as

$$(2p_1-k)(2p_2+k) \xrightarrow{CL} (2p_1-k)(2p_2+k)+2p_1k-2p_2k.$$
 (19)



- What does the term quadratic in the commutator (C²) contribute?
- Superficial UV divergence ⇒ Cannot naively drop k_⊥
- Proceed directly: we need

$$\begin{split} 2Q_{e}^{2}\alpha\mathcal{B}(p_{1},p_{2})|_{C^{2}}\mathcal{M}_{\mathcal{B}\mu} &\equiv \frac{\int d^{4}k(iQ_{e}^{2}e^{2})}{8\pi^{4}} \frac{1}{k^{2}+i\epsilon} \\ &\frac{\frac{1}{4}\bar{v}(p_{2})[\cancel{k},\gamma^{\alpha}]\gamma_{\mu}[\cancel{k},\gamma_{\alpha}](-ie)(v_{A}-a_{A}\gamma_{5})u(p_{1})}{(k^{2}-2kp_{1}+i\epsilon)(k^{2}+2kp_{2}+i\epsilon)} \bigg|_{CL}, \end{split}$$

where we define

$$\mathcal{M}_{B\mu} = -ie\bar{v}(p_2)\gamma_{\mu}(v_A - a_A\gamma_5)u(p_1). \tag{21}$$

• *CL* now further restricted to contributions singular as $m^2/s \rightarrow 0$.





 Four terms in the numerator of eq.(20) from the respective sum of gamma matrix products

$$\{ k \gamma^{\alpha} \gamma_{\mu} k \gamma_{\alpha} - k \gamma^{\alpha} \gamma_{\mu} \gamma_{\alpha} k - \gamma^{\alpha} k \gamma_{\mu} k \gamma_{\alpha} + \gamma^{\alpha} k \gamma_{\mu} \gamma_{\alpha} k \} =$$

$$\{ \gamma^{\lambda} \gamma^{\alpha} \gamma_{\mu} \gamma^{\lambda'} \gamma_{\alpha} - \gamma^{\lambda} \gamma^{\alpha} \gamma_{\mu} \gamma_{\alpha} \gamma^{\lambda'} - \gamma^{\alpha} \gamma^{\lambda} \gamma_{\mu} \gamma^{\lambda'} \gamma_{\alpha} + \gamma^{\alpha} \gamma^{\lambda} \gamma_{\mu} \gamma_{\alpha} \gamma^{\lambda'} \} k_{\lambda} k_{\lambda'} \equiv$$

$$N_{\mu}^{\lambda \lambda'} k_{\lambda} k_{\lambda'}$$

• This defines $N_{\mu}^{\lambda\lambda'}$.



Using standard methods, we need

$$I_{\mu} = 2 \int_{0}^{1} d\alpha_{1} \int_{0}^{1-\alpha_{1}} d\alpha_{2} \frac{\int d^{n}k'(iQ_{e}^{2}e^{2})}{8\pi^{4}} \frac{\frac{1}{4}\bar{v}(p_{2})N_{\mu}^{\lambda\lambda'}\left[\frac{k'^{2}}{n}g_{\lambda\lambda'} + \Delta_{\lambda}\Delta_{\lambda'}\right](-ie)(v_{A} - a_{A}\gamma_{5})u(p_{1})}{[k'^{2} - \Delta^{2} + i\epsilon]^{3}}\Big|_{CL},$$
(22)

where $\Delta = \alpha_1 p_1 - \alpha_2 p_2$.

ullet Equations of motion \Rightarrow term involving Δ is not collinearly enhanced.



• The term contracted with $g_{\lambda\lambda'}$ gives us

$$I_{\mu} = \left\{ \frac{-3Q_{e}^{2}\alpha}{4\pi} \mathcal{M}_{B\mu} \right\} \bigg|_{CL} \equiv 0$$
 (23)

- \Rightarrow No collinearly enhanced contribution from I_{μ} .
- Eq.(19) gives the complete collinear enhancement of B.
- Change in B does not affect its IR behavior shift terms are IR finite ⇒ Entire YFS IR resummation is unaffected.
- Shifted terms can be seen to extend the YFS IR exponentiation to obtain the entire exponentiated $\frac{3}{2}Q_e^2\alpha L$.



We have

$$2\alpha Q_{e}^{2}\Delta B(p_{1},p_{2}) = \frac{\int d^{4}k(iQ_{e}^{2}e^{2})}{8\pi^{4}} \frac{1}{k^{2} + i\epsilon} \frac{2p_{1}k - 2p_{2}k}{(k^{2} - 2kp_{1} + i\epsilon)(k^{2} + 2kp_{2} + i\epsilon)}$$

$$= 2\int_{x_{i} \geq 0, i=1,2,3} d^{3}x\delta(1 - x_{1} - x_{2} - x_{3}) \frac{\int d^{4}k'(iQ_{e}^{2}e^{2})}{8\pi^{4}}$$

$$\frac{2(p_{1} - p_{2})p_{x}}{(k'^{2} - d + i\epsilon)^{3}}$$
(24)

where $d = p_x^2$ with $p_y = x_1 p_1 - x_2 p_2$.

⇒ We get

$$2Q_{\theta}^{2}\alpha\Re\Delta B(p_{1},p_{2}) = Q_{\theta}^{2}\frac{\alpha}{\pi}L.$$
 (25)

• We see that indeed the entire term $\frac{3}{2}Q_{e}^{2}\frac{\alpha}{\pi}L$ is now exponentiated by our collinearly improved YFS virtual IR function Bc

$$B_{CL} = B + \Delta B$$

$$= \int \frac{d^4k}{k^2} \frac{i}{(2\pi)^3} \left[\left(\frac{2p - k}{2kp - k^2} - \frac{2q + k}{2kq + k^2} \right)^2 - \frac{4pk - 4qk}{(2pk - k^2)(2qk + k^2)} \right].$$
(26)

BAYLOR

See S. Jadach, Durham talk, 2002, for integrated form of B_{CL}



- What about the real YFS IR algebra? Collinear enhancement desired in some applications
- \Rightarrow Recall the original YFS EEX formulation of the respective algebra \Rightarrow the formula for the YFS IR function \tilde{B} given above in eq.(8).
- See Fig. 5.

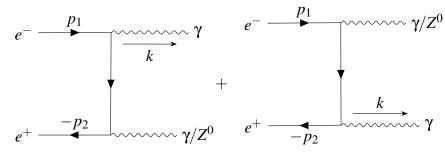


Figure: Real corrections which generate the YFS infrared function \tilde{B} .

Figure 2: Real corrections which generate the YFS infra BAYLOR

lacktriangle Following the steps in the usual YFS algebra for real emission \Rightarrow

$$\begin{split} 2\alpha Q_{e}^{2}\tilde{B}\mathcal{M}_{B\mu}^{\dagger}\mathcal{M}_{B\mu'} &= \frac{\int d^{3}k(-1)e^{2}Q_{e}^{2}}{2k_{0}(2\pi)^{3}} \bigg[\frac{\bar{u}(p_{1})(2p_{1}^{\lambda}-k^{\lambda}+\frac{1}{2}[k,\gamma^{\lambda}])\gamma_{\mu}(\nu_{A}-a_{A}\gamma_{5})\nu(p_{2})}{k^{2}-2kp_{1}} \\ &+ \frac{\bar{u}(p_{1})\gamma_{\mu}(\nu_{A}-a_{A}\gamma_{5})(-2p_{2}^{\lambda}+k^{\lambda}+\frac{1}{2}[k,\gamma^{\lambda}])\nu(p_{2})}{k^{2}-2kp_{2}} \bigg] \\ &\left[\frac{\bar{v}(p_{2})\gamma_{\mu'}(\nu_{A}-a_{A}\gamma_{5})(2p_{1\lambda}-k_{\lambda}-\frac{1}{2}[k,\gamma_{\lambda}])u(p_{1})}{k^{2}-2kp_{1}} \\ &+ \frac{\bar{v}(p_{2})(-2p_{2\lambda}+k_{\lambda}-\frac{1}{2}[k,\gamma_{\lambda}])\gamma_{\mu'}(\nu_{A}-a_{A}\gamma_{5})u(p_{1})}{k^{2}-2kp_{2}} \bigg] \bigg|_{k^{2}=0} + K_{\mu\mu'} \\ &\left. (27) \end{split}$$

where $K_{\mu\mu'}$ is infrared finite,

•

$$\mathcal{M}_{B\mu} = \bar{v}(p_2)\gamma_{\mu}(v_A - a_A\gamma_5)u(p_1) \tag{28}$$





- If we drop the commutator terms on the RHS of eq.(27) we recover the usual YFS formula for $2\alpha Q_{\alpha}^{2}\tilde{B}$.
- We again isolate collinearly enhanced contributions by using the representation in eq.(yfsalg4) for k, respecting the condition $k^2 = 0$. \Rightarrow Maintain $0 = (c_1^2 + c_2^2)m^2 + 2c_1c_2p_1p_2 - |k_{\perp}|^2.$
- $\bullet \Rightarrow$ Collinear enhancement of \tilde{B} :

$$2\alpha Q_e^2 \tilde{B}_{CL} = \frac{-\alpha Q_e^2}{4\pi^2} \int \frac{d^3k}{k_0} \left\{ \left(\frac{p_1}{kp_1} - \frac{p_2}{kp_2} \right)^2 + \frac{1}{kp_1} \left(2 - \frac{kp_2}{p_1 p_2} \right) + \frac{1}{kp_2} \left(2 - \frac{kp_1}{p_1 p_2} \right) \right\}.$$
(29)

Agreement with Berends et al.



- What about CEEX?
- In Fig. 5, use of amplitude-level isolation of real IR divergences, K-S photon polarization vectors ⇒

$$\mathcal{M}_{\mu} = \mathcal{M}_{B\mu} \mathfrak{s}_{CL,\sigma}(k), \tag{30}$$

with

$$\begin{split} \mathfrak{s}_{CL,\sigma}(k) &= \sqrt{2} Q_{e} e \left[-\sqrt{\frac{p_{1}\zeta}{k\zeta}} \frac{\langle k\sigma | \hat{p}_{1} - \sigma \rangle}{2p_{1}k} + \delta_{\lambda-\sigma} \sqrt{\frac{k\zeta}{p_{1}\zeta}} \frac{\langle k\sigma | \hat{p}_{1}\lambda \rangle}{2p_{1}k} \right. \\ &+ \sqrt{\frac{p_{2}\zeta}{k\zeta}} \frac{\langle k\sigma | \hat{p}_{2} - \sigma \rangle}{2p_{2}k} + \delta_{\lambda\sigma} \sqrt{\frac{k\zeta}{p_{2}\zeta}} \frac{\langle \hat{p}_{2}\lambda | k - \sigma \rangle}{2p_{2}k} \right]. \end{split} \tag{31}$$

Here, $\zeta \equiv (1, 1, 0, 0)$ and $\hat{p} = p - \zeta m^2/(2\zeta p)$.

• Upon taking the modulus squared of $\mathfrak{s}_{CL,\sigma}(k)$ we see that the extra non-IR divergent contributions reproduce the known collinear big log contribution which is missed by the usual YFS algebra.

Outlook

- Amplitude-based resummation (Staszek's contributions thereto were essential) allows improved control of IR and Collinear limits
- MC realizations are needed for current and future precision collider physics, using residuals made amenable by computer algebraic methods.
- Enhanced toolbox available to extend the (CEEX) YFS MC method to the other important processes at present and future colliders.
- Some New Physics may hang in the balance at both LHC, FCC, and other future colliders.
- We wish Staszek were still here to share the possible excitement with us, we really do miss him very much.

