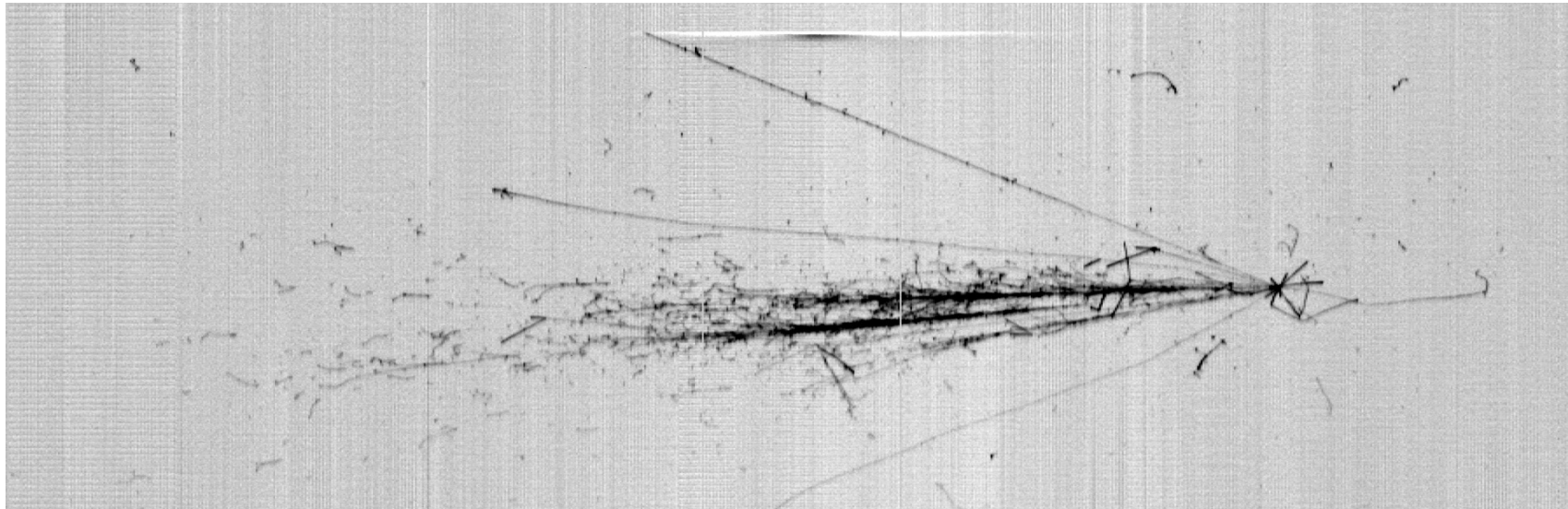


# Search for sterile neutrinos including recent results from the ICARUS T600

KRZYSZTOF CIESLIK  
for the ICARUS Collaboration

*The Henryk Niewodniczanski Institute of Nuclear Physics,  
Polish Academy of Sciences, Krakow*

***Matter To The Deepest 2013, 03.09.2013 r.***



# Outline

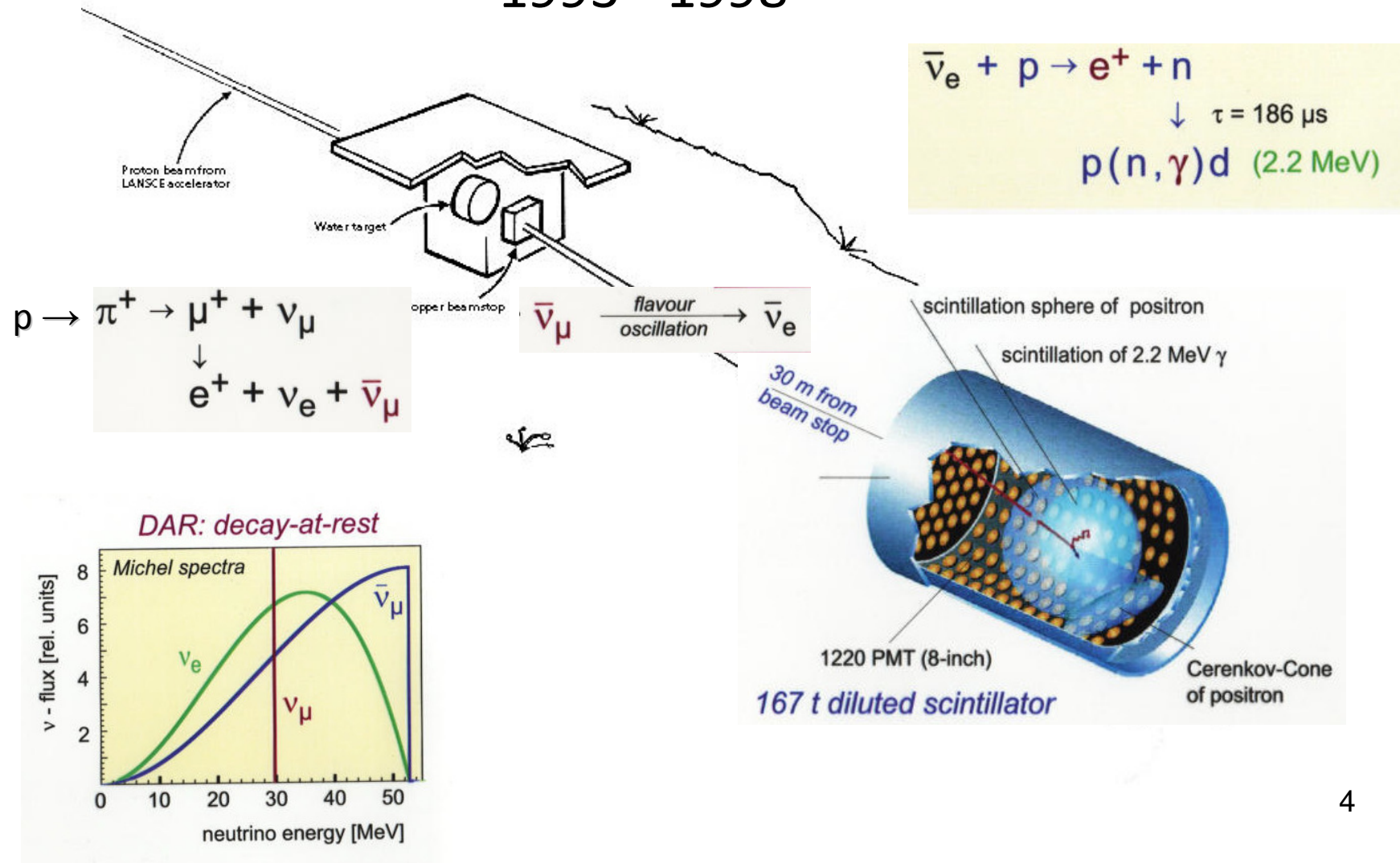
- Sterile neutrinos (LSND/MiniBooNE anomaly)
- The ICARUS experiment
- Search for sterile neutrinos in the CNGS beam
  1. *Experimental search for the „LSND anomaly” with the ICARUS detector in the CNGS beam, Eur. Phys. J. C 73(2013) (arXiv:1209.0122 Sep 2012)*
  2. *New improved results: (arXiv:1307.4699 Jul 2013)*
- Conclusions

# Sterile neutrinos

- Sterile neutrinos were hypothesized in 1957 by B.Pontecorvo as particles not interacting via any of fundamental interactions except gravity.
- They are extremely difficult to detect. If they are heavy enough, they may also contribute to the dark matter.
- Sterile neutrinos may mix with standard neutrinos via a mass term. The „LSND anomaly” and results from the MiniBooNE experiment may be considered as the experimental hints for sterile neutrinos.

# LSND (Liquid Scintillator Neutrino Detector)

1993 - 1998



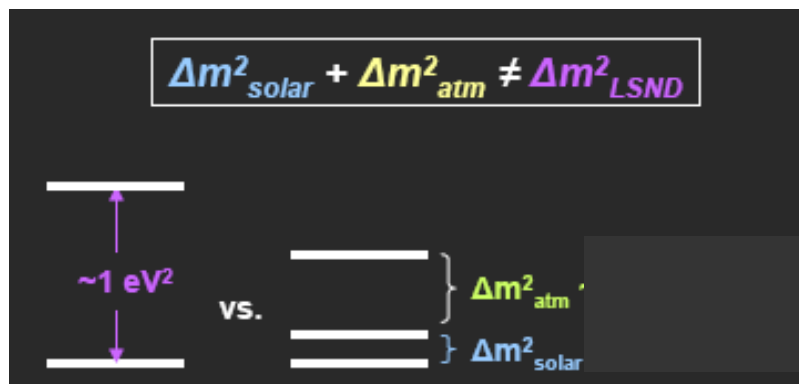
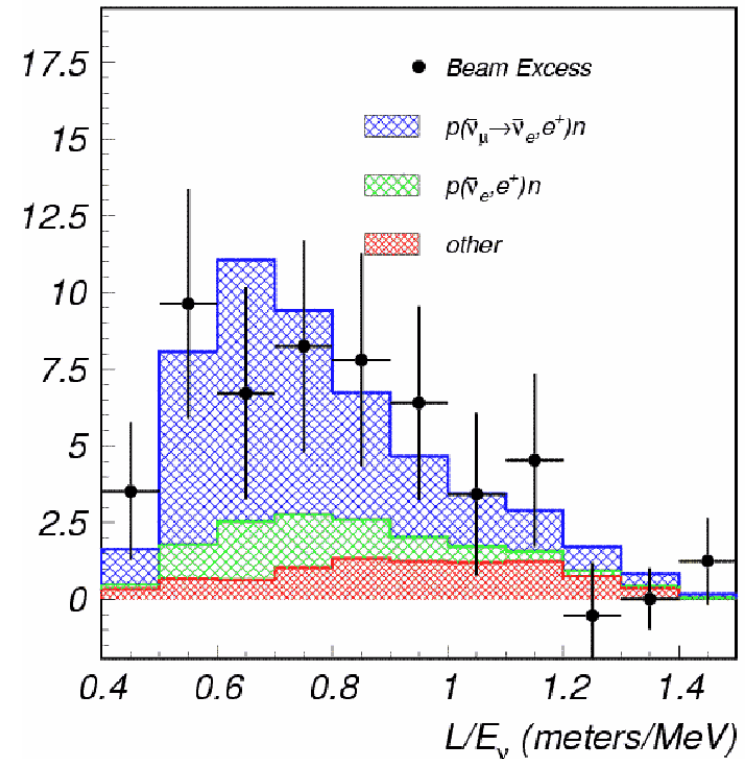
# LSND anomaly

LSND found an excess of  $\bar{\nu}_e$  in  $\bar{\nu}_\mu$  beam

Excess:  $87.9 \pm 22.4 \pm 6.0$  ( $3.8\sigma$ )

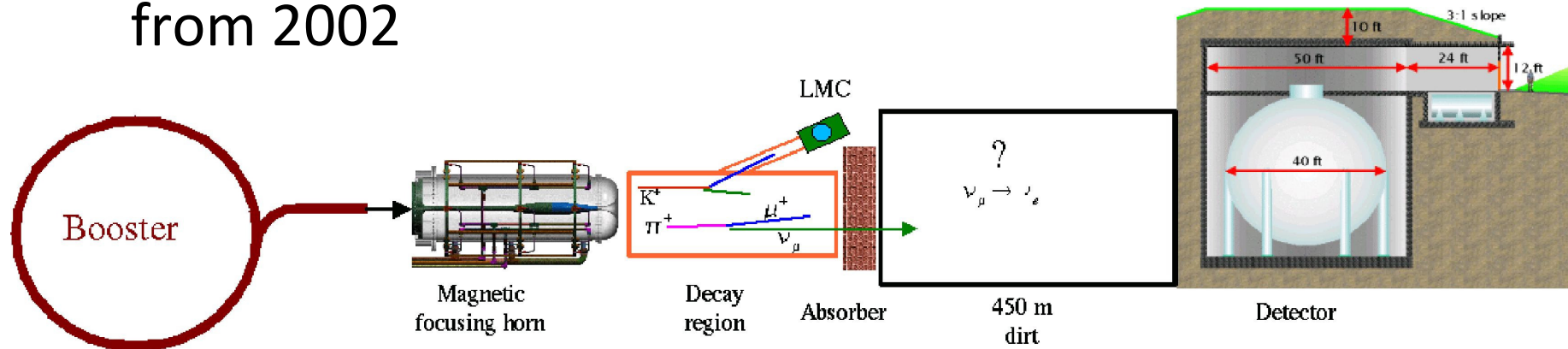
Oscillation probability:  $0.264 \pm 0.067 \pm 0.045$

$\Delta m^2$  :  $0.2 - 10 \text{ eV}^2$



# MiniBooNE

from 2002



- Test the LSND anomaly
- Keep L/E same, change beam and energy
- 8 GeV proton beam (Be target)

**neutrino energy (E):**

MiniBooNE:  $\sim 700$  MeV

LSND:  $\sim 30$  MeV

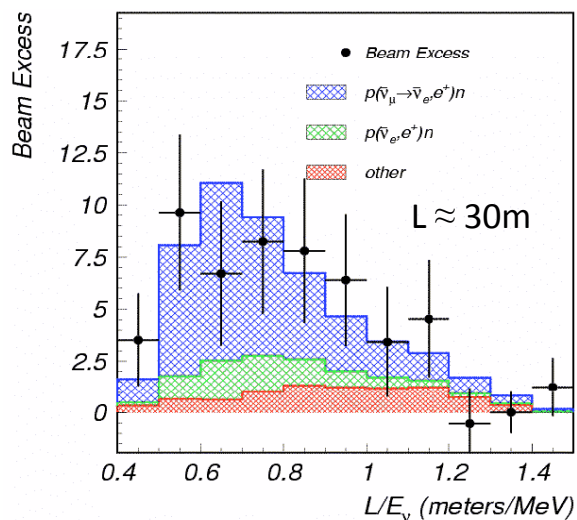
**baseline (L):**

MiniBooNE:  $\sim 540$  m

LSND:  $\sim 30$  m

- Mineral Oil Cherenkov Detector
- 800 tons, 12 m diameter sphere
- 1280 eight-inch PMT's
- 240 PMT for VETO
- 611,000  $\nu$  events

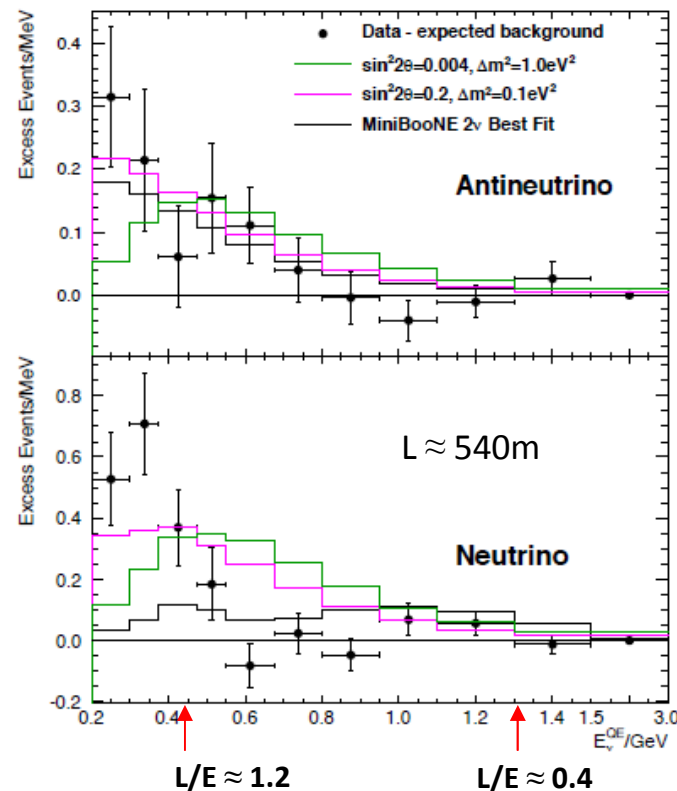
# LSND and MiniBooNE



LSND has observed an excess of  $\bar{\nu}_e$  events in  $\bar{\nu}_\mu$  beam,  
 $87.9 \pm 22.4 \pm 6.0$  ( $3.8 \sigma$ )

Experiments showing negative evidence:

KARMEN, NOMAD, BUGEY, NUTEV



for  $200 < E_{QE} < 1250$  MeV

antineutrino:  $78.4 \pm 28.5$  ( $2.8 \sigma$ )

neutrino:  $162 \pm 47.8$  ( $3.4 \sigma$ )

for neutrinos the energy distribution is marginally compatible with a two neutrino oscillation formalism

MiniBoone results do not fully confirm the „LSND anomaly”

## Positive hints

| <i>Anomaly</i>   | <i>Source</i>    | <i>Type</i>      | <i>Channel</i>                         | <i>Significance</i>              |
|--|------------------|------------------|--|----------------------------------|
| <b>LSND</b>  | Short baseline   | Decay at rest    | $-\nu\mu \rightarrow \nu e$<br>CC      | 3.8 $\sigma$                     |
| <b>MiniBoone</b>   | Short baseline   | Neutrino beam    | $-\nu\mu \rightarrow \nu e$<br>CC      | <b>3.4 <math>\sigma</math></b>   |
| <b>MiniBoone</b>   | Short baseline   | Anti-Neutr. beam | anti- $\nu\mu \rightarrow \nu e$<br>CC | <b>2.8 <math>\sigma</math></b>   |
| <b>Gallium</b>   | Electron capture | Source           | $\nu$ disapp.                          | 2.7 $\sigma$                     |
| <b>Reactors</b>  | Fission          | Beta decay       | $\nu$ disapp.                          | 3.0 $\sigma$                     |
| Zhang, Qian, Vogel: „ <i>Reactor antineutrino with known <math>\theta_{13}</math></i> ”<br>(arXiv:1303.0900), Mar 2013 |                  |                  |  | <b>→ 1.4 <math>\sigma</math></b> |



# The ICARUS Collaboration

M. Antonello<sup>a</sup>, P. Aprili<sup>a</sup>, B. Baibussinov<sup>b</sup>, M. Baldo Ceolin<sup>b</sup>, P. Benetti<sup>c</sup>, E. Calligarich<sup>c</sup>, N. Canci<sup>a</sup>, S. Centro<sup>b</sup>, A. Cesana<sup>f</sup>, K. Cieslik<sup>g</sup>, D. B. Cline<sup>h</sup>, A.G. Cocco<sup>d</sup>, A. Dabrowska<sup>g</sup>, D. Dequal<sup>b</sup>, A. Dermenev<sup>i</sup>, R. Dolfini<sup>c</sup>, C. Farnese<sup>b</sup>, A. Fava<sup>b</sup>, A. Ferrari<sup>j</sup>, G. Fiorillo<sup>d</sup>, D. Gibin<sup>b</sup>, A. Gigli Berzolari<sup>c</sup>, S. Gninenko<sup>i</sup>, A. Guglielmi<sup>b</sup>, M. Haranczyk<sup>g</sup>, J. Holeczek<sup>l</sup>, A. Ivashkin<sup>i</sup>, J. Kisiel<sup>l</sup>, I. Kochanek<sup>l</sup>, J. Lagoda<sup>m</sup>, S. Mania<sup>l</sup>, G. Mannocchi<sup>n</sup>, A. Menegolli<sup>c</sup>, G. Meng<sup>b</sup>, C. Montanari<sup>c</sup>, S. Otwinowski<sup>h</sup>, L. Periale<sup>n</sup>, A. Piazzoli<sup>c</sup>, P. Picchi<sup>n</sup>, F. Pietropaolo<sup>b</sup>, P. Plonski<sup>o</sup>, A. Rappoldi<sup>c</sup>, G.L. Raselli<sup>c</sup>, M. Rossella<sup>c</sup>, C. Rubbia<sup>a,j</sup>, P. Sala<sup>f</sup>, E. Scantamburlo<sup>e</sup>, A. Scaramelli<sup>f</sup>, E. Segreto<sup>a</sup>, F. Sergiampietri<sup>p</sup>, D. Stefan<sup>a</sup>, J. Stepaniak<sup>m</sup>, R. Sulej<sup>m,a</sup>, M. Szarska<sup>g</sup>, M. Terrani<sup>f</sup>, F. Varanini<sup>b</sup>, S. Ventura<sup>b</sup>, C. Vignoli<sup>a</sup>, H. Wang<sup>h</sup>, X. Yang<sup>h</sup>, A. Zalewska<sup>g</sup>, K. Zaremba<sup>o</sup>

- a *Laboratori Nazionali del Gran Sasso dell'INFN, Assergi, Italy*
- b *Dipartimento di Fisica e INFN, Università di Padova, Italy*
- c *Dipartimento di Fisica Nucleare e Teorica e INFN, Università di Pavia, Italy*
- d *Dipartimento di Scienze Fisiche, INFN e Università Federico II, Napoli, Italy*
- e *Dipartimento di Fisica, University of L'Aquila, Italy*
- f *INFN, Sezione di Milano e Politecnico, Milano, Italy*
- g *Institute of Nuclear Physics, Polish Academy of Science, Krakow, Poland*
- h *Department of Physics and Astronomy, University of California, Los Angeles, USA*
- i *INR RAS, prospekt 60-letiya Oktyabrya 7a, Moscow, Russia*
- j *CERN, Ch1211 Geneve 23, Geneva, Switzerland*
- l *Institute of Physics, University of Silesia, 12 Bankowa st., 40-007 Katowice, Poland*
- m *A. Soltan Institute for Nuclear Studies, 05-400 Swierk/Otwock, Warszawa, Poland*
- n *Laboratori Nazionali di Frascati (INFN), Frascati, Italy*
- o *Institute for Radioelectronics, Warsaw Univ. of Technology, Warsaw, Poland*
- p *Dipartimento di Fisica, Università di Pisa, Italy*

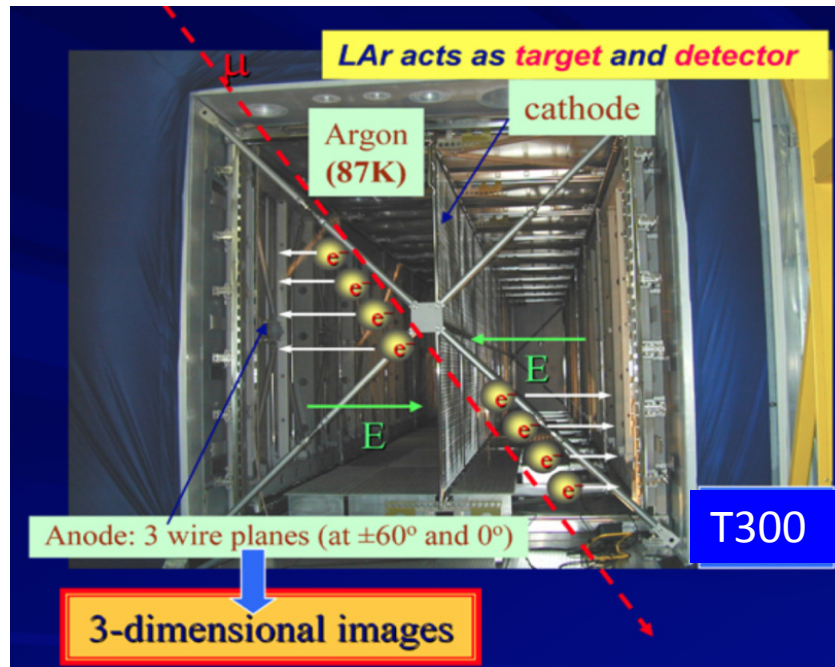


# The ICARUS detector (T600)

The **Liquid Argon Time Projection Chamber** (*electronic bubble chamber*)

[C. Rubbia: CERN-EP/77-08 (1977)] capable of providing a 3D imaging of any charged particle with:

- high granularity (spatial resolution of the detector  $\sim 1 \text{ mm}^3$ );
- excellent calorimetric properties.

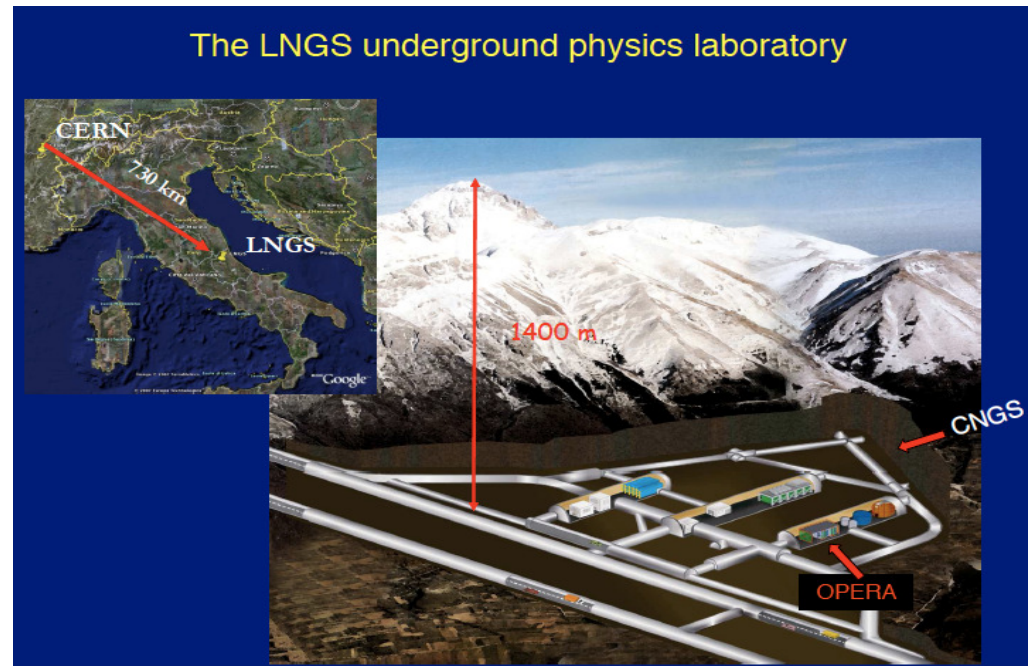
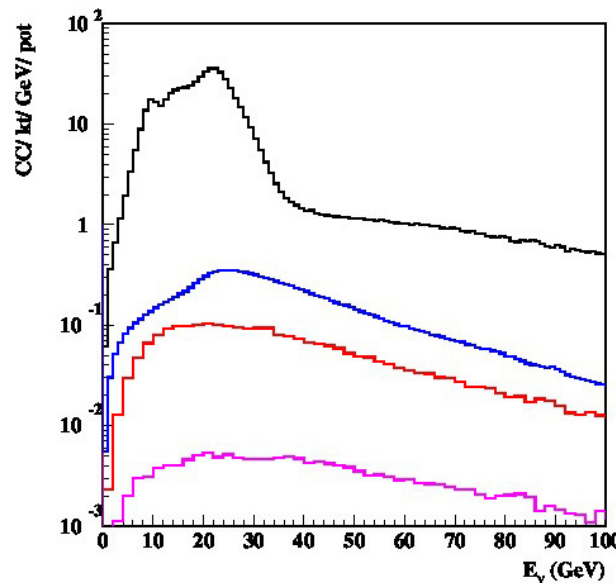


- ionization and scintillation signals are exploited,
- PMTs signal is used for triggering,
- continuous probing of wire signals as a function of time allows 3D reconstruction.
- very pure argon :  $\tau_{\text{ele}} > 5 \text{ ms}$

- Total LAr mass 600 t, active mass 476 t
- Two identical T300 modules (2 TPC chambers for each module).
- TPC characteristics:
  - (17.9 x 3.1 x 1.5 for each TPC) m<sup>3</sup>;
  - drift length = 1.5 m;
  - $E_{\text{drift}} = 0.5 \text{ kV/cm}$ ;  $v_{\text{drift}} = 1.6 \text{ mm}/\mu\text{s}$ .
- 3 readout wire planes/chamber at  $0^\circ, +60^\circ, -60^\circ$ , 3 mm plane and wire spacing:
  - $\sim 53000$  wires;
  - two induction planes and one collection
- PMTs for scintillation light (128 nm):
  - (20+54) PMTs.

# CNGS – CERN Neutrinos to Gran Sasso

Conventional beam based on protons from the SPS accelerator at CERN

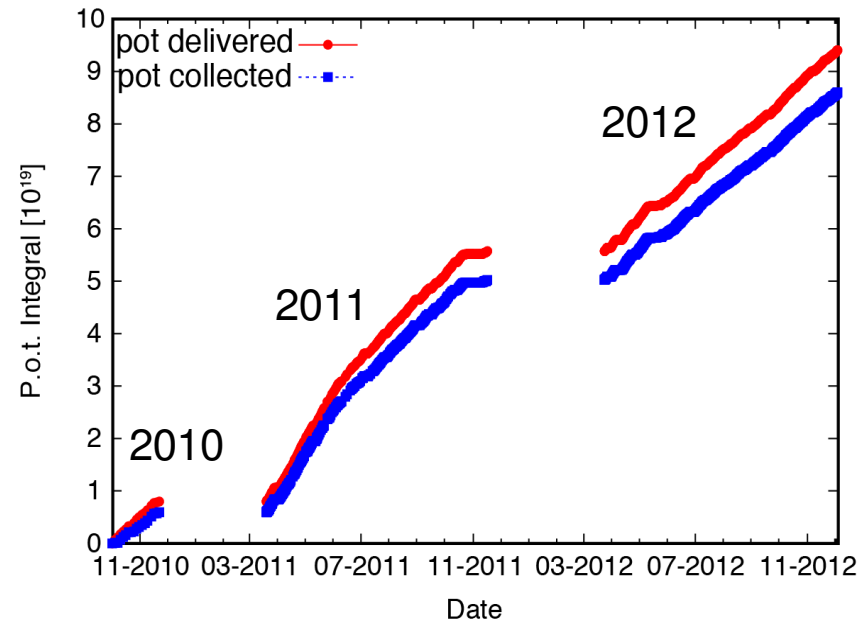


$E \rightarrow 10 - 30 \text{ GeV}, \quad L \approx 730 \text{ km}$

$\nu_e/\nu_\mu \sim 0.8\%, \quad \bar{\nu}_\mu/\nu_\mu \sim 2.1\%, \quad \bar{\nu}_e/\nu_\mu \sim 0.07\%$

# CNGS data taking

- CNGS data useful for analysis (01.10.2010 – 03.12.2012)
- Technical run with cosmics (Dec. 2012 – Jun. 2013)
- Trigger based on PMT signals, in coincidence with proton extraction
- detector live-time > 93%
- total  $8.6 \times 10^{19}$  pot collected



# LAr T600 reconstruction performance

## Tracking:

- The tracking is done with high sampling ( $\sim 3\text{mm}$ )
- Muon momentum via multiple scattering

## Total energy reconstruction from charge integration

### ENERGY RESOLUTIONS:

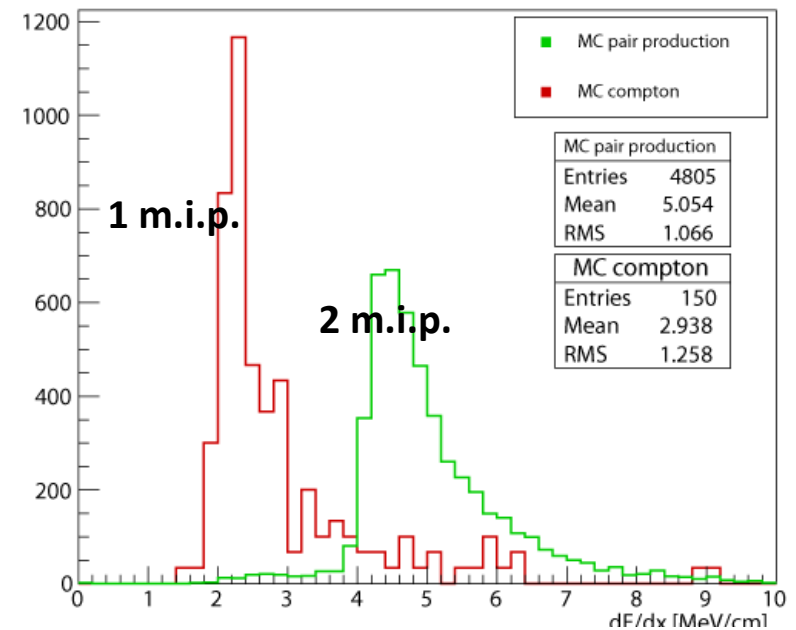
Low energy electrons  $\sigma(E)/E = 11\%/\sqrt{E(\text{MeV})} + 2\%$

Electromagnetic showers  $\sigma(E)/E = 3\%/\sqrt{E(\text{GeV})}$

Hadron shower (pure LAr)  $\sigma(E)/E \approx 30\%/\sqrt{E(\text{GeV})}$

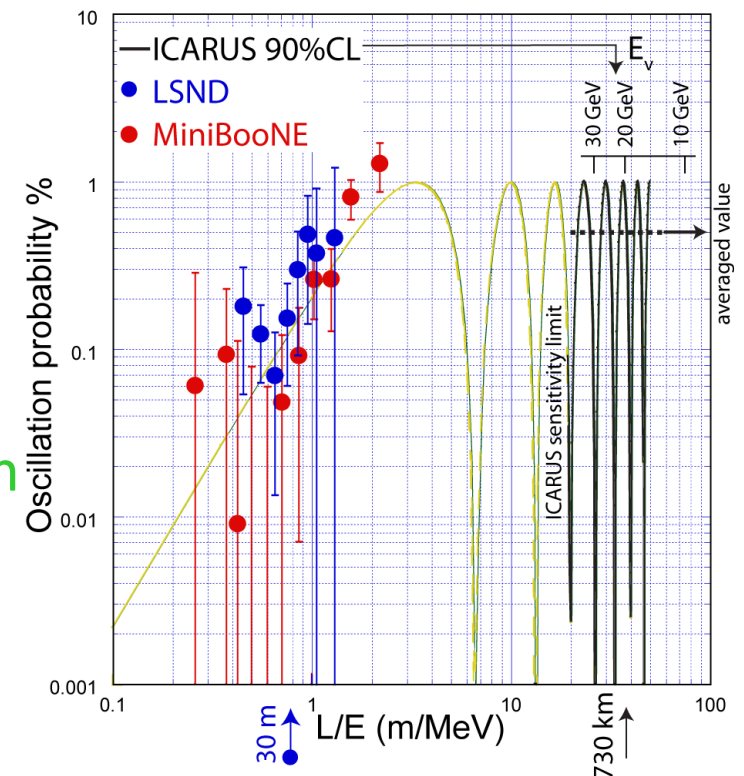
## Measurement of local energy deposition $dE/dx$ :

- Very good  $e/\pi^0$  separation by means of  $dE/dx$  in the first part of the cascade
- Particle identification by  $dE/dx$  vs range



# A search for the LSND effect in the ICARUS T600

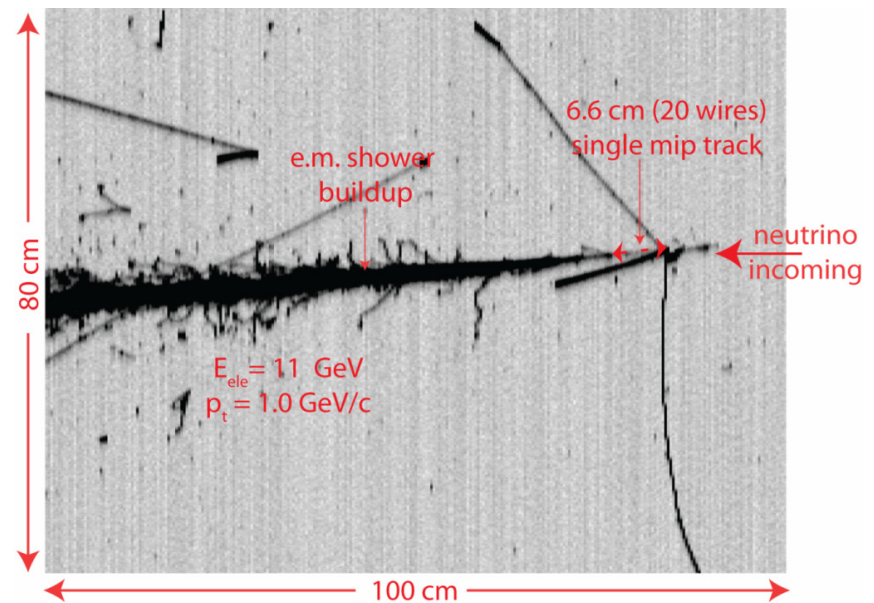
- Search for  $\nu_{\mu} \rightarrow \nu_e$  appearance in CNGS beam neutrinos
  - $L = 730 \text{ km}$ ,  $E = 10 - 30 \text{ GeV}$
- Differences w.r.t. the LSND experiment:
  - $L/E \approx 1 \text{ m/MeV}$  at LSND
  - $L/E \approx 36.5 \text{ m/MeV}$  at CNGS
  - LSND-like short distance oscillation signal averages to:  $\sin^2(1.27\Delta m^2 L/E) \approx \frac{1}{2}$
  - $\langle P \rangle_{\nu_{\mu} \rightarrow \nu_e} \approx \frac{1}{2} \sin^2(2\theta)$
- ICARUS operates in a  $L/E$  region in which contributions from standard neutrino oscillations are not yet too relevant



# Selection of $\nu_e$ events

- the signature of  $\nu_e$  is observed visually
- primary vertex min. 5 cm from wall and min. 50 cm from downstream wall to allow shower identification
- visible energy  $< 30$  GeV ( $\sim 50\%$  reduction on the  $\nu_e$  from the beam and only 15% signal events rejected)
- $\nu_\mu$  CC events identified by track without hadronic interaction with length  $> 250$  cm from the primary vertex (a muon candidate)
- A charged track from primary vertex, m.i.p. on 8 wires ( $dE/dx < 3.1$  MeV/cm), developing into EM-shower
- Clear separation ( $150$  mrad) from other tracks at the vertex in at least one of two transverse views
- Selection efficiency validated on MC sample of the  $\nu_e$  events (3 independent scanners):

$$\eta = 0.74 \pm 0.05$$



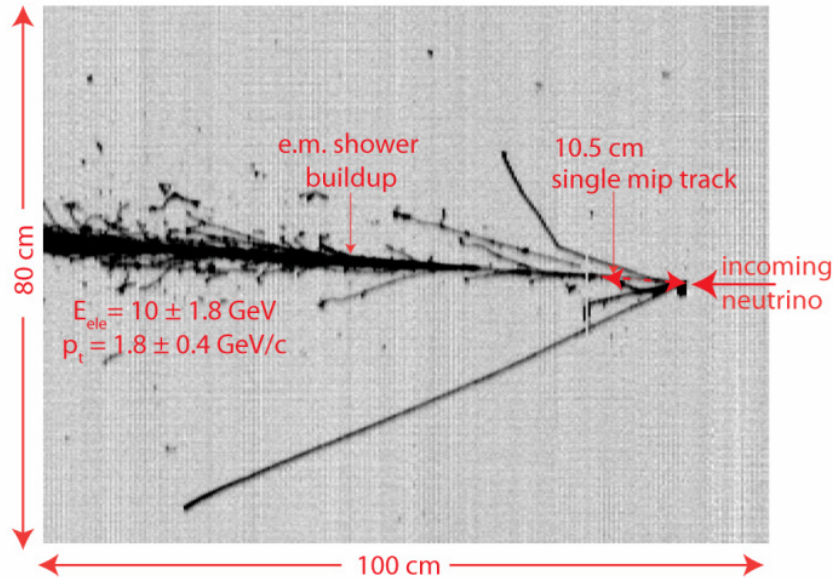
MC event (only the vertex region is shown)  
 $E_V = 11$  GeV,  $p_t = 1.0$  GeV/c

# Event rates

- Presented results ([arXiv:1307.4699](https://arxiv.org/abs/1307.4699) Jul 2013) refer to 1995 neutrino interactions ( $6.0 \cdot 10^{19}$  pot)
- The expected number of  $\nu_e$  events due to conventional sources:
  - $5.7 \pm 0.8$  events due to  $\nu_e$  beam contamination,
  - $2.3 \pm 0.5$  events due to the oscillations  $\nu_\mu \rightarrow \nu_e$
  - $1.3 \pm 0.1$  events due to the oscillations  $\nu_\mu \rightarrow \nu_\tau$  with  $\tau \rightarrow e$
- Taking into account the selection efficiency, the expected number of the  $\nu_e$  events is then:  $6.4 \pm 0.9$  (syst. only)
- $4 \nu_e$  events observed in the data

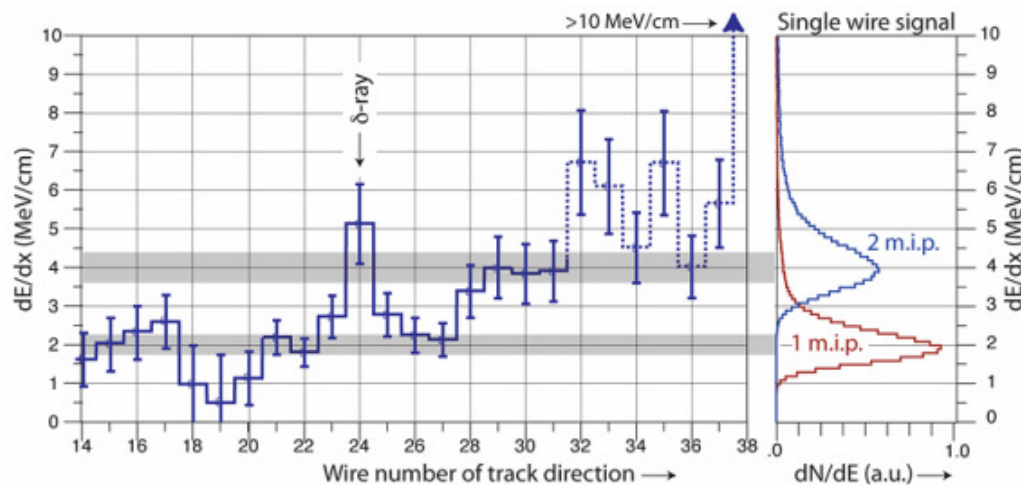


# One of four candidate signal events observed in the data



$$E_{\text{tot}} = 11.5 \pm 1.8 \text{ GeV},$$

$$p_t = 1.8 \pm 0.4 \text{ GeV}/c$$



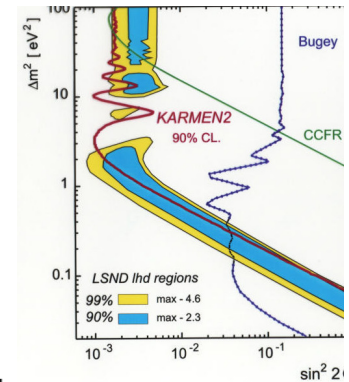
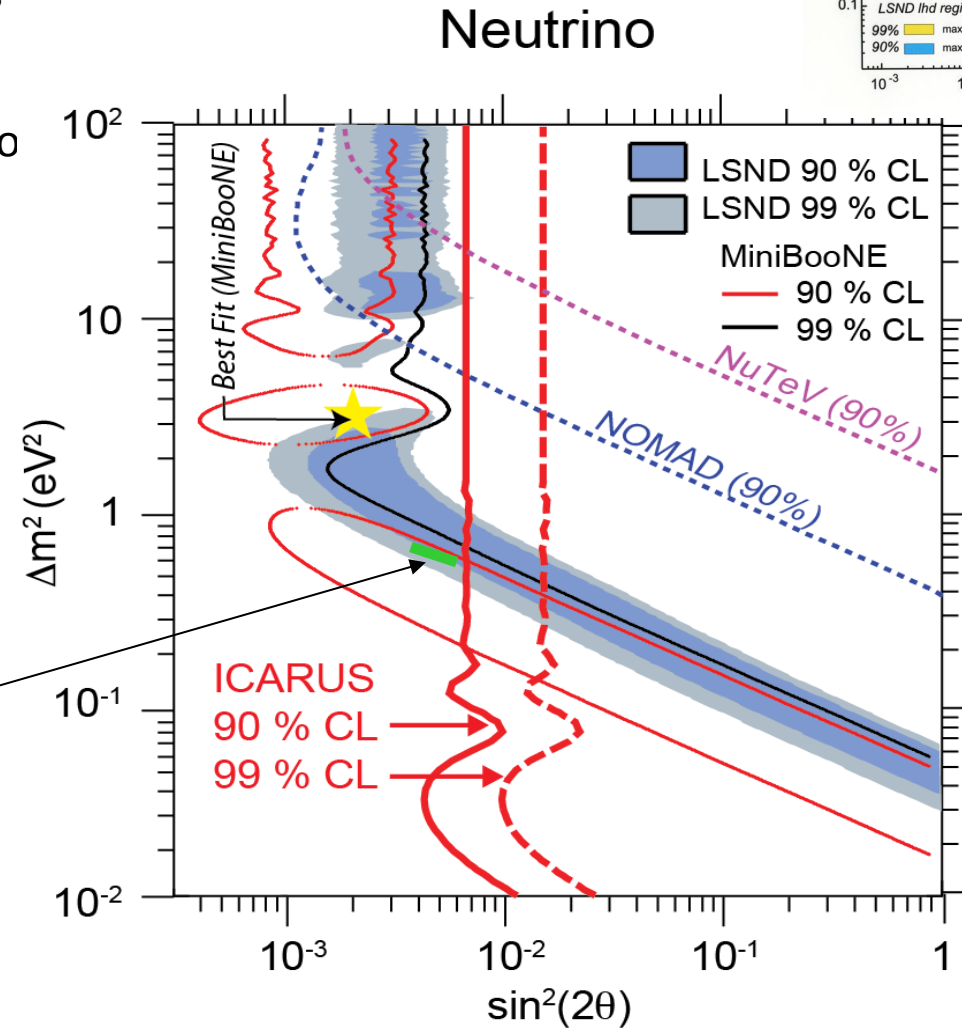
In all events: single electron shower clearly separated from hadronic component in the transverse plane

# ICARUS results on the LSND anomaly

- **4  $\nu_e$  events** observed considering  **$6.4 \pm 0.9$**  expected background
- Limits on number of events due to LSND anomaly:
  - **3.68 (90% CL)**
  - **8.34 (99% CL)**
- the corresponding limits on oscillation probability are:
  - **$P(\nu_\mu \rightarrow \nu_e) \leq 3.4 \cdot 10^{-3}$  (90% CL)**
  - **$P(\nu_\mu \rightarrow \nu_e) \leq 7.6 \cdot 10^{-3}$  (99% CL)**

$$P(\nu_\mu \rightarrow \nu_e) \approx 1/2 \sin^2(2\theta)$$

„LSND anomaly” surviving area 90% CL based on all experimental results



ICARUS result strongly limits the window of possible parameters for LSND anomaly indicating a narrow region:  **$0.4 < \Delta m^2 < 0.7$**  and  **$0.002 < P(\nu_\mu \rightarrow \nu_e) < 0.003$**

# Conclusions

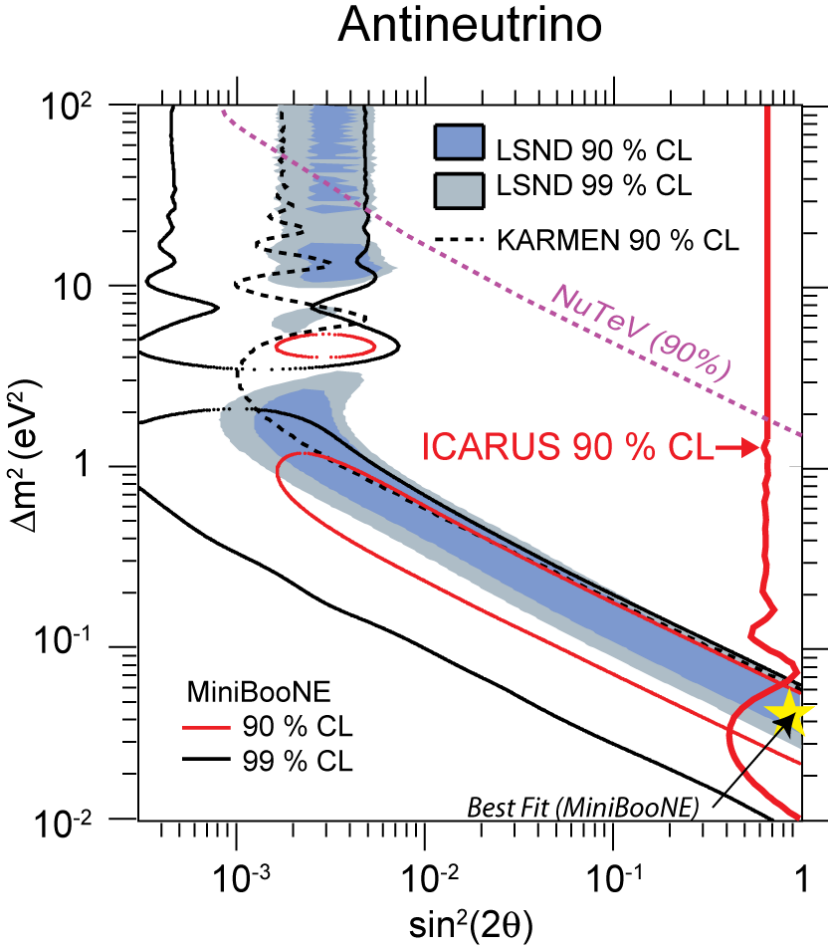
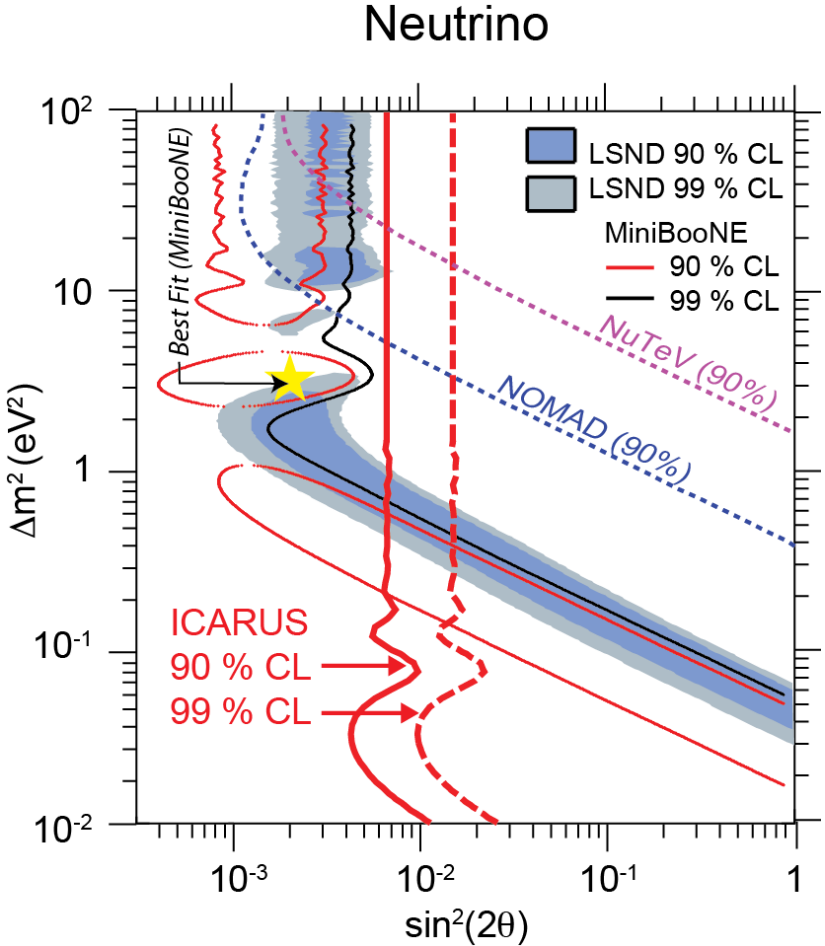
- The ICARUS result strongly limits the window of possible parameters for LSND anomaly indicating a narrow region around:

$$(\Delta m^2, \sin^2 2\theta) = (0.5 \text{ eV}^2, 0.005)$$

- where there is an overall agreement (90 % CL) between:
  - the present ICARUS limit
  - the limits of KARMEN
  - the positive signals of LSND and MiniBooNE
- The OPERA experiment also performing on the CNGS beam has confirmed our finding with a bit greater limit:  $\sin^2(2\theta) < 7.2 \times 10^{-3}$
- Definitive answer to the sterile  $\nu$  search would be given by the experiment at CERN with ICARUS T600 and a new 150 t detectors at 1.6 km and 460 m baselines exposed to a new 2GeV neutrino beam.

**BACK-UP**

# ICARUS results on the LSND-like anomaly



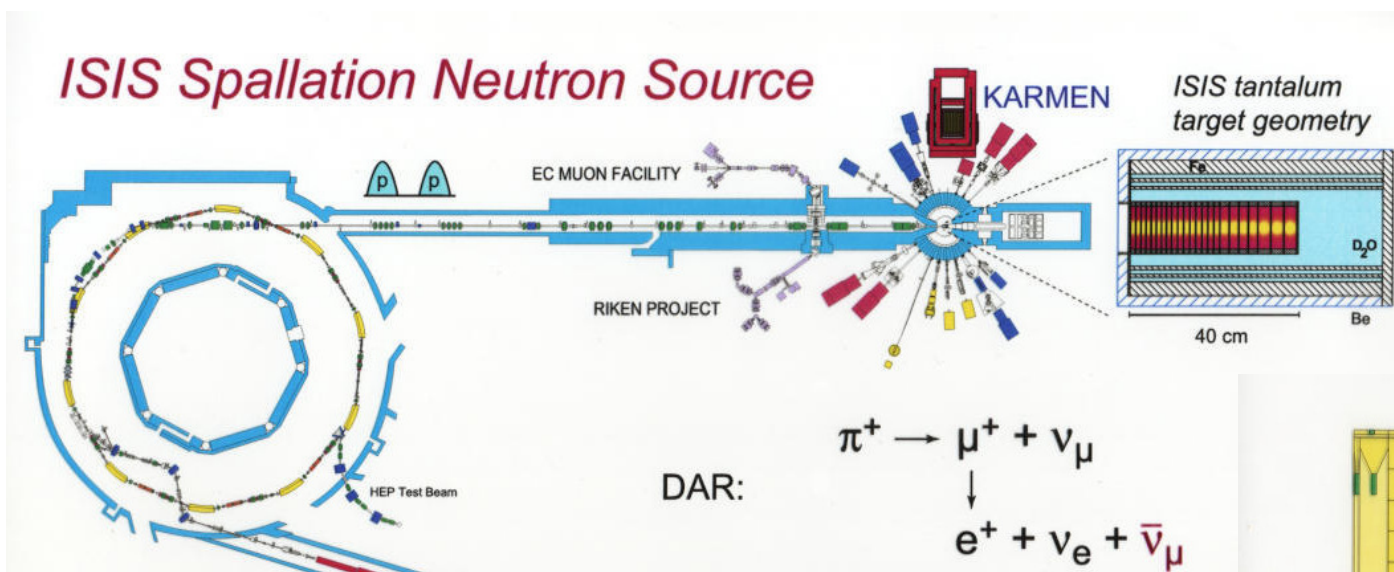
1997 - 2001

# KARMEN

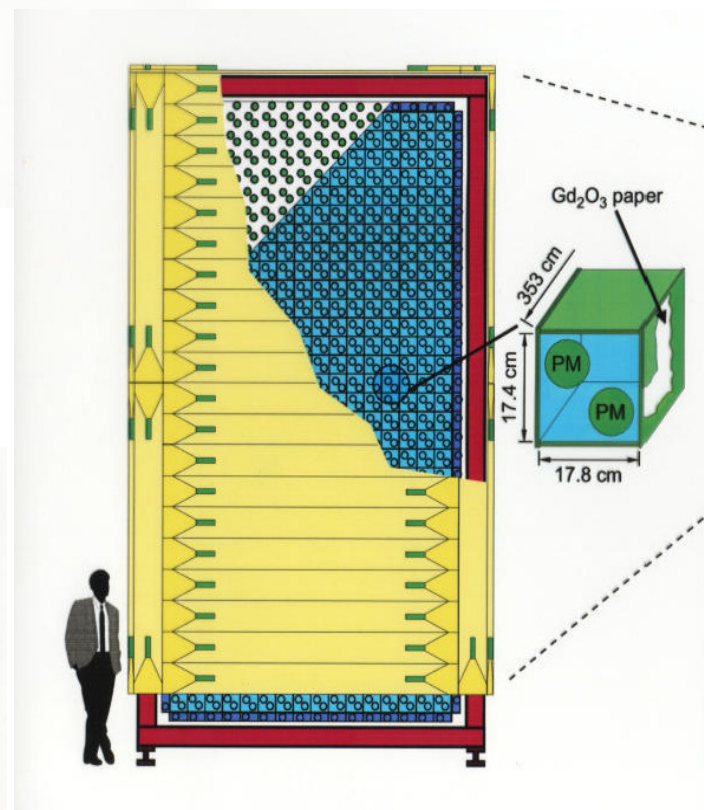
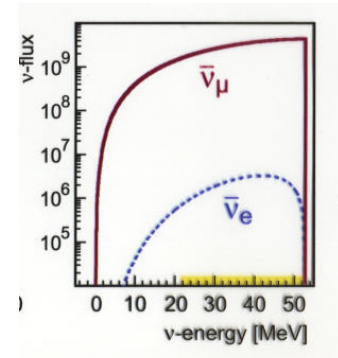


therm. + capt.

- $Q = -1.8 \text{ MeV}$
- $\text{Gd}(n, \gamma)$
- $\Sigma E_\gamma = 8 \text{ MeV}$
- $p(n, \gamma)$
- $E_\gamma = 2.2 \text{ MeV}$

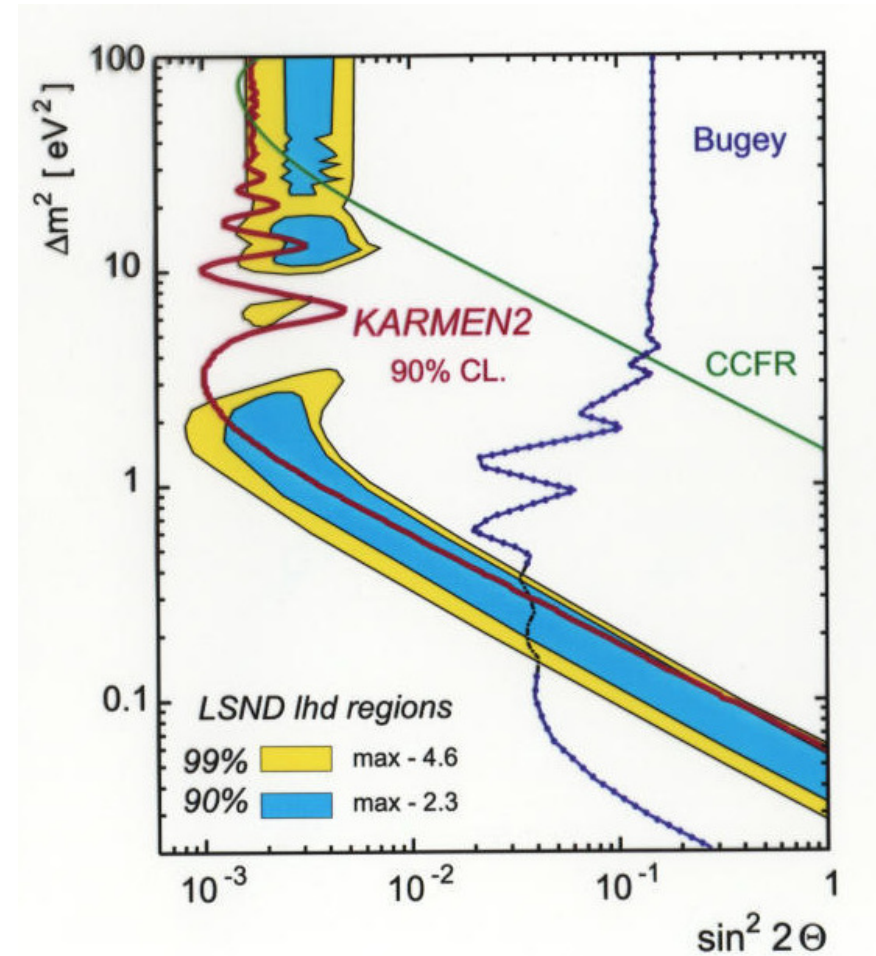


- 50 t scintillator (5.6 x 3.2 x 3.5) m<sup>3</sup>
- $L \approx 18 \text{ m}$
- $E < 50 \text{ MeV}$
- $L/E = 0.4 - 1$



# KARMEN – no oscillation excess

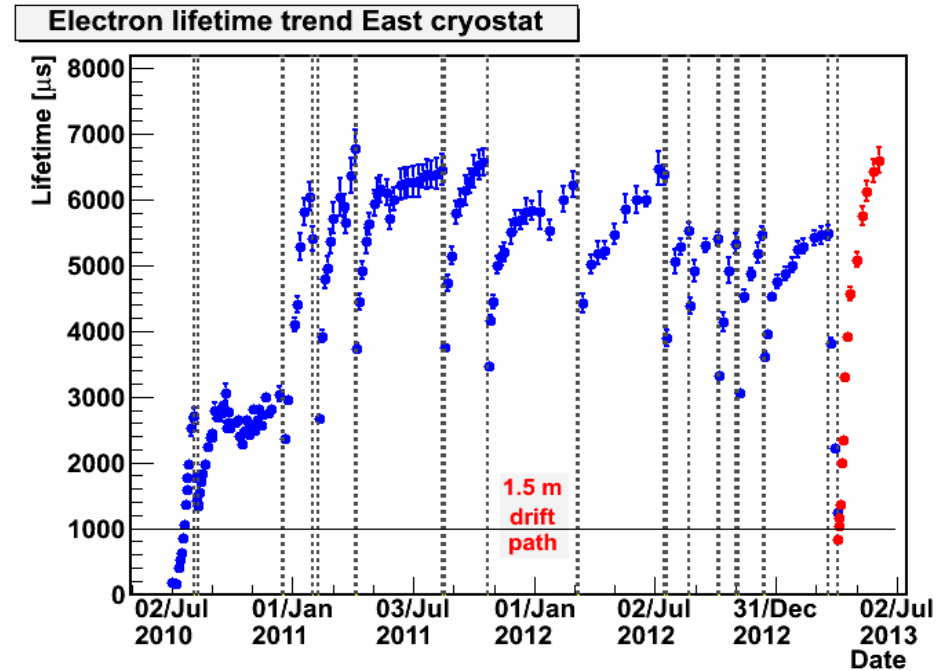
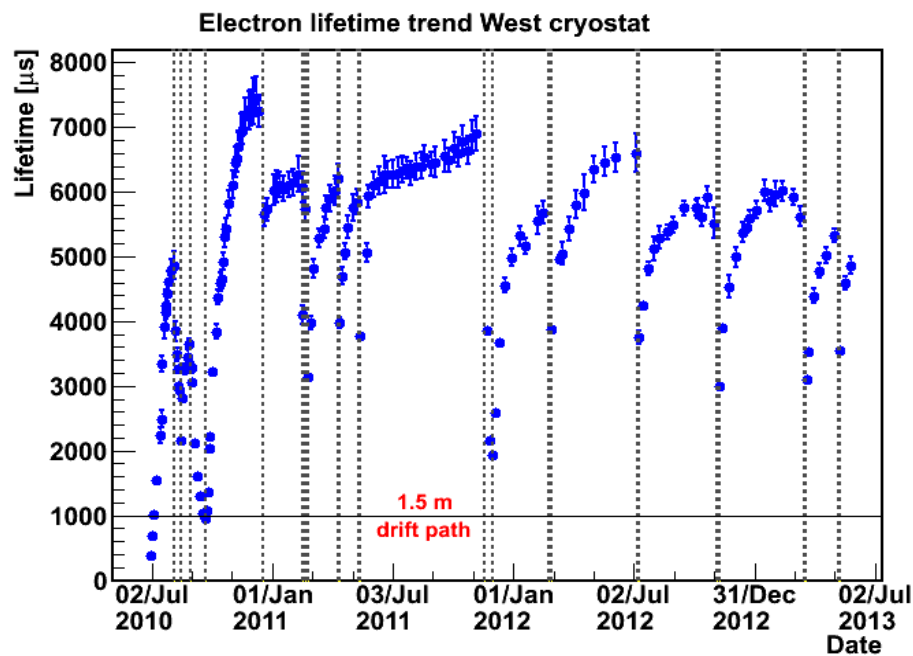
- candidate events : 15
- background :  $15.8 \pm 0.5$ 
  - cosmic :  $3.9 \pm 0.2$
  - $\nu_e$  from CC :  $9.9 \pm 0.4$
  - $-\nu_e$  contamination :  $2.0 \pm 0.2$
- oscillation limit:
  - for  $\Delta m^2 > 1 \text{ eV}^2$  :  $\sin^2\theta < 0.0017$
  - $0.2 < \Delta m^2 < 1 \text{ eV}^2$  :  $10^{-3} < \sin^2\theta < 3 \cdot 10^{-2}$



# LAr purification

Key feature: LAr purity from electro-negative molecules ( $O_2$ ,  $H_2O$ ,  $CO_2$ ).

LAr continuously filtered, electron ionization life-time is measured

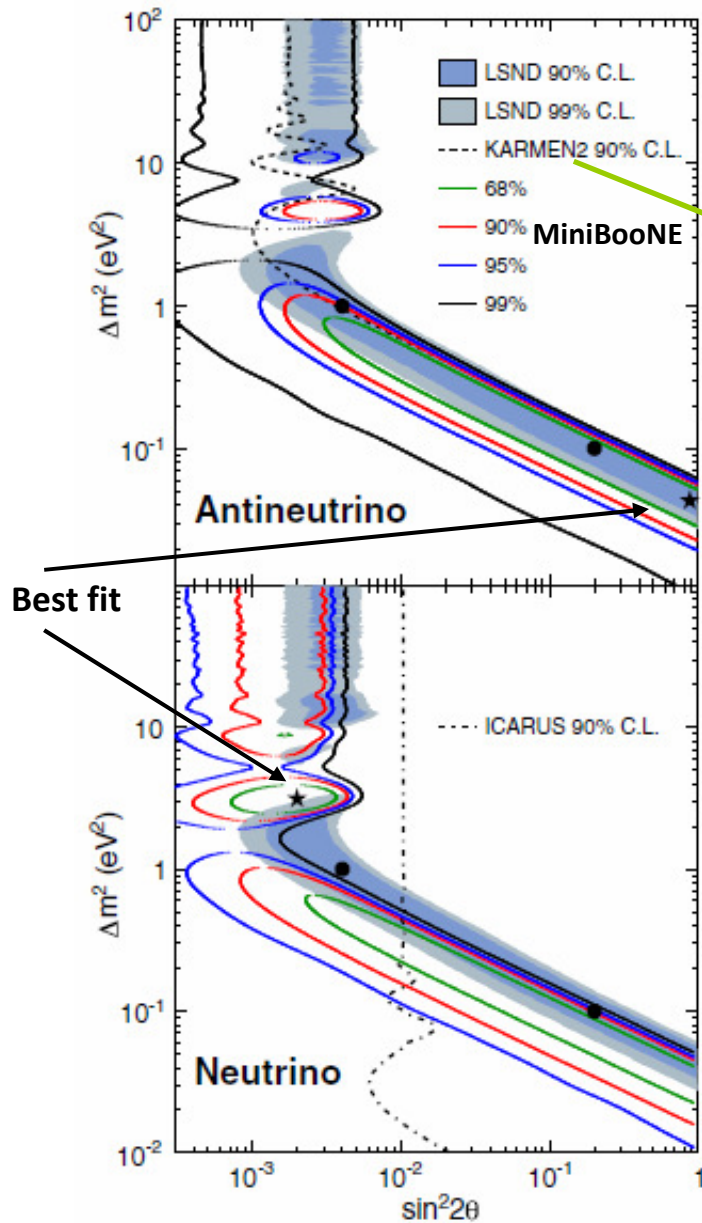


corresponding to a maximum charge attenuation of 17% at 1.5m

Would allow operation at larger drift distances

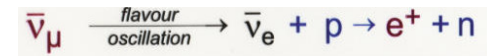


# LSND anomaly area

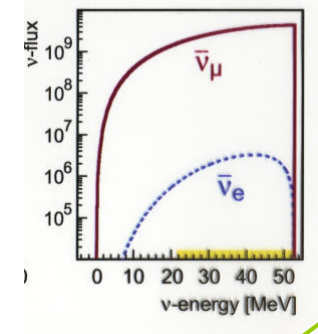


## KARMEN – negative evidence

1997 - 2001



- 50 t scintillator (5.6 x 3.2 x 3.5) m<sup>3</sup>
- $L \approx 18$  m     $E < 50$  MeV
- background :  $15.8 \pm 0.5$
- candidate events : 15



Other exp. with the negative evidence:

NOMAD, BUGEY, NUTEV

MiniBooNE : PRL 110, 161801 (19 April 2013)