

On extended GMSB models

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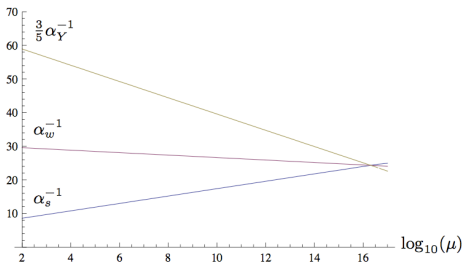
based on: arXiv:1305.6277 [hep-ph]

Matter To The Deepest 2013, Ustroń, 1-6 September

1. MSSM

Still the best candidate for BSM is **softly broken MSSM**:

- solves problem of quadratic corrections to m_{h^0}
- dark matter candidate \rightarrow LSP
- better unification of gauge couplings at 10^{16} GeV \rightarrow hint for GUT model



Problems:

- one needs **additional sectors**: SUSY breaking and mediation
- fine-tuning
- **a lot of parameters** (soft terms) \rightarrow explain them using RGE and some simple set of initial conditions at high scale \rightarrow **GUT model**

2. LHC vs. MSSM

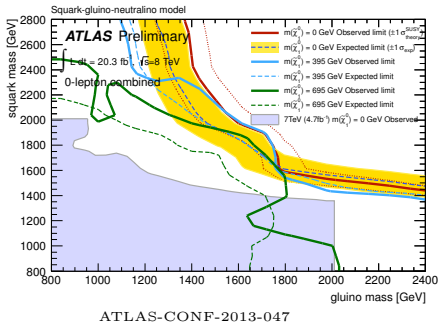
What do the LHC searches tell us about MSSM?

- no SUSY signal so far
- relevant exclusions only for 1st and 2nd family
- still \tilde{Q}_3, \dots can be as light as 500 GeV

BUT important information comes from Higgs mass measurement:

- $m \sim 125$ GeV \rightarrow need for large loop corrections

ASSUME other MSSM Higgses are much heavier and masses of $\tilde{Q}_{1,2}$ and \tilde{g} are bigger than 1.8 TeV.



3. 1-loop corrections to m_{h^0}

- dominant contribution from top quarks and stops (due to $y_t \sim 1$):

$$\Delta(m_{h^0}^2) = h^0 - \text{---} \left(\begin{array}{c} t \\ \text{---} \text{---} \text{---} \end{array} \right) \text{---} + h^0 - \text{---} \left(\begin{array}{c} \tilde{t} \\ \text{---} \text{---} \text{---} \end{array} \right) \text{---} + h^0 - \text{---} \left(\begin{array}{c} \tilde{t} \\ \text{---} \text{---} \text{---} \end{array} \right) \text{---}$$

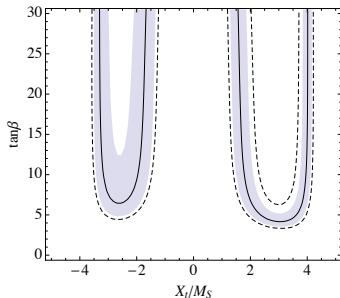
$$m_{h^0}^2 = m_Z^2 \cos^2 2\beta + \frac{3m_t^4}{4\pi^2 v^2} \left[\ln \frac{M_S^2}{m_t^2} + \frac{X_t^2}{M_S^2} \left(1 - \frac{X_t^2}{12M_S^2} \right) \right] \approx (125 \text{ GeV})^2,$$

$$M_S = \sqrt{m_{\tilde{t}_1} m_{\tilde{t}_2}}$$

$$X_t = A_t - \mu \cot \beta$$

A-terms:

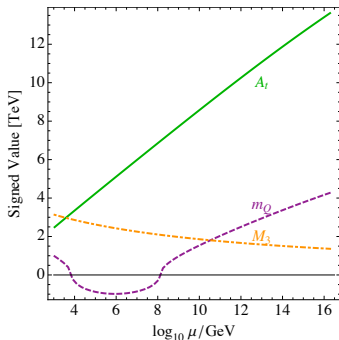
$$V_{\text{soft}} \supset y_t A_t H_u \tilde{Q}_3 \tilde{U}_3^c \longrightarrow y_t A_t h_0 \tilde{t}_1 \tilde{t}_2$$



Draper et al. 1112.3068

4. How to generate large A -terms?

- value of A -term gives initial condition for RGE evolution

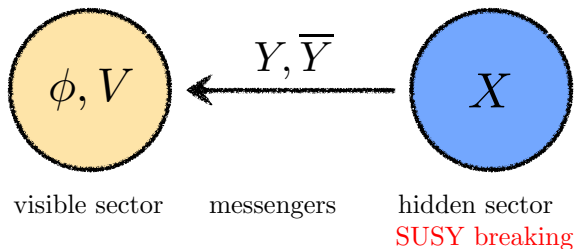


Draper et al. 1112.3068

$$\mu \frac{dA_t}{d\mu} \sim y_t^2 A_t + g_3^2 M_3$$

- how to get A -terms in GUT model?

5. SUSY breaking mediation



- such structure is dictated by SUSY
- mediation = interactions between Y, \bar{Y} and other fields
- **singlet** $\langle X \rangle = M + \theta^2 F \rightarrow$ spontaneous SUSY breaking
- X does not interact via superpotential with visible sector
- messengers have large masses e.g. $M \sim 10^{14}$ GeV

6. SUSY breaking mediation

- supergravity
 - no control over FCNC at all \rightarrow arbitrary mixings between families
- gauge interactions
 - no FCNC effects at M scale (small mixing generated via RGE)
 - $A \approx 0$ at M scale
 - $m_{h^0} \sim 125$ GeV $\rightarrow M \gtrsim 10^{14}$ GeV (i.e. A needs long RGE evolution)
- Yukawa (and gauge) interactions \rightarrow EXTENDED GMSB MODELS
 - FCNC effects controlled by Yukawa-type couplings of messengers
 - A -terms $\neq 0$ at M scale
 - easy to satisfy $m_{h^0} \sim 125$ GeV even for $M \sim 10^8$ GeV
 - rich phenomenology

7. Messenger couplings in EGMSSB models

Focus on: $SU(5)$ unification model with messengers in $5 + \bar{5}$ and $10 + \bar{10}$

- matter Φ_i in $5, \bar{5}$ or 10 (in MSSM only $5\ 10\ 10$ and $\bar{5}\ \bar{5}\ 10$)
- pair of messengers $Y = (Y, \bar{Y})$

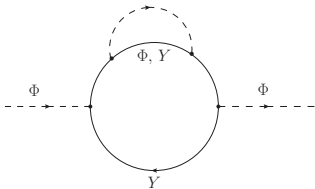
$$W_Y = \eta YYY + h_I^i \Phi_i YY + h_{II}^{ij} \Phi_i \Phi_j Y$$

- allowed couplings: $5\ 10\ 10$, $\bar{5}\ \bar{5}\ 10$, $\bar{5}\ \bar{10}\ \bar{10}$, $5\ 5\ \bar{10}$
- $h_{I,II}$ quite well explored (Yukawa-Deflected Gauge Mediation)
usually some hierarchy for $h_{I,II}$ is assumed
- couplings of three messengers $\eta \rightarrow$ additional effects!
 - relevant only if occur together with h_I or h_{II}
 - do not contribute to A -terms (nor to 1-loop masses)

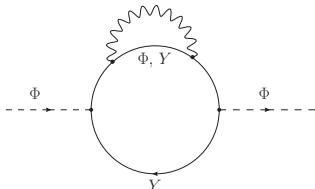
8. Soft terms

- 2-loop contributions to soft masses

$$W_Y = \eta Y Y Y + h_I^i \Phi_i Y Y + h_{II}^{ij} \Phi_i \Phi_j Y$$



$$\sim h^4, h^3 y, h^2 y^2$$

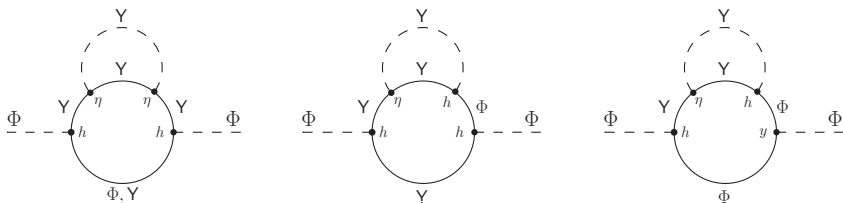


$$\sim h^2 g^2$$

$$m_{\tilde{\Phi}, h}^2 \sim \frac{1}{(4\pi)^4} (h^4 + h^3 y + h^2 y^2 + h^2 g^2) \left(\frac{F}{M} \right)^2$$

9. New contributions to soft masses

$$W_Y = \eta YYY + h_I^i \Phi_i YY + h_{II}^{ij} \Phi_i \Phi_j Y$$



$$m_{\Phi, \eta}^2 \sim \frac{1}{(4\pi)^4} (\eta^2 h^2 + \eta h^3 + \eta h^2 y) \left(\frac{F}{M} \right)^2$$

10. Phenomenological constraints

Assumptions:

- no rapid proton decay via

$$\phi_{\bar{5}}\phi_{\bar{5}}\phi_{10}, \quad \frac{1}{M}\phi_{\bar{5}}\phi_{10}\phi_{10}\phi_{10} \quad \frac{1}{M^2}(\phi_{10}^\dagger\phi_{10})^2$$

- absence of μ/B_μ problem
- no $\mu H_u H_d$ term in the superpotential; but Higgs mass via

$$\frac{1}{M_{GUT}} X^\dagger H_u H_d$$

One needs additional selection rules \rightarrow e.g. [global \$U\(1\)\$ symmetry](#)

[Examples?](#)

11. The simplest model

- $U(1)$ charges

H_5	$H_{\bar{5}}$	$\phi_{\bar{5}}$	ϕ_{10}	Y_5	$Y_{\bar{5}}$	Y_{10}	$Y_{\bar{10}}$	X
$-8q$	$-7q$	$3q$	$4q$	$17q$	$-2q$	$14q$	q	$-15q$

- Yukawa-type couplings of messengers

$$W_Y = \frac{1}{2} h_{14} \phi_{10} Y_{\bar{5}} Y_5 + \frac{1}{2} \eta_2 Y_{\bar{5}} Y_{\bar{10}} Y_{10}.$$

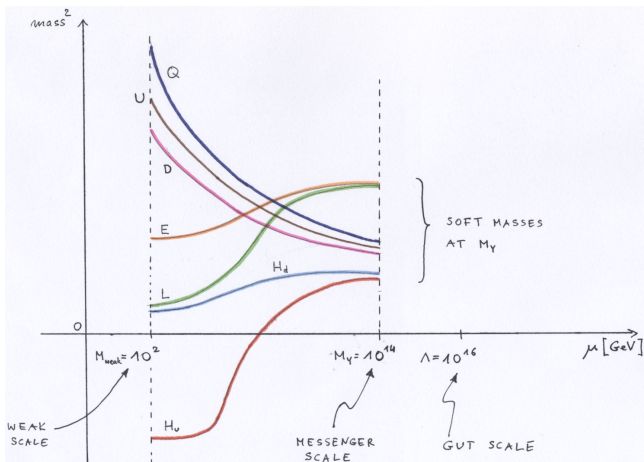
- soft terms

$$m_{\tilde{Q},\eta}^2 = 6\alpha_{h_{14}} \alpha_{\eta_2} \frac{\xi^2}{16\pi^2}, \quad m_{\tilde{Q},h}^2 = \alpha_{h_{14}} \left(6\alpha_{h_{14}} - \frac{7}{15}\alpha_1 - 3\alpha_2 - 6\alpha_3 \right) \frac{\xi^2}{16\pi^2}$$

$$A_t = A_b = -\alpha_{h_{14}} \frac{\xi}{4\pi}$$

12. Top-down analysis

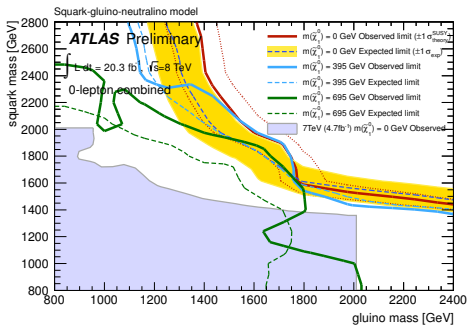
Reverse the initial problem and evolve parameters from M to EWSB scale:



13. Phenomenology

Find spectrum and check if phenomenology is correct i.e.

- $m_{h^0} \approx 125$ GeV
- no tachyons
- scalar potential bounded from below, no CCB
- $a_\mu, b \rightarrow s\gamma$
- ATLAS bounds on gluino and squarks of 1. and 2. generation



ATLAS-CONF-2013-047

14. The simplest model

$$W_Y = \frac{1}{2} \eta_2 Y_5^- Y_{10}^- Y_{10}^- + \frac{1}{2} h_{14} \phi_{10} Y_5^- Y_5^-$$

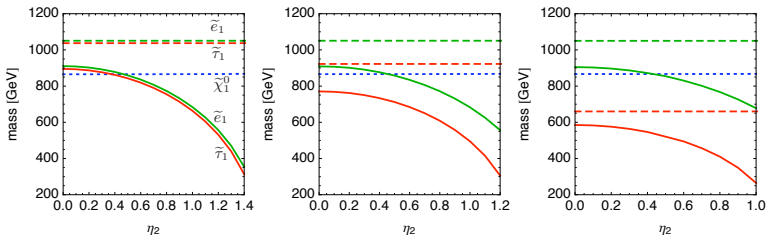


Figure : Plot of the particles masses vs. η_2 coupling for $\tan \beta = 10$ (left plot), $\tan \beta = 30$ (middle plot) and $\tan \beta = 50$ (right plot). h_{14} is set to 1.2, while $\xi = F/M$ scale is 1.6×10^5 GeV. Dashed lines show masses of the particles when $h_{14} = \eta_2 = 0$, which corresponds to the standard GMSB case. $\tilde{\tau}_1$ and \tilde{e}_1 are mostly right-handed.

15. The simplest model

$$W_Y = \frac{1}{2} \eta_2 Y_{\bar{5}} Y_{10} Y_{10} + \frac{1}{2} h_{14} \phi_{10} Y_{\bar{5}} Y_{\bar{5}}$$

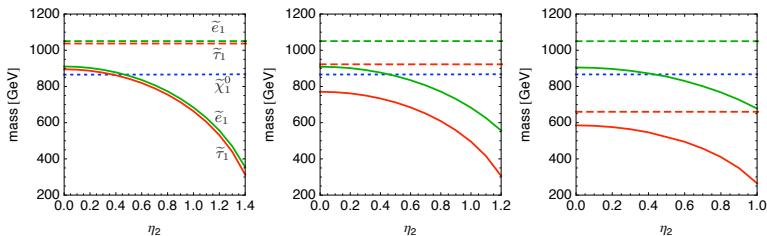


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$$m_{\tilde{Q}, \eta}^2 = 6\alpha_{h_{14}} \alpha_{\eta_2} \frac{\xi^2}{16\pi^2}, \quad m_{\tilde{Q}, h}^2 = \alpha_{h_{14}} \left(6\alpha_{h_{14}} - \frac{7}{15}\alpha_1 - 3\alpha_2 - 6\alpha_3 \right) \frac{\xi^2}{16\pi^2}$$

$$\mu \frac{d}{d\mu} m_{\tilde{E}_3}^2 = \dots + \frac{6}{10} g_1^2 m_{\tilde{Q}, \eta}^2$$

16. Conclusions

- **Extended GMSB models** naturally accommodate for left-right top squarks mixing
- in some cases superpotential couplings of **three messengers** are relevant to mass spectrum
- additional selection rules (e.g. global $U(1)$) are necessary to satisfy phenomenological constraints