

# Recent measurements of $\theta_{13}$ mixing angle in neutrino oscillation experiments



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# Layout of the presentation

- Neutrino oscillations
- How to measure  $\theta_{13}$  mixing angle
- Results from reactor experiments
  - Daya Bay
  - Reno
  - Double Chooz
- Results from long baseline experiments
  - T2K
  - Minos

# Neutrino oscillations

FLAVOR

PMNS mixing matrix

MASS

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & \cos\theta_{23} & \sin\theta_{23} \\ 0 & -\sin\theta_{23} & \cos\theta_{23} \end{pmatrix} \begin{pmatrix} \cos\theta_{13} & 0 & \sin\theta_{13}e^{-i\delta} \\ 0 & 1 & 0 \\ -\sin\theta_{13}e^{-i\delta} & 0 & \cos\theta_{13} \end{pmatrix} \begin{pmatrix} \cos\theta_{12} & \sin\theta_{12} & 0 \\ \sin\theta_{12} & \cos\theta_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

„atmospheric”  
SK, K2K, T2K, MINOS

CHOOZ,  
DayaBay,  
Reno,  
DblChooz,  
T2K

„solar”  
SNO, KamLand, SK,  
Borexino

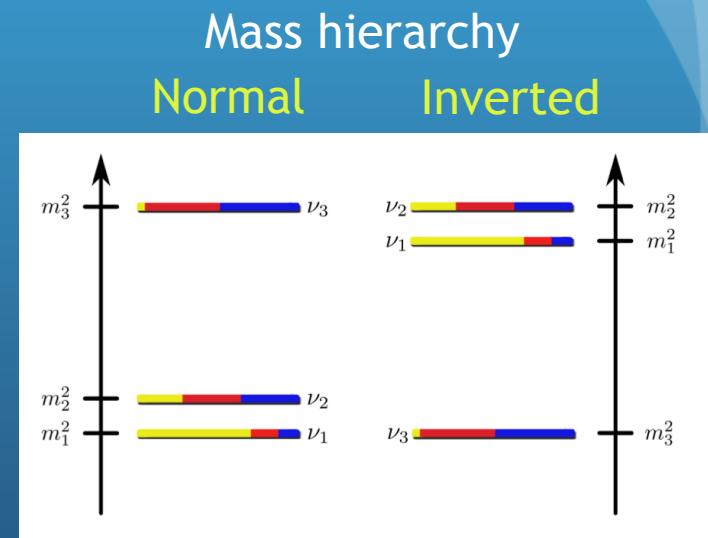
$$\Delta m_{31}^2 = \begin{cases} 2.53^{+0.08}_{-0.10} \\ -(2.40^{+0.10}_{-0.07}) \end{cases} \times 10^{-3} \text{ eV}^2$$

$$\Delta m_{21}^2 = (7.62 \pm 0.19) \times 10^{-5} \text{ eV}^2$$

mixing angles, squared mass differences, CP violation phase -  
fundamental parameters of nature

# What's so interesting about $\theta_{13}$ ?

- The last unknown mixing angle
  - difficult measurement - to get  $\theta_{13}$  one has to measure small effects (deficit or excess of interactions of neutrinos of certain flavor)
- If it is non-zero we have the possibility to study CP violation in neutrino sector, also matter effects (mass hierarchy!)



# ... and ways of measuring it

- we need to look at oscillations involving electron neutrinos or antineutrinos
- we can look at disappearance -> reactor experiments

Leading terms!



$$P_{\text{sur}} \approx 1 - \sin^2 2\theta_{13} \sin^2(1.267 \Delta m_{31}^2 L/E),$$

Energy ~ a few MeV  
Distance ~ a few km

$$\bar{\nu}_e \rightarrow \bar{\nu}_e$$

- we can look at appearance -> long-baseline experiments with  $\nu_\mu$  beam



$$P(\nu_\mu \rightarrow \nu_e) = \sin^2 2\theta_{13} \sin^2 \theta_{23} \sin^2(1.27 \Delta m_{23}^2 L/E)$$

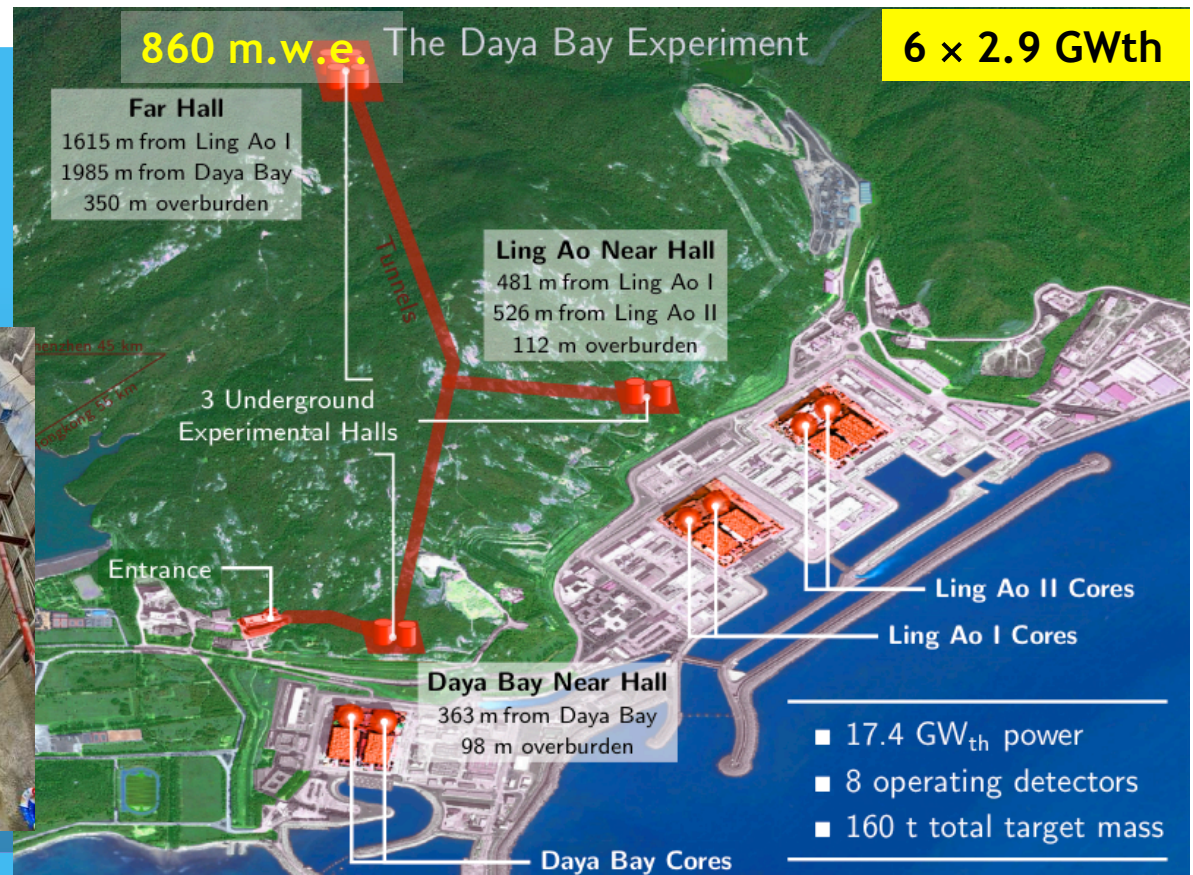
Energy ~ a few GeV  
Distance ~ a few hundred km

$$\nu_\mu \rightarrow \nu_e$$

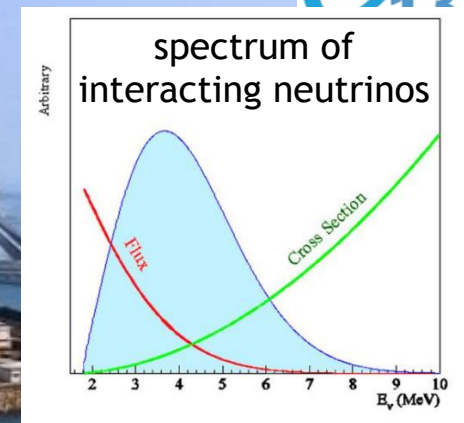
Second order terms depend on  $\delta$  and mass hierarchy

- in both cases near detectors are useful to minimize systematic errors

# Reactor experiments: Daya Bay

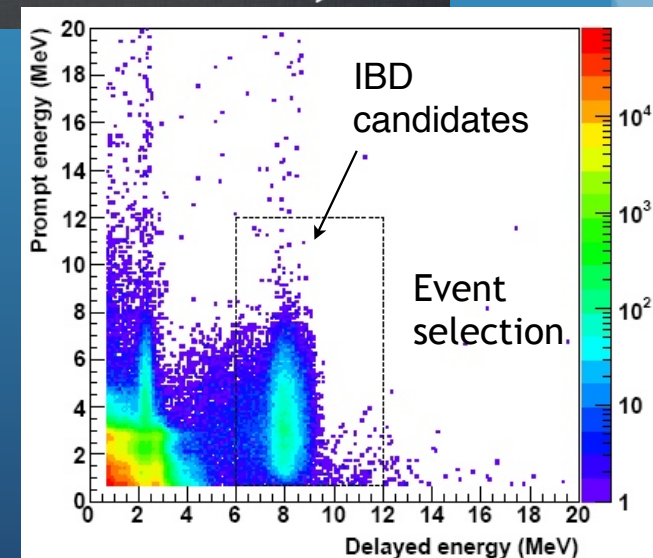
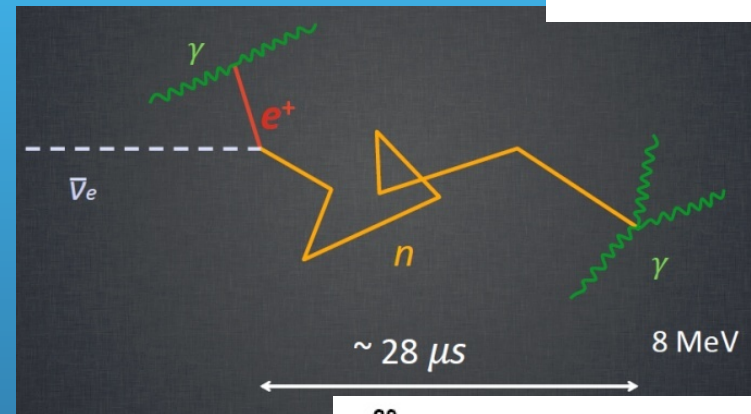
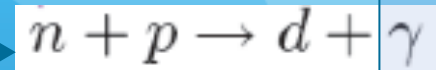
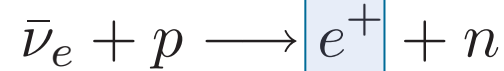


Dàya Bay (大亚湾) and Lǐng ào (岭澳) nuclear power plants



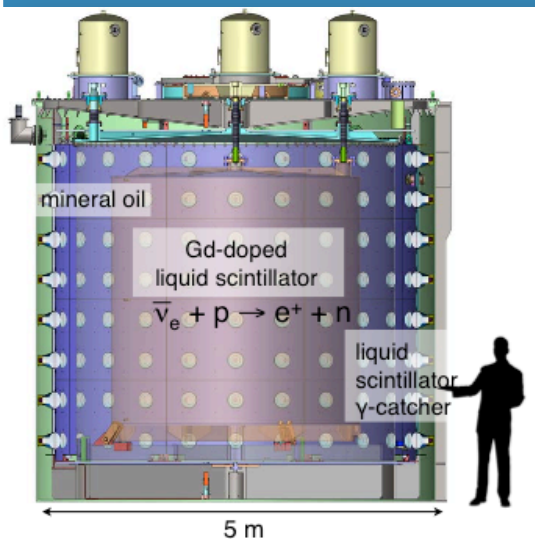
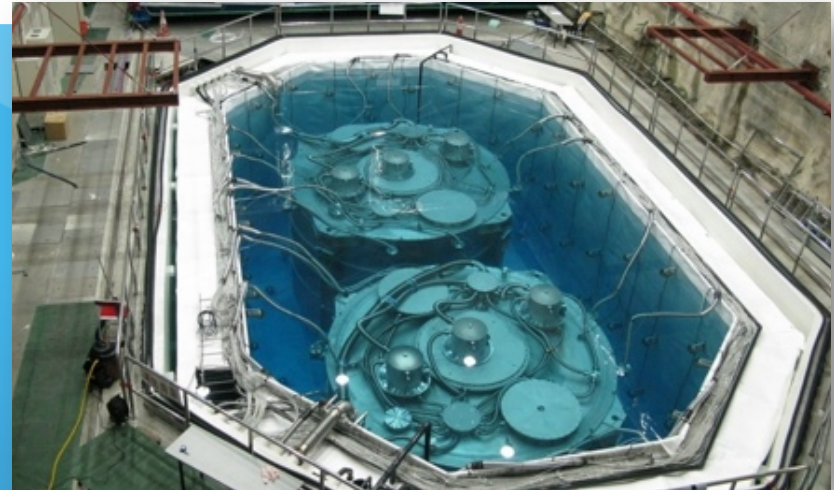
# Daya Bay: detection mechanism

- Gadolinium-doped liquid scintillator (Gd-LS)
- Inverse beta decay
  - **prompt signal:** positron (its energy is correlated with neutrino energy)
  - **delayed signal:** neutron capture
- **Systematic error minimisation:**
  - Identical far and near detectors
  - Identical detection modules
  - Good background rejection (3-zone modules, water system, RPCs)



# Daya Bay: detection modules

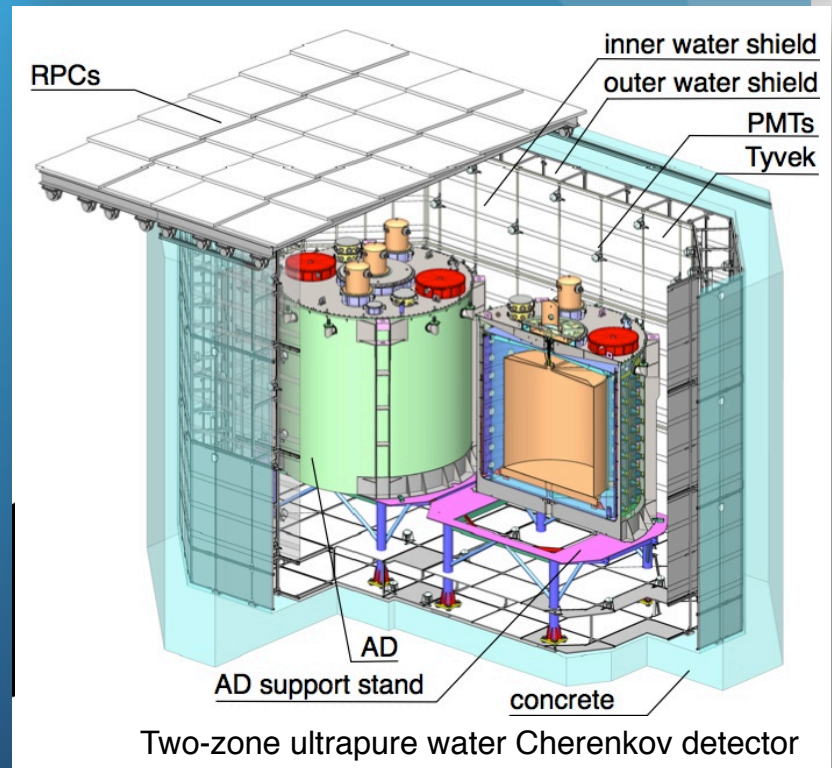
- 3-zone detection modules immersed in water, shielded by RPC plane
  - target - 20t of Gd-LS (0.1% Gd)
  - gamma catcher - 21t of LS
  - radiation shield - 37t of mineral oil
  - 192 8inch photomultipliers on the walls (scintillation light detectors)
  - 3 calibration modules



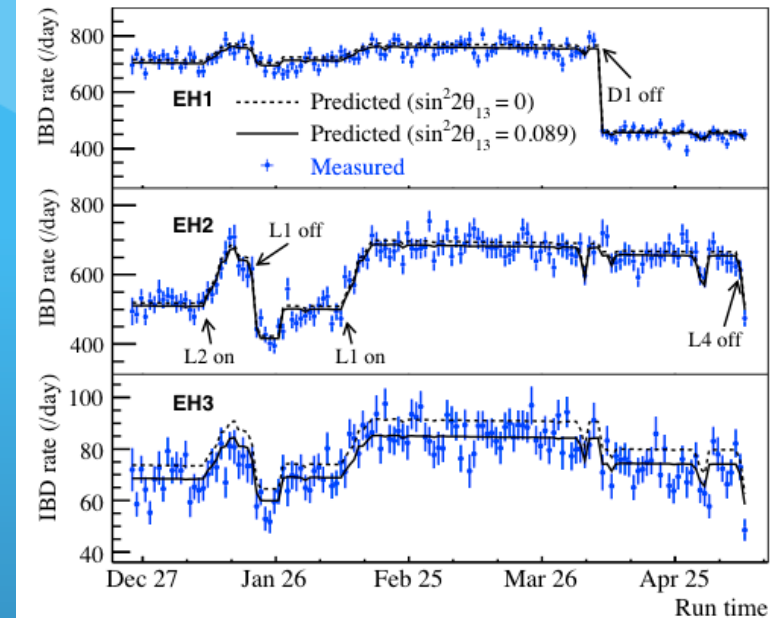
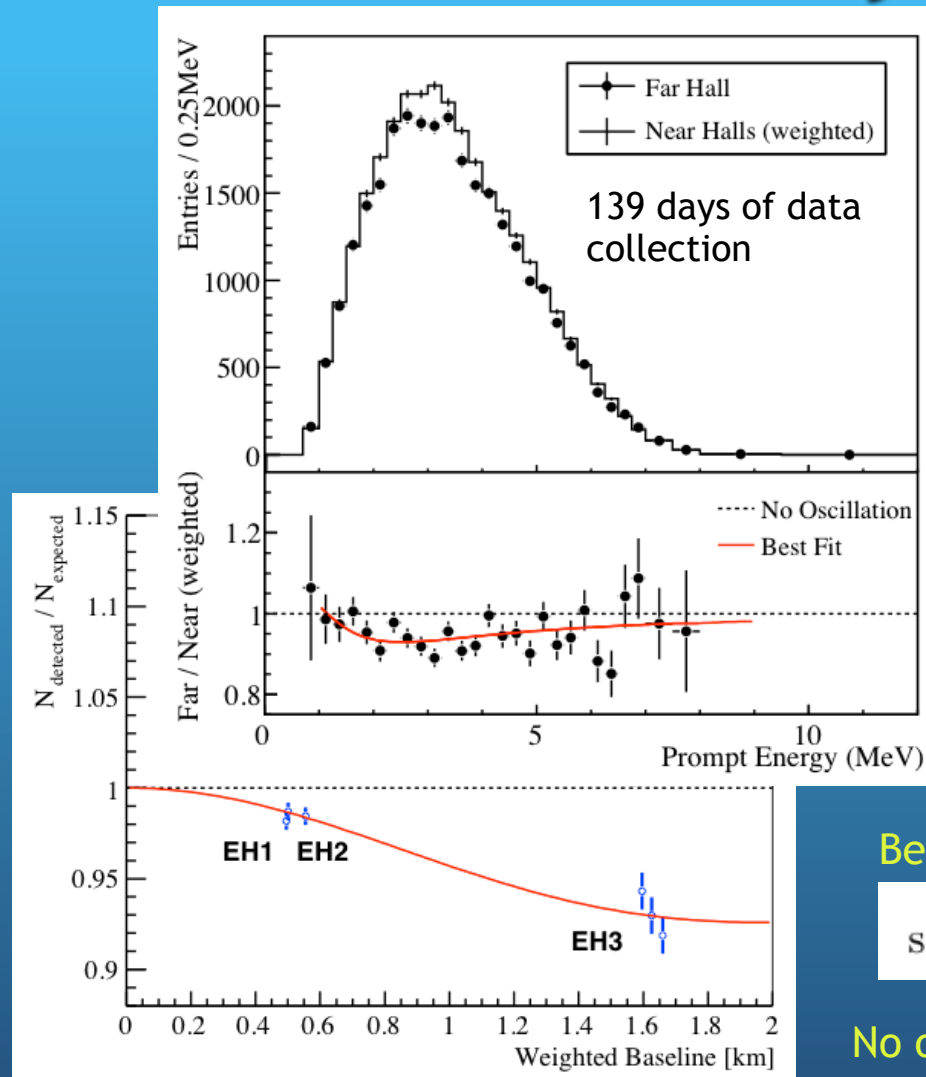
## Muon water system

- Cherenkov water shield - at least 2.5m in every direction, 1200/1950t of water in two separated vessels, each equipped with photomultipliers

Detects muons that can produce spallation neutrons, attenuates gamma rays from surroundings, moderates neutrons



# Results: rate only



Time variations: flux prediction + detector MC vs data in the three stations

Far to near ratio:

$$R = 0.944 \pm 0.007(\text{stat.}) \pm 0.003(\text{syst.}),$$

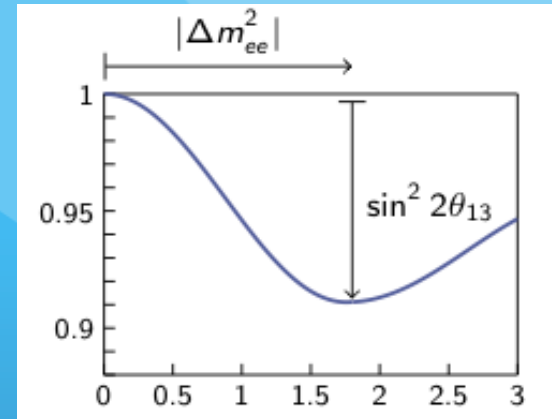
Best fit, assuming  $\Delta m_{31}^2 = 2.32 \times 10^{-3} \text{eV}^2$

$$\sin^2 2\theta_{13} = 0.089 \pm 0.010(\text{stat.}) \pm 0.005(\text{syst.})$$

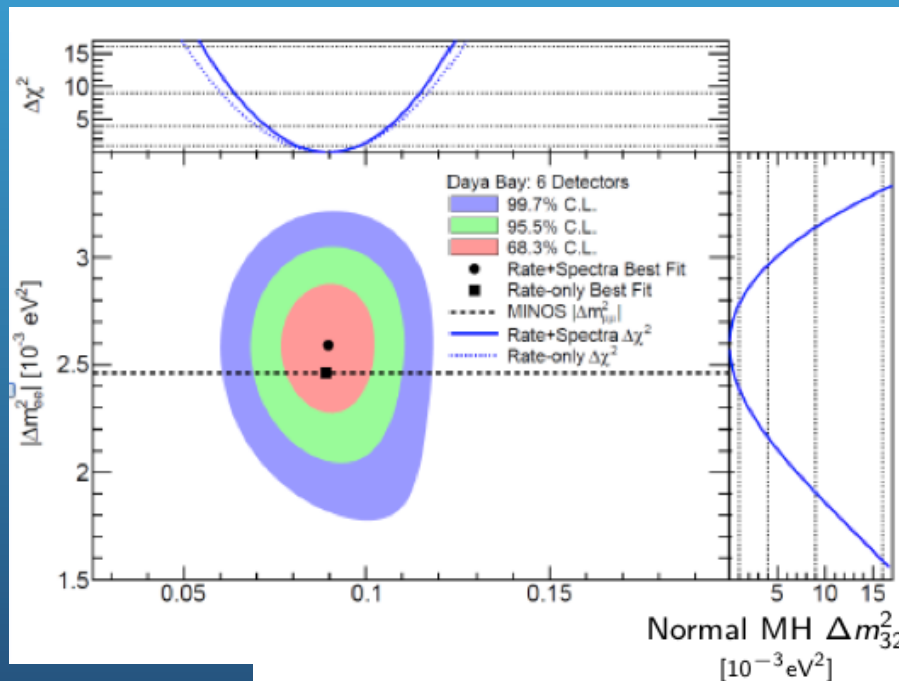
No oscillations excluded at the level of  $7.7\sigma$

# Results: rate+shape

- New analysis, made public just a week ago (no paper yet)
- We can take advantage of measuring energy spectra and fit energy dependence (obtaining  $\theta_{13}$  as well as the relevant  $\Delta m^2$ )



$$\sin^2(\Delta m_{ee}^2 \frac{L}{4E}) \equiv \cos^2 \theta_{12} \sin^2(\Delta m_{31}^2 \frac{L}{4E}) + \sin^2 \theta_{12} \sin^2(\Delta m_{32}^2 \frac{L}{4E})$$



217 days of data collection

$$\sin^2 2\theta_{13} = 0.090^{+0.008}_{-0.009}$$

$$|\Delta m_{ee}^2| = 2.59^{+0.19}_{-0.20} \cdot 10^{-3} \text{ eV}^2$$

$$\chi^2/N_{\text{DoF}} = 162.7/153$$

From Daya Bay  $\Delta m_{ee}^2$

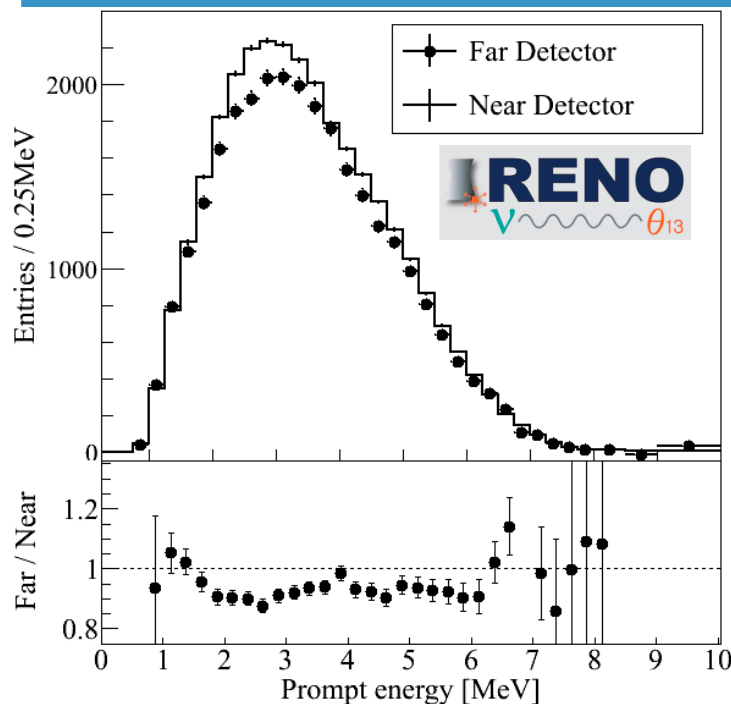
$$2.54^{+0.19}_{-0.20}$$

Inverted MH  $\Delta m_{32}^2$   
[ $10^{-3} \text{ eV}^2$ ]

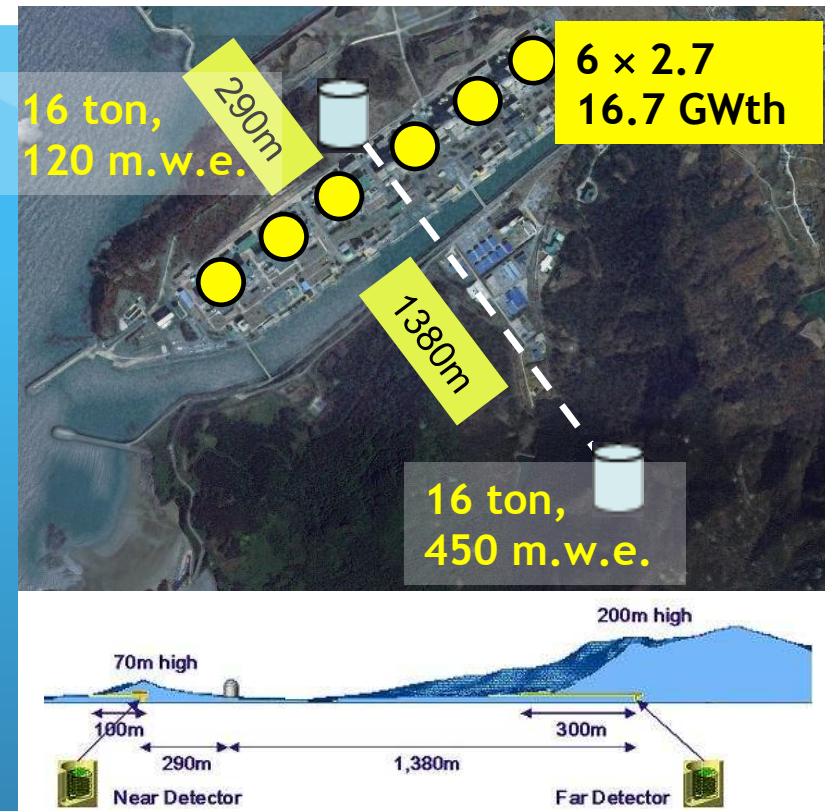
$$-2.64^{+0.19}_{-0.20}$$

# Reactor experiments: RENO

- Yeonggwang power plant in South Korea
- started operation in 2011
- rate results published, rate+shape analysis in progress



Published results: PRL 108, 191802 (2012), 2013 from NuTel conference



RENO 2013 results (402 days of data collection)  
(improved energy calibration, background estimation/reduction)

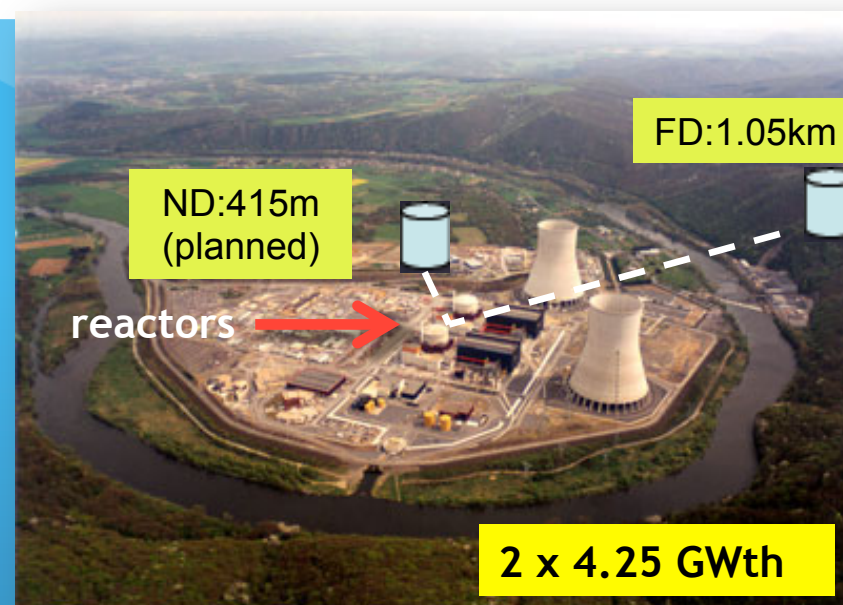
$$R = \frac{\Phi_{observed}^{Far}}{\Phi_{expected}^{Far}} = 0.929 \pm 0.006(stat) \pm 0.009(syst)$$

$$\sin^2 2\theta_{13} = 0.100 \pm 0.010(stat) \pm 0.015(syst)$$

There is still room for improvement: syst. error can be reduced to 0.05 for 3 yrs of data ( $5.6\sigma \rightarrow 12\sigma$ )

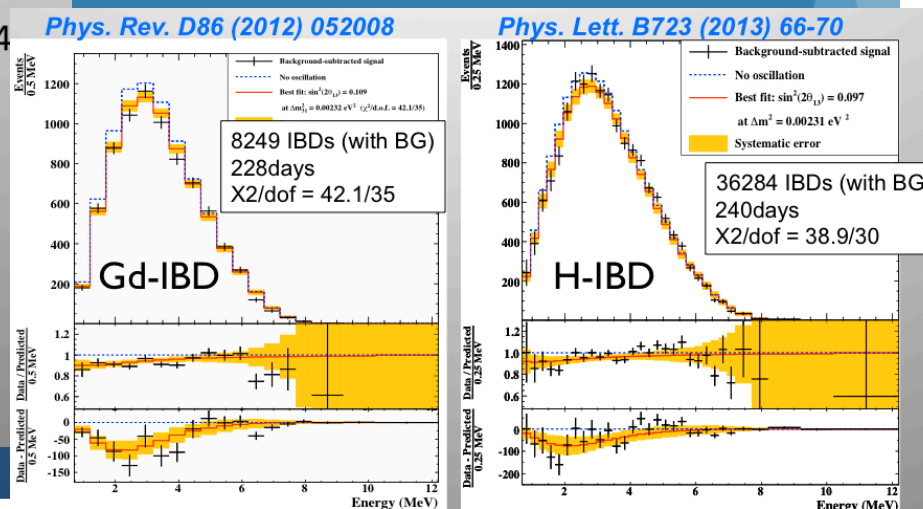
# Reactor experiments: Double Chooz

- North border of France
- Measurements utilizing only far detector (near detector in construction, starts taking data in 2014)



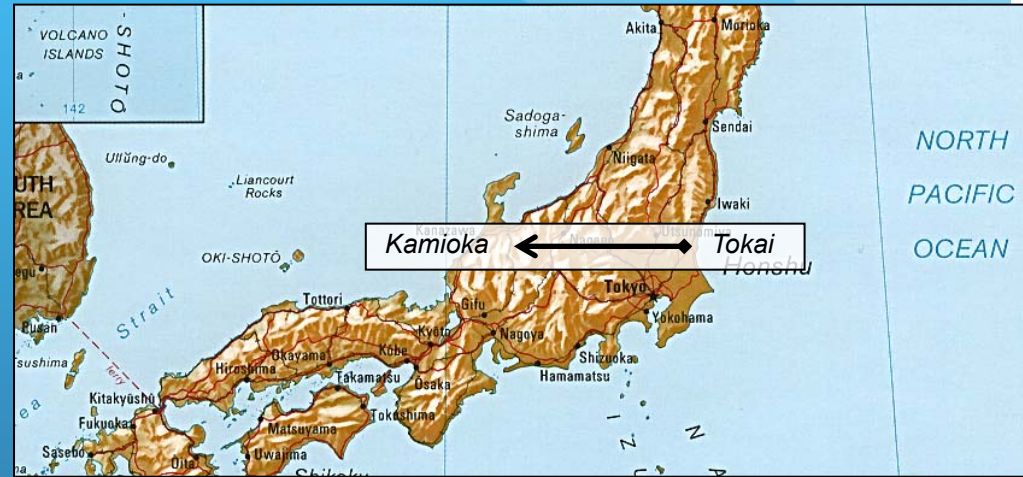
- Robust results with “independent” analyses
  - Gd-capture (rate+shape):  $\sin^2(2\theta_{13}) = 0.109 \pm 0.039$
  - H-capture (rate+shape):  $\sin^2(2\theta_{13}) = 0.097 \pm 0.048$
  - Nuclear power variation (Gd):  $\sin^2(2\theta_{13}) = 0.10 \pm 0.04$
  - Nuclear power variation (H):  $\sin^2(2\theta_{13}) = 0.13 \pm 0.07$
- New Combined fit results:
  - Rate and shape:  $\sin^2(2\theta_{13}) = 0.109 \pm 0.035$
  - Nuclear power variation:  $\sin^2(2\theta_{13}) = 0.097 \pm 0.035$
- Currently working on improved analysis (Gd+H)
- First results with near detector in 2014 (final precision  $\approx 10\%$ )

## Rate+shape analysis:



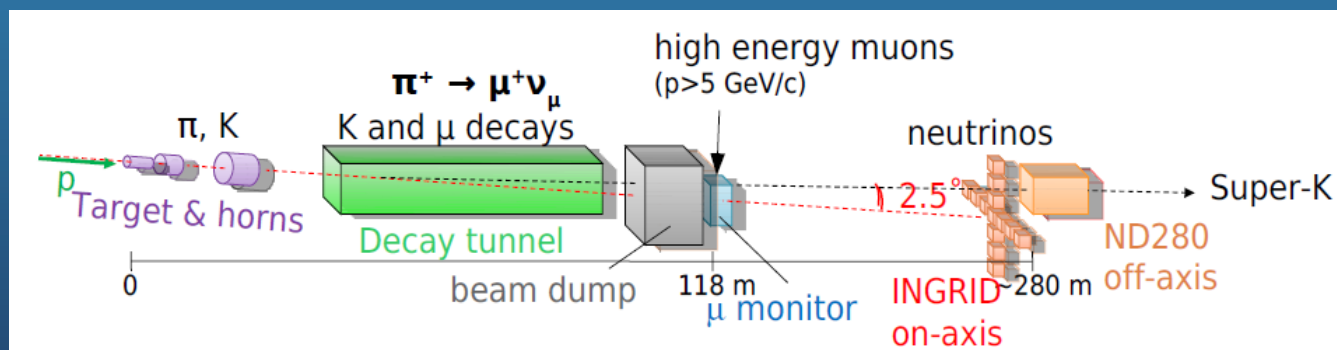
# Accelerator experiments: T2K

- ▶ **Tokai2Kamioka**: long baseline experiment with narrow-band beam
- ▶ Neutrinos produced in J-PARC laboratory in **Tokai** (30GeV proton beam hits a graphite target)
- ▶ **Near detector** 280m from the production point measures non-oscillated beam
- ▶ **Far detector** - Super-Kamiokande, large water Cherenkov detector in Kamioka mine studies effects of oscillations



**Main goal: neutrino oscillation studies**

- ▶ muon neutrino disappearance
- ▶ electron neutrino appearance



# The T2K Collaboration



*~500 members, 59 Institutes, 11 countries*

## Canada

TRIUMF  
U. Alberta  
U. B. Columbia  
U. Regina  
U. Toronto  
U. Victoria  
U. Winnipeg  
York U.

## France

CEA Saclay  
IPN Lyon  
LLR E. Poly.  
LPNHE Paris

## Germany

Aachen U.

## Italy

INFN, U. Bari  
INFN, U. Napoli  
INFN, U. Padova  
INFN, U. Roma

## Japan

ICRR Kamioka  
ICRR RCCN  
Kavli IPMU  
KEK  
Kobe U.  
Kyoto U.  
Miyagi U. Edu.  
Osaka City U.  
Okayama U.

Tokyo Metropolitan U.  
U. Tokyo

## Poland

IFJ PAN, Cracow  
NCBJ, Warsaw  
U. Silesia, Katowice  
U. Warsaw  
Warsaw U. T.  
Wrocław U.

## Russia

INR

## Spain

IFAE, Barcelona  
IFIC, Valencia

## Switzerland

ETH Zurich  
U. Bern  
U. Geneva

## United Kingdom

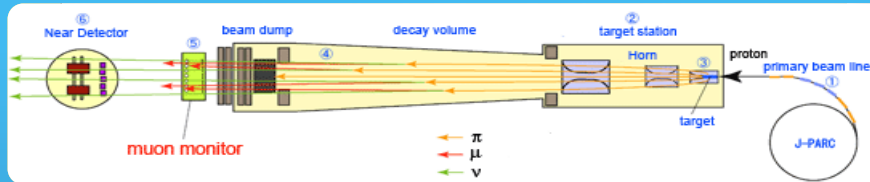
Imperial C. London  
Lancaster U.  
Oxford U.  
Queen Mary U. L.  
STFC/Daresbury  
STFC/RAL  
U. Liverpool

U. Sheffield  
U. Warwick

## USA

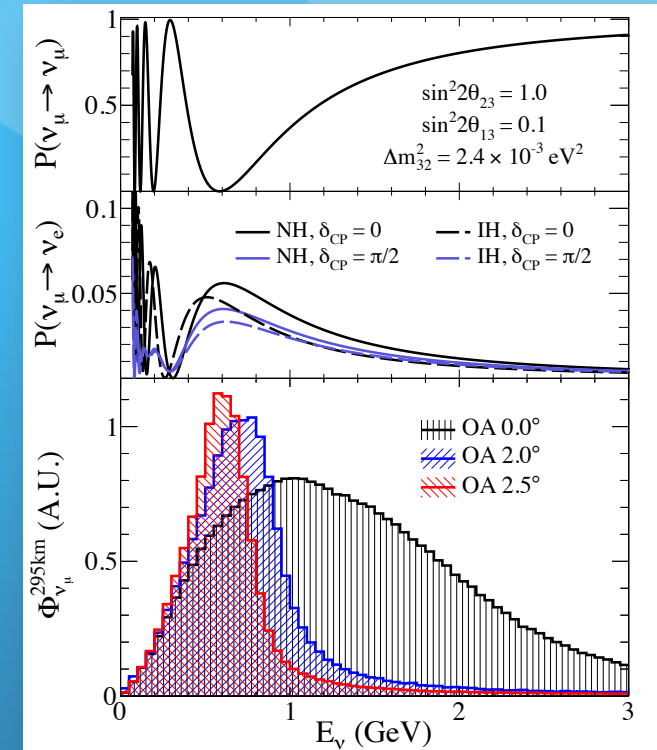
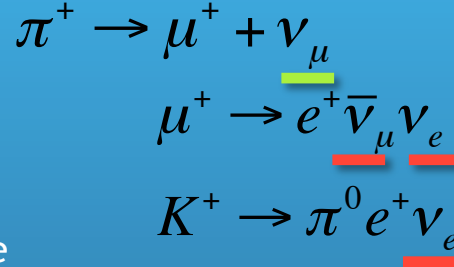
Boston U.  
Colorado S. U.  
Duke U.  
Louisiana S. U.  
Stony Brook U.  
U. C. Irvine  
U. Colorado  
U. Pittsburgh  
U. Rochester  
U. Washington

# T2K: beam

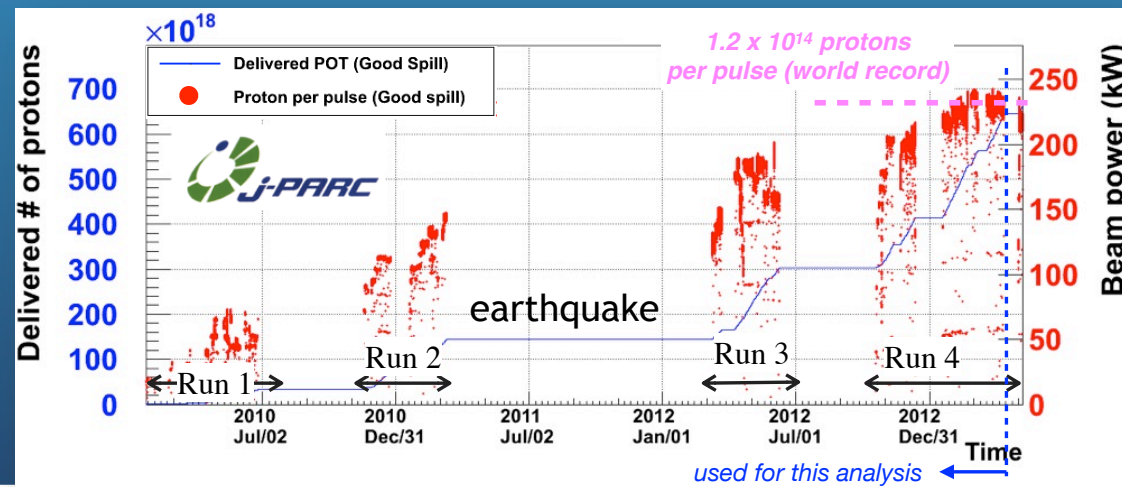


- Proton beam hits the target, produces hadrons, mainly pions
- The pions decay:  

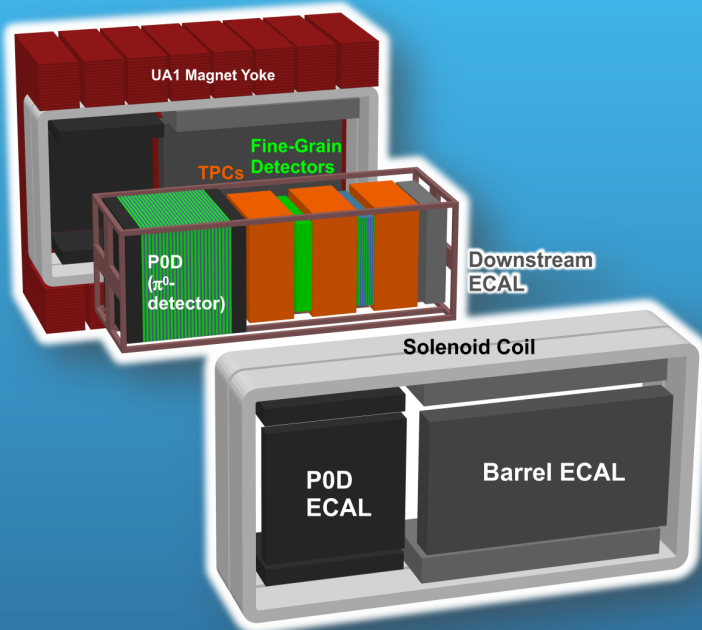
$$\pi^+ \rightarrow \mu^+ + \nu_\mu$$
- Beam contamination:
- Detectors positioned off-axis to get favorable spectrum shape
- Hadron production measured with target replica in NA61 experiment for better predictions of neutrino flux



Off-axis angle chosen: 2.5 degrees

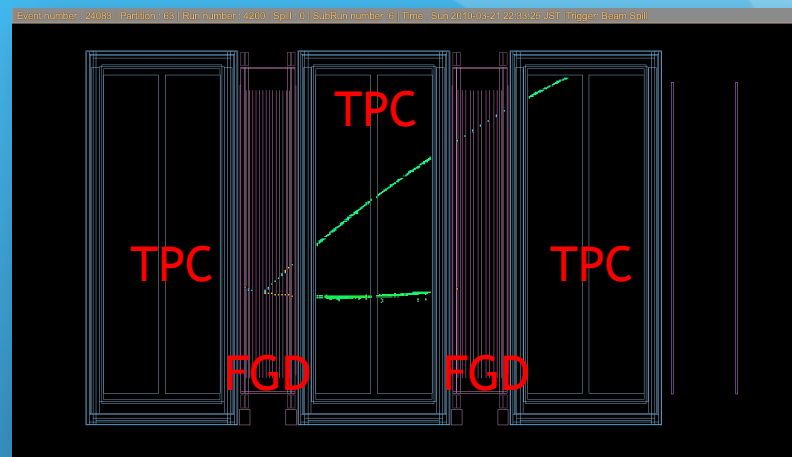


# T2K, near detector: ND280



## ND280 role:

- beam monitoring
- cross-section measurements
- Measurement of muon neutrino momentum for non-oscillated beam, electron neutrino contribution estimation



- FGD detectors- target for neutrino interactions, proton identification
- TPCs - identification and momentum reconstruction for muons, protons, charged pions and electrons

## ND280 role in neutrino oscillation analysis:

- Constrains flux estimates (external constraints from NA61 experiment)
  - Constrains cross-section parameters
- These are used to calculate MC for detector prediction

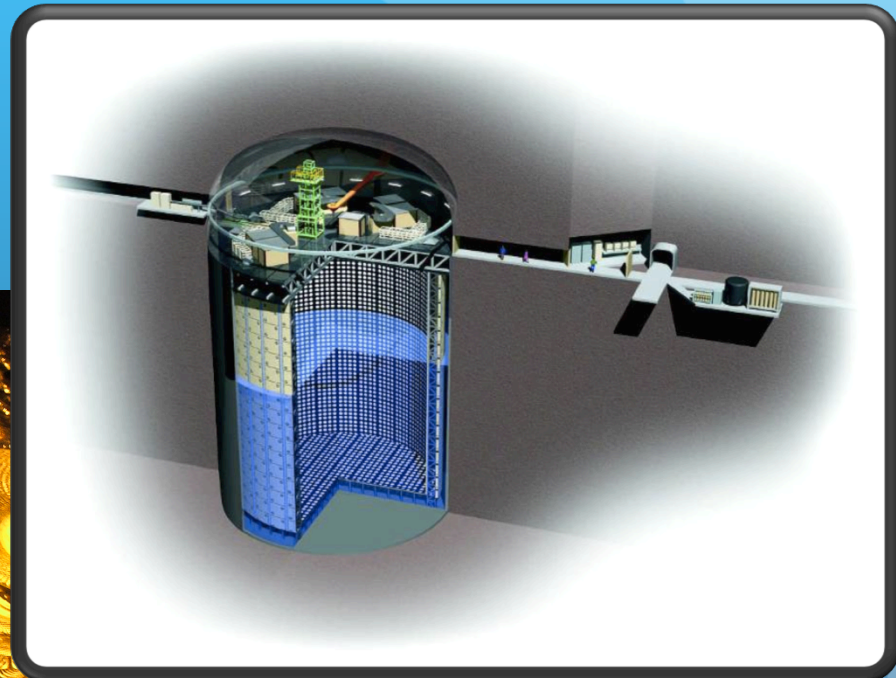
# T2K: Super-Kamiokande

## Large water Cherenkov detector

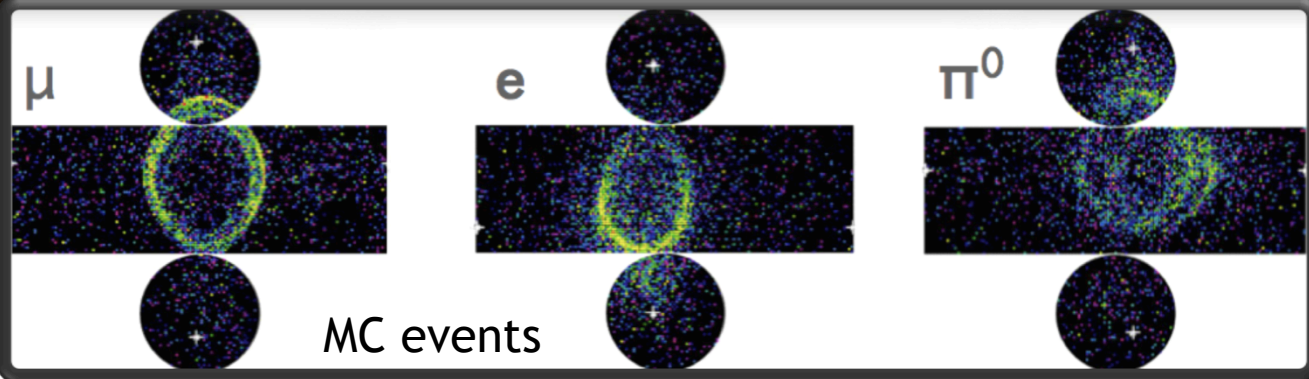
50 kton of water, 22.5 kton fiducial volume, >11,000 photomultipliers on the walls observe Cherenkov light

Many years of experience, very well known detection technique, systematic errors known and understood

Studies also atmospheric, solar neutrinos

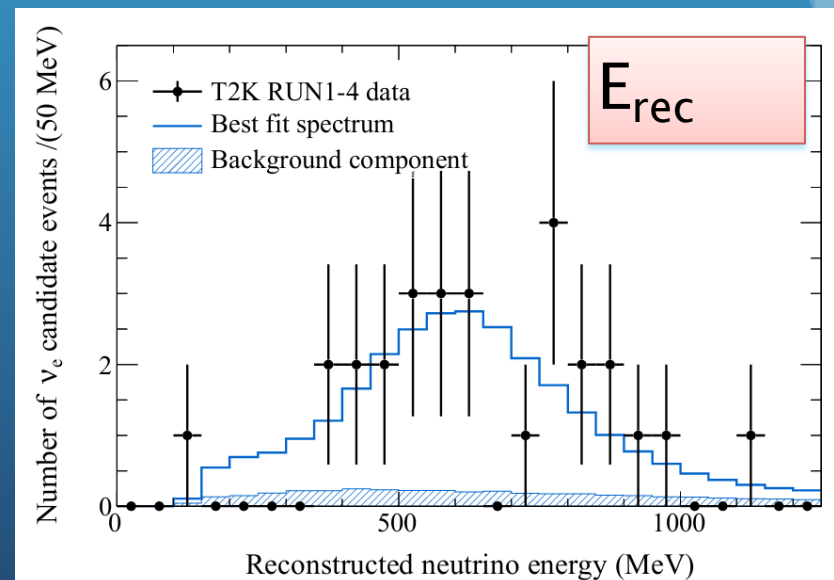
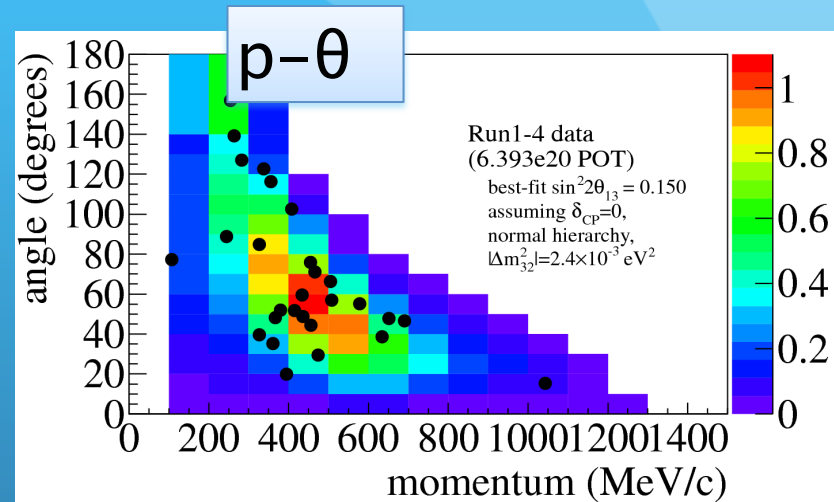


Good muon/electron separation



# T2K: 2013 results

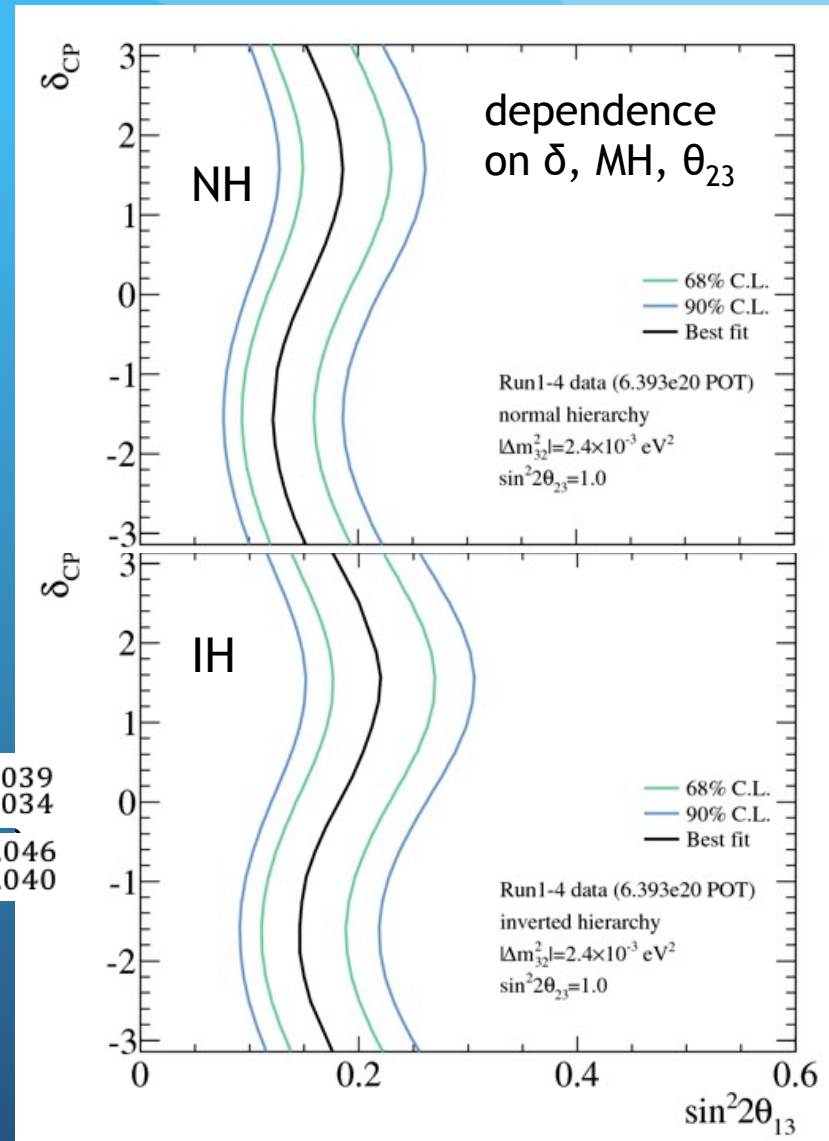
- Neutrino oscillation parameters extracted in two ways:
  - using reconstructed neutrino energy distribution
  - using observed electron momentum and angle
- **28 events observed**
  - for  $\sin^2 2\theta_{13}=0.1$ ,  $\sin^2 2\theta_{23}=1$ ,  $\delta=0$  we would see  $20.4 \pm 1.8$  events
  - expected background  $4.64 \pm 0.53$



# T2K: 2013 results

- Electron neutrino appearance result - 2012 update
  - new data (run 4)
  - new SK reconstruction algorithm (fiTQun), better  $\pi^0$  background rejection
  - near detector CC inclusive measurement improved by using new event categories
- Best fit results for  $\delta=0$  (68% C.L. error)
  - normal hierarchy  $\sin^2 2\theta_{13} = 0.150^{+0.039}_{-0.034}$
  - inverted hierarchy  $\sin^2 2\theta_{13} = 0.182^{+0.046}_{-0.040}$
- $7.5\sigma$  significance for non-zero  $\theta_{13}$

First ever observation ( $>5\sigma$ ) of an explicit  $\nu$  appearance channel!



NOTE: These are 1D contours for various value of  $\delta_{CP}$ , not 2D contours

# T2K: $\theta_{23}$ uncertainty and reactor results

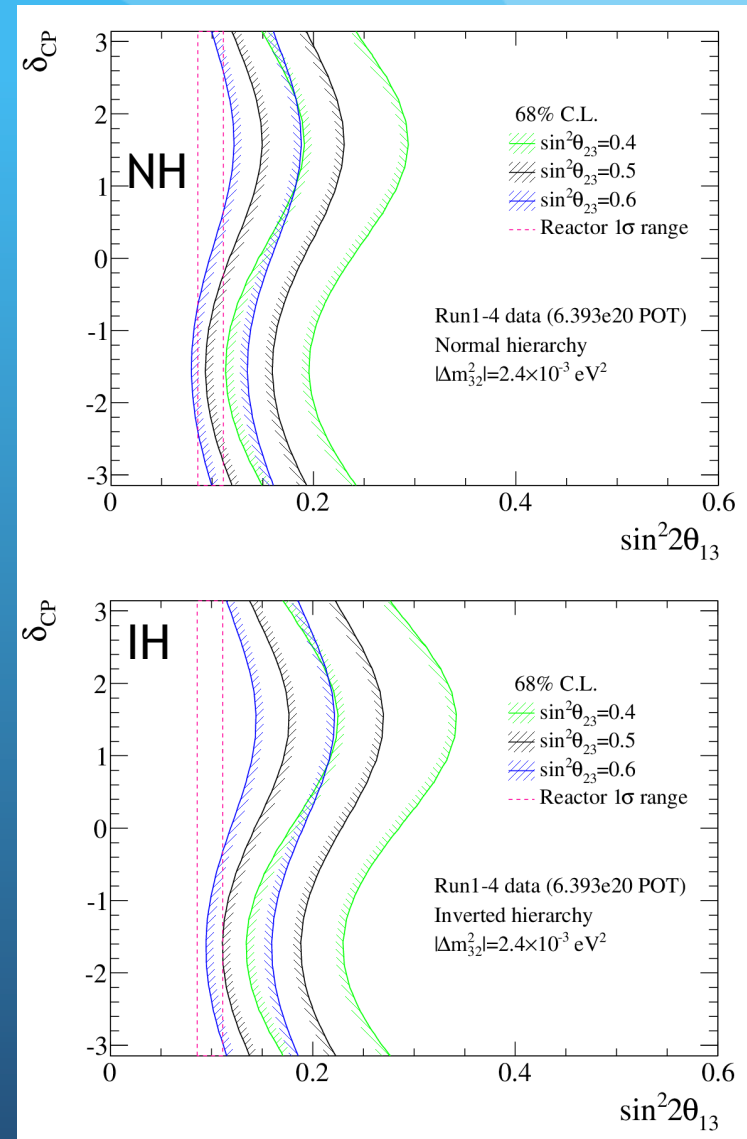
- $\nu_e$  appearance contours depend on the value of  $\theta_{23}$
- This needs to be determined more precisely by studying  $\nu_\mu$  disappearance

$$P(\nu_\mu \rightarrow \nu_e) \approx \sin^2 2\theta_{13} \sin^2 \theta_{23} \sin^2 (1.27 \Delta m_{23}^2 L / E)$$

$$P(\nu_\mu \rightarrow \nu_\mu) \approx 1 - \sin^2 2\theta_{23} \cos^4 \theta_{13} \sin^2 (1.27 \Delta m_{23}^2 L / E)$$

Comparison with reactor results (PDG2012):  
 $\sin^2 2\theta_{13} = 0.098 \pm 0.013$

Leading terms

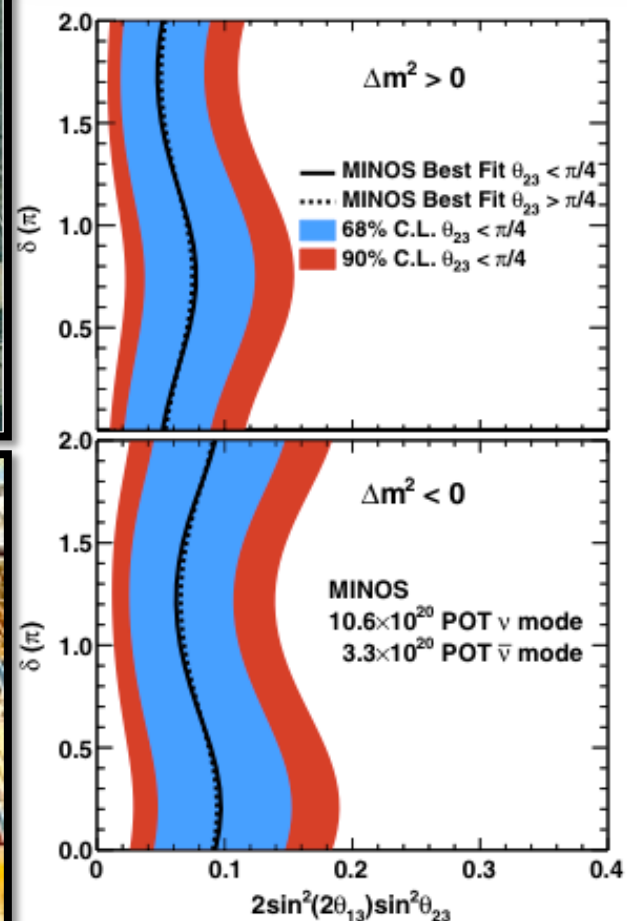
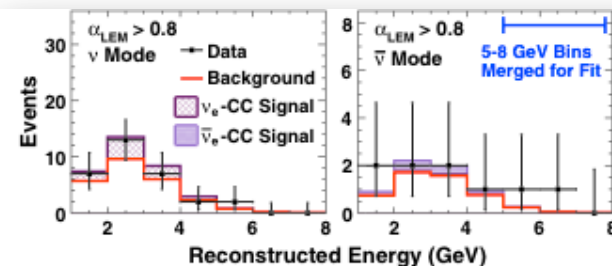
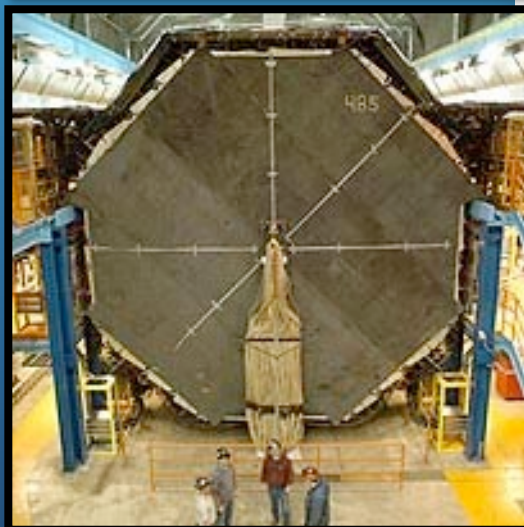


NOTE: These are 1D contours for various value of  $\delta_{CP}$ , not 2D contours

# Accelerator expts: MINOS

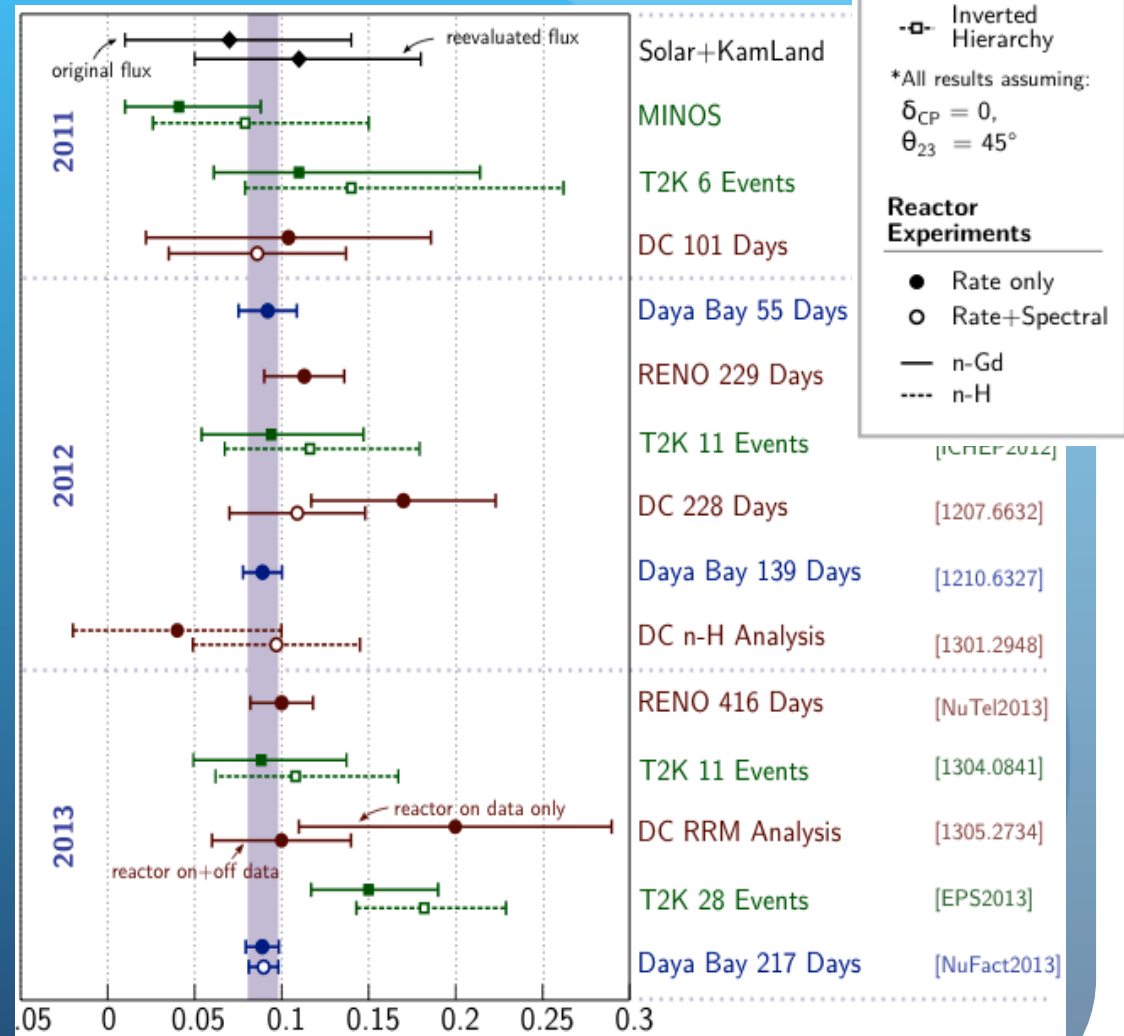
- LB (735km) experiment in USA, started in 2005
- NuMI beam from FermiLab - mainly muon (anti)neutrinos
- Near detector in FermiLab, far detector in Soudan mine (3.8kT fiducial) - magnetized tracking calorimeters
- Neutrino ( $10.6 \cdot 10^{20}$  POT) and antineutrino ( $3.3 \cdot 10^{20}$  POT) beam data collected
- Result:

$$2\sin^2(2\theta_{13})\sin^2(\theta_{23}) = 0.093^{+0.054}_{-0.049}$$



# Summary

- $\theta_{13}$  mixing angle has been successfully measured in reactor and long-baseline experiments
- more precise measurements coming soon
- we can now study CP violation effects



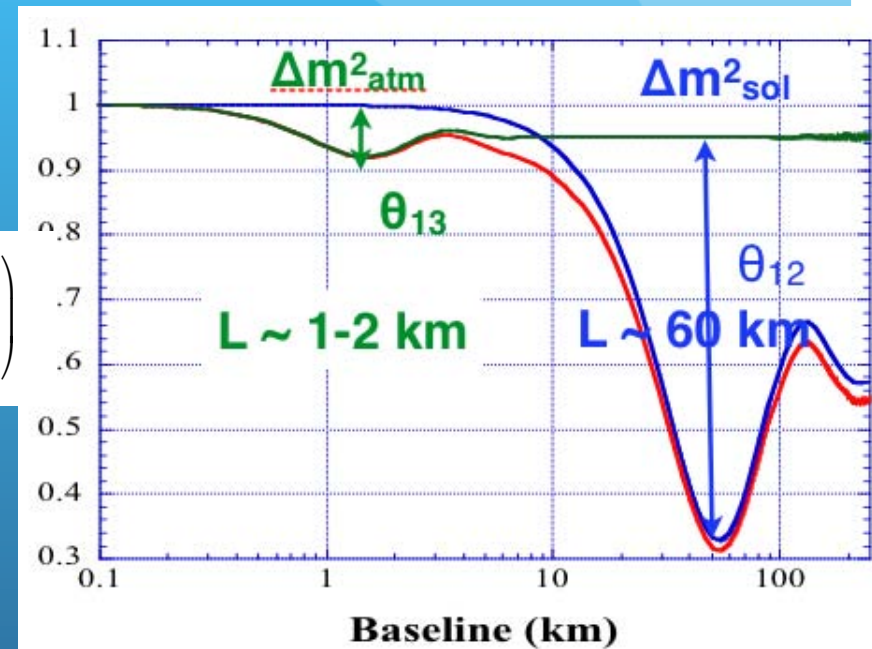
Plot taken from Soeren Jetter's Daya Bay talk (NuFact 2013)

backup

# Reactor experiments

- probability for electron antineutrino disappearance

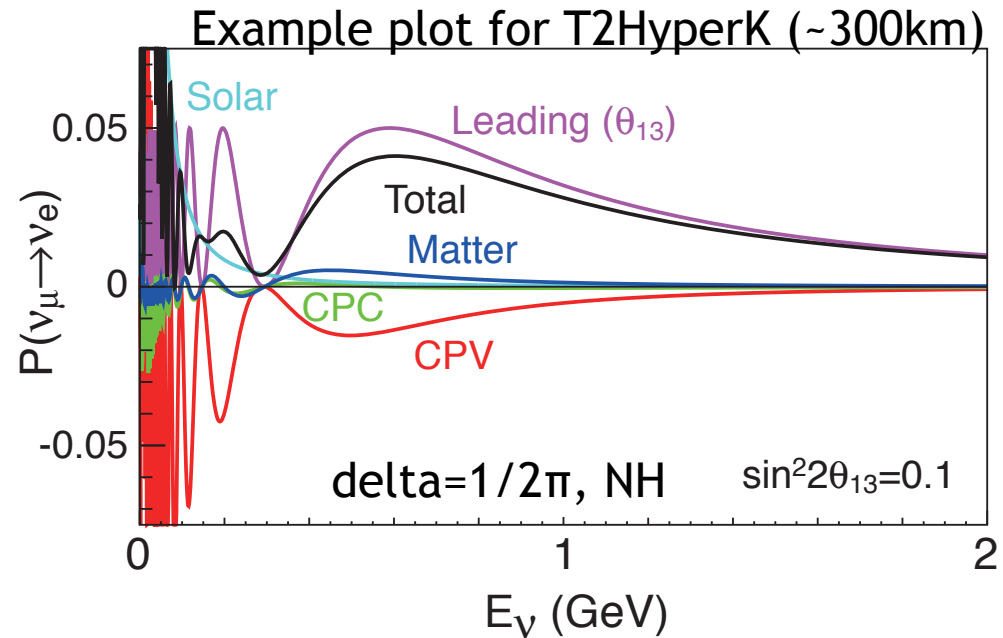
$$P_{ee} \approx 1 - \sin^2 2\theta_{13} \sin^2 \left( \frac{\Delta m_{31}^2 L}{4E_\nu} \right) - \cos^4 \theta_{13} \sin^2 2\theta_{12} \sin^2 \left( \frac{\Delta m_{21}^2 L}{4E_\nu} \right)$$



$$E_{\bar{\nu}} \cong \underbrace{T_{e^+}}_{10-40 \text{ keV}} + T_n + \underbrace{(M_n - M_p)}_{1.8 \text{ MeV: threshold}} + m_{e^+}$$

# Long baseline experiments

- Probability of electron neutrino appearance



$$P(\nu_\mu \rightarrow \nu_e) = 4C_{13}^2 S_{13}^2 S_{23}^2 \cdot \sin^2 \Delta_{31} \quad \text{leading term}$$

CP conserving

$$+ 8C_{13}^2 S_{12} S_{13} S_{23} (C_{12} C_{23} \cos \delta - S_{12} S_{13} S_{23}) \cdot \cos \Delta_{32} \cdot \sin \Delta_{31} \cdot \sin \Delta_{21}$$

$$- 8C_{13}^2 C_{12} C_{23} S_{12} S_{13} S_{23} \sin \delta \cdot \sin \Delta_{32} \cdot \sin \Delta_{31} \cdot \sin \Delta_{21}$$

CP violating

$$+ 4S_{12}^2 C_{13}^2 (C_{12}^2 C_{23}^2 + S_{12}^2 S_{23}^2 S_{13}^2 - 2C_{12} C_{23} S_{12} S_{23} S_{13} \cos \delta) \cdot \sin^2 \Delta_{21}$$

$$- 8C_{13}^2 S_{13}^2 S_{23}^2 \cdot \frac{aL}{4E_\nu} (1 - 2S_{13}^2) \cdot \cos \Delta_{32} \cdot \sin \Delta_{31}$$

solar term

$$+ 8C_{13}^2 S_{13}^2 S_{23}^2 \frac{a}{\Delta m_{31}^2} (1 - 2S_{13}^2) \cdot \sin^2 \Delta_{31},$$

matter effects

$C_{ij}, S_{ij}, \Delta_{ij}$

for  $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$   $\delta \rightarrow -\delta$   $a \rightarrow -a$ .

$\cos \theta_{ij}, \sin \theta_{ij}, \Delta m_{ij}^2 L / 4E_\nu$

$\alpha \sim \rho \cdot E_\nu$

# Oscillation parameter values

$$\Delta m_{31}^2 = \begin{cases} 2.53_{-0.10}^{+0.08} \\ -(2.40_{-0.07}^{+0.10}) \end{cases} \times 10^{-3} \text{ eV}^2$$
$$\sin^2 \theta_{23} = \begin{cases} 0.49_{-0.05}^{+0.08} & \text{for } \Delta m_{31}^2 > 0 \\ 0.53_{-0.07}^{+0.05} & \text{for } \Delta m_{31}^2 < 0 \end{cases}$$

$$\sin^2 \theta_{12} = 0.320_{-0.017}^{+0.015}$$

$$\Delta m_{21}^2 = (7.62 \pm 0.19) \times 10^{-5} \text{ eV}^2$$