

# Searches for physics/particles beyond the Standard Model at the LHC

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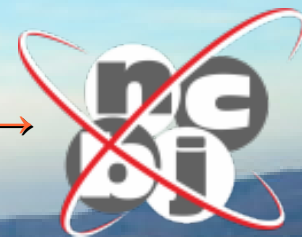
on behalf of the CMS and the ATLAS Collaborations

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## OUTLINE

- BSM searches @ LHC (introduction)
- SUSY
  - mainstream, variables, approaches, summary plots, SMS examples
- Exotica
  - mainstream, summary plots, reinterpretations (LLP examples)
  - The first look at 13 TeV data!
- DM @ LHC
- Conclusions

## *Matter To The Deepest*

*Recent Developments In Physics Of Fundamental Interactions  
XXXIX International Conference of Theoretical Physics*

**Ustroń, 15/09/2015**

# Searches for BSM @ LHC (an introduction)



At the LHC BSM searches are organized into two big subgroups: SUSY and Exotic(a). The later is a short name for all non - (mainstream) SUSY.

In the CMS, Exotica was divided further to establish Beyond 2 Generation (B2G) subgroup.

What is needed to perform searches for direct BSM phenomena?

- Well performing collider + efficient & precise multi-purpose detectors (ATLAS & CMS).
- A choice of a search topology (several topologies for one phenomenon are possible and *vice versa*: several phenomena could be searched for in a given topology).
- Trigger: high level HLT (software) but not forgetting level one L1 (instrumental).
- Physics objects (jet,  $E_T^{\text{miss}}$ , leptons, photon): definition, performance, validation, time stability *etc.*

It is done for all analyses, but some searches need custom objects (long-lived, monopoles ...).

- Preselection: to assure homogeneous trigger efficiency (and suppress background).
- A choice of signal sensitive variables (at least two).

Definition of one (or several) Signal enriched Regions SR (both criteria satisfied).

What rest form Control Regions CR in which at least one criterion is not satisfied.

Use data driven methods to find out and validate transfer factors from CR to SR to obtain a data driven estimate of background level in the SR (including background systematics – not explained here), however, while data driven methods are preferred,

sometimes it is necessary to use MC to estimate the level of background.

It could be more complicated if "shape analysis" is used instead of "counting experiment".

- Claim discovery if statistically significant excess is found or, if not, interpret the result in term of constraints on selected models.

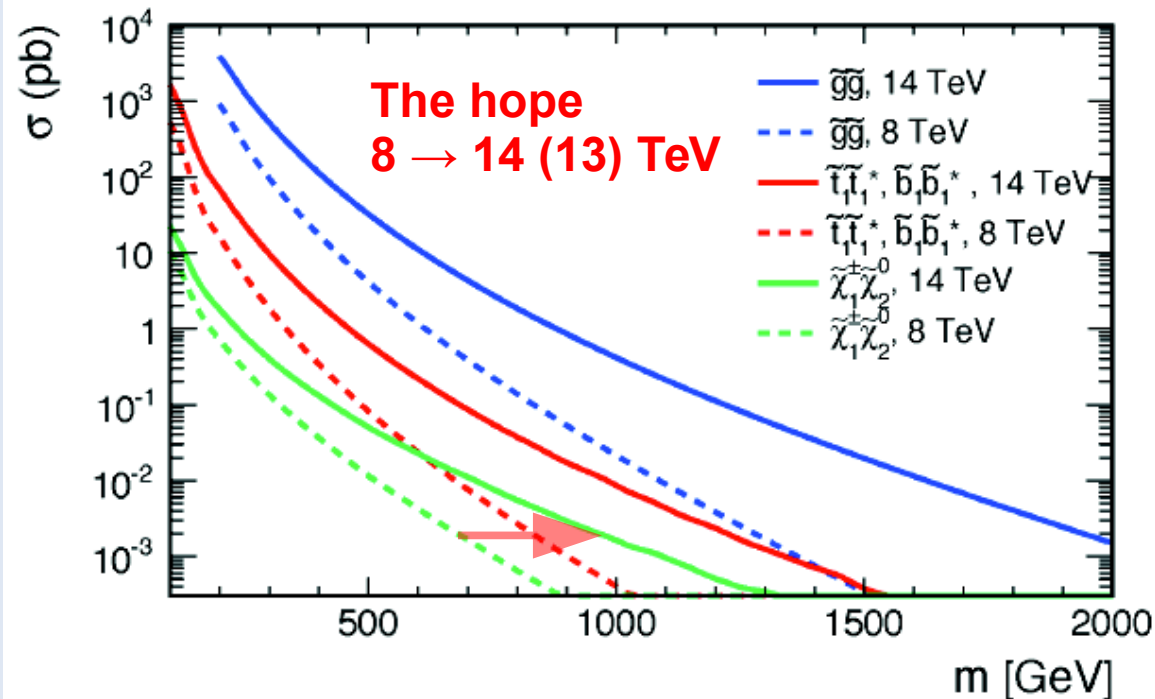
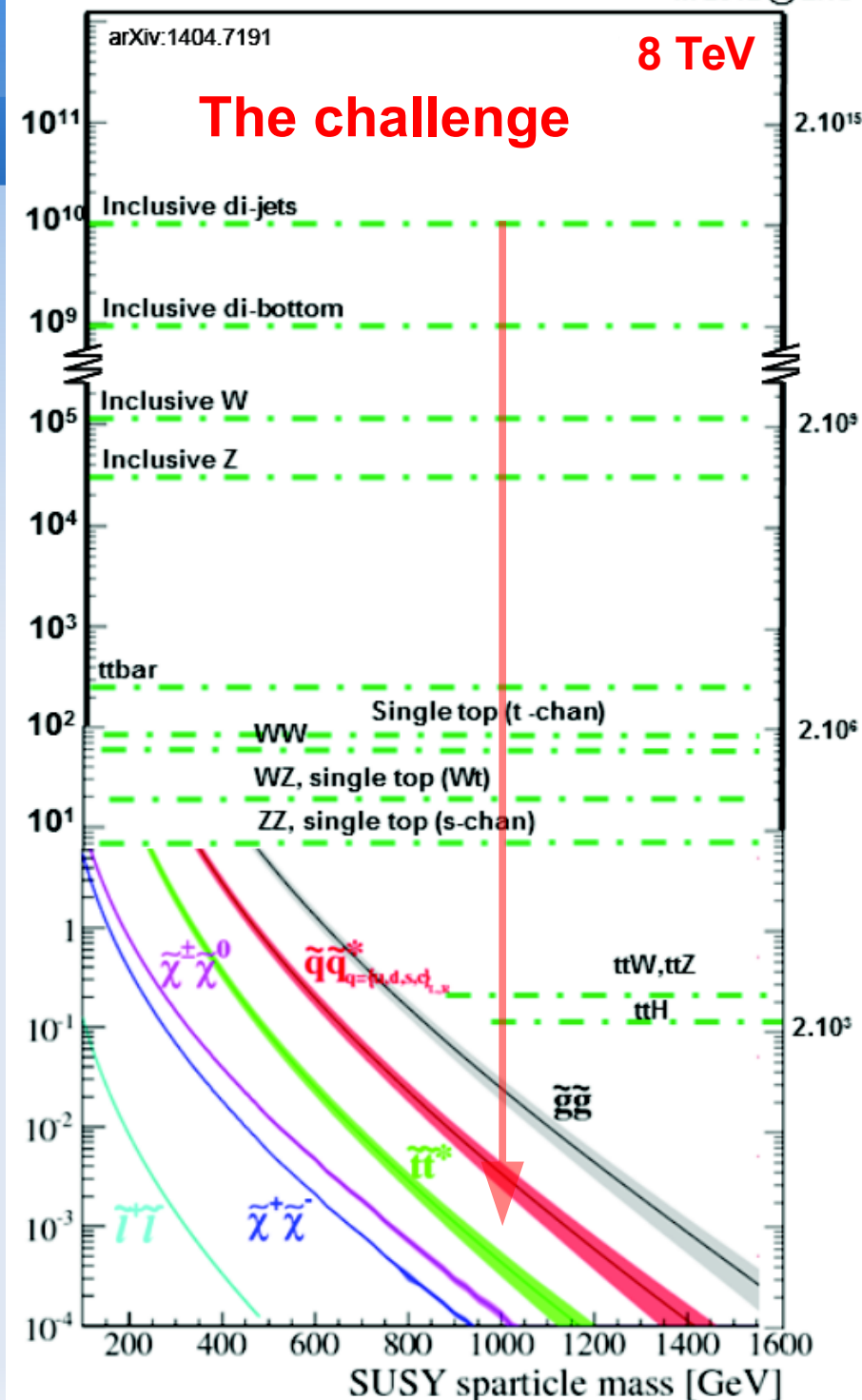
# SUSY

SUSY is the benchmark for general BSM searches  
(many SUSY searches could be reinterpreted in other models)

From inclusive searches for missing transverse energy  $E_T^{\text{miss}}$   
and given number of jets, leptons (SS, OS *etc.*), photons

*via* targeted searches for specific scenarios

to exploration of challenging (eg. compressed spectra)  
or "less standard" (RPV, non-prompt decays *etc.*) ones.



# Searching for SUSY



What makes our (experimentalists) life hard?

- Triggering on soft signals
- Background estimation (using data driven methods whenever possible)
- Background systematics (which limits sensitivity for challenging scenarios)

What help us?

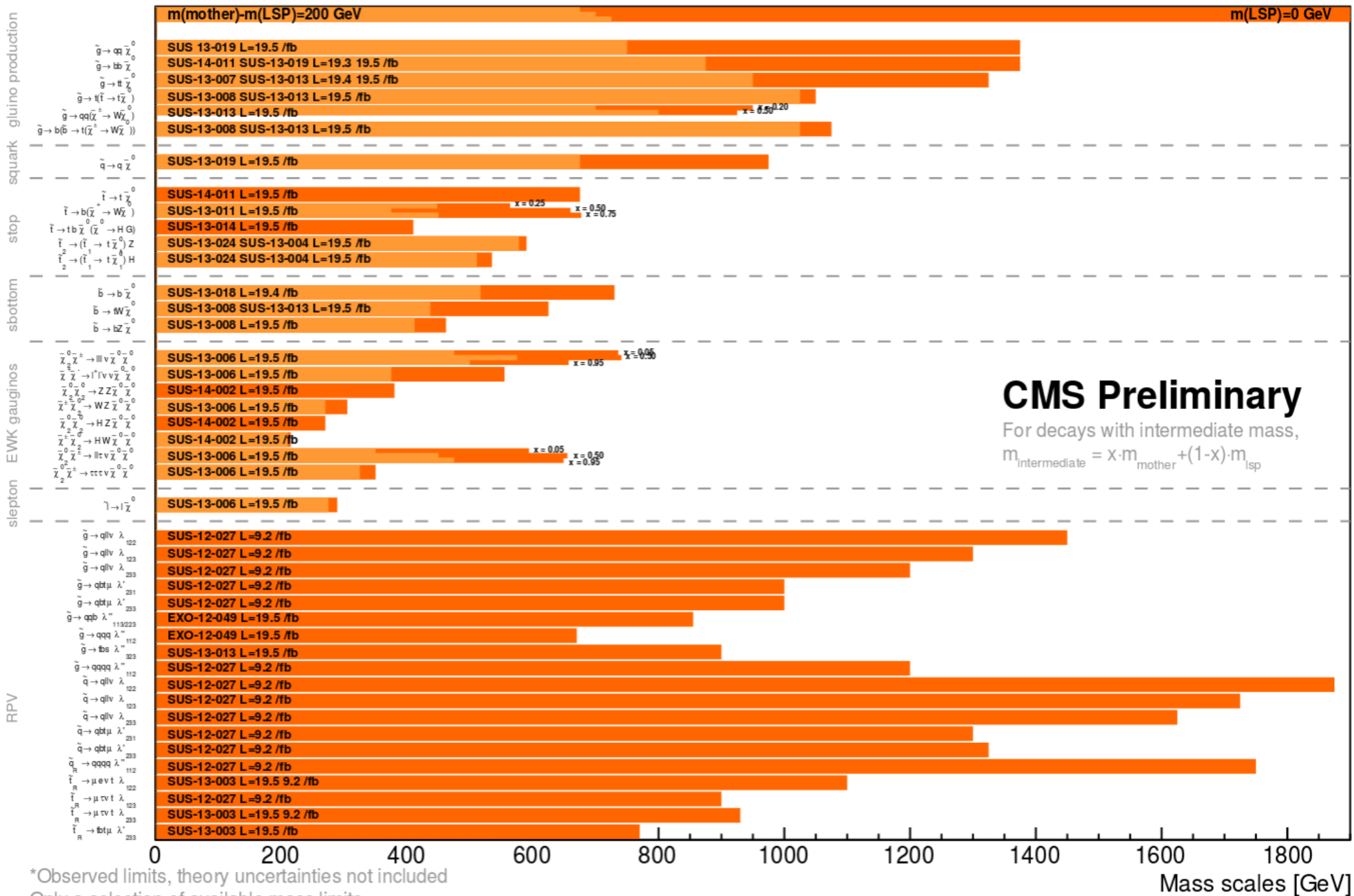
- Growing set of kinematical variables  $E_T^{\text{miss}}$ ,  $H_T$ ,  $S_T$  ( $\rightarrow S_T^{\text{lept.}}$ ), razor ( $\rightarrow$  super-razor) *etc.*
- Multivariate analyses
- New models/signatures, analysis methods, background suppression ideas, *etc.*

What guide us?

- Searches are inspired and/or interpreted
  - $\rightarrow$  in specific models (CMSSM, NUHM, NMSSM *etc.*)
  - $\rightarrow$  phenomenological pMSSM
  - $\rightarrow$  simplified models SMS
- Our desire for discovery of BSM physics



## Summary of CMS SUSY Results\* in SMS framework



\*Observed limits, theory uncertainties not included

Only a selection of available mass limits

Probe \*up to\* the quoted mass limit

Model		$e, \mu, \tau, \gamma$	Jets	$E_T^{\text{miss}}$	$\int \mathcal{L} dt [\text{fb}^{-1}]$	Mass limit		$\sqrt{s} = 7 \text{ TeV}$	$\sqrt{s} = 8 \text{ TeV}$	Reference
Inclusive Searches	MSUGRA/CMSSM	0-3 $e, \mu/1\text{-}2 \tau$	2-10 jets/3 $b$	Yes	20.3	$\tilde{q}, \tilde{g}$	1.8 TeV		$m(\tilde{q})=m(\tilde{g})$	1507.05525
	$\tilde{q}\tilde{q}, \tilde{q} \rightarrow q\tilde{\chi}_1^0$	0	2-6 jets	Yes	20.3	$\tilde{q}$	850 GeV		$m(\tilde{\chi}_1^0)=0 \text{ GeV}, m(1^{\text{st}} \text{ gen. } \tilde{q})=m(2^{\text{nd}} \text{ gen. } \tilde{q})$	1405.7875
	$\tilde{q}\tilde{q}, \tilde{q} \rightarrow q\tilde{\chi}_1^0$ (compressed)	mono-jet	1-3 jets	Yes	20.3	$\tilde{q}$	100-440 GeV		$m(\tilde{q})-m(\tilde{\chi}_1^0)<10 \text{ GeV}$	1507.05525
	$\tilde{q}\tilde{q}, \tilde{q} \rightarrow q(\ell\ell/\ell\nu/\nu\nu)\tilde{\chi}_1^0$	2 $e, \mu$ (off-Z)	2 jets	Yes	20.3	$\tilde{q}$	780 GeV		$m(\tilde{\chi}_1^0)=0 \text{ GeV}$	1503.03290
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{q}\tilde{\chi}_1^0$	0	2-6 jets	Yes	20.3	$\tilde{g}$	1.33 TeV		$m(\tilde{\chi}_1^0)=0 \text{ GeV}$	1405.7875
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{q}\tilde{\chi}_1^\pm \rightarrow q\tilde{q}W^\pm\tilde{\chi}_1^0$	0-1 $e, \mu$	2-6 jets	Yes	20	$\tilde{g}$	1.26 TeV		$m(\tilde{\chi}_1^0)<300 \text{ GeV}, m(\tilde{\chi}^\pm)=0.5(m(\tilde{\chi}_1^0)+m(\tilde{g}))$	1507.05525
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{q}(\ell\ell/\ell\nu/\nu\nu)\tilde{\chi}_1^0$	2 $e, \mu$	0-3 jets	-	20	$\tilde{g}$	1.32 TeV		$m(\tilde{\chi}_1^0)=0 \text{ GeV}$	1501.03555
	GMSB ( $\tilde{\ell}$ NLSP)	1-2 $\tau$ + 0-1 $\ell$	0-2 jets	Yes	20.3	$\tilde{g}$	1.6 TeV		$\tan\beta > 20$	1407.0603
	GGM (bino NLSP)	2 $\gamma$	-	Yes	20.3	$\tilde{g}$	1.29 TeV		$c\tau(\text{NLSP})<0.1 \text{ mm}$	1507.05493
	GGM (higgsino-bino NLSP)	$\gamma$	1 $b$	Yes	20.3	$\tilde{g}$	1.3 TeV		$m(\tilde{\chi}_1^0)<900 \text{ GeV}, c\tau(\text{NLSP})<0.1 \text{ mm}, \mu<0$	1507.05493
	GGM (higgsino-bino NLSP)	$\gamma$	2 jets	Yes	20.3	$\tilde{g}$	1.25 TeV		$m(\tilde{\chi}_1^0)<850 \text{ GeV}, c\tau(\text{NLSP})<0.1 \text{ mm}, \mu>0$	1507.05493
	GGM (higgsino NLSP)	2 $e, \mu$ (Z)	2 jets	Yes	20.3	$\tilde{g}$	850 GeV		$m(\text{NLSP})>430 \text{ GeV}$	1503.03290
Gravitino LSP	0	mono-jet	Yes	20.3	$\tilde{G}^{1/2}$ scale	865 GeV		$m(\tilde{G})>1.8 \times 10^{-4} \text{ eV}, m(\tilde{g})=m(\tilde{q})=1.5 \text{ TeV}$	1502.01518	
3 <sup>rd</sup> gen. $\tilde{g}$ med.	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow b\tilde{b}\tilde{\chi}_1^0$	0	3 $b$	Yes	20.1	$\tilde{g}$	1.25 TeV		$m(\tilde{\chi}_1^0)<400 \text{ GeV}$	1407.0600
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow t\tilde{t}\tilde{\chi}_1^0$	0	7-10 jets	Yes	20.3	$\tilde{g}$	1.1 TeV		$m(\tilde{\chi}_1^0)<350 \text{ GeV}$	1308.1841
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow t\tilde{t}\tilde{\chi}_1^0$	0-1 $e, \mu$	3 $b$	Yes	20.1	$\tilde{g}$	1.34 TeV		$m(\tilde{\chi}_1^0)<400 \text{ GeV}$	1407.0600
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow b\tilde{t}\tilde{\chi}_1^\pm$	0-1 $e, \mu$	3 $b$	Yes	20.1	$\tilde{g}$	1.3 TeV		$m(\tilde{\chi}_1^0)<300 \text{ GeV}$	1407.0600
3 <sup>rd</sup> gen. squarks direct production	$\tilde{b}_1\tilde{b}_1, \tilde{b}_1 \rightarrow b\tilde{\chi}_1^0$	0	2 $b$	Yes	20.1	$\tilde{b}_1$	100-620 GeV		$m(\tilde{\chi}_1^0)<90 \text{ GeV}$	1308.2631
	$\tilde{b}_1\tilde{b}_1, \tilde{b}_1 \rightarrow t\tilde{\chi}_1^\pm$	2 $e, \mu$ (SS)	0-3 $b$	Yes	20.3	$\tilde{b}_1$	275-440 GeV		$m(\tilde{\chi}_1^\pm)=2 m(\tilde{\chi}_1^0)$	1404.2500
	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow b\tilde{\chi}_1^\pm$	1-2 $e, \mu$	1-2 $b$	Yes	4.7/20.3	$\tilde{t}_1$	110-167 GeV		$m(\tilde{\chi}_1^\pm)=2 m(\tilde{\chi}_1^0), m(\tilde{\chi}_1^0)=55 \text{ GeV}$	1209.2102, 1407.0583
	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow Wb\tilde{\chi}_1^0$ or $t\tilde{\chi}_1^0$	0-2 $e, \mu$	0-2 jets/1-2 $b$	Yes	20.3	$\tilde{t}_1$	90-191 GeV		$m(\tilde{\chi}_1^0)=1 \text{ GeV}$	1506.08616
	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow c\tilde{\chi}_1^0$	0	mono-jet/ $c$ -tag	Yes	20.3	$\tilde{t}_1$	90-240 GeV		$m(\tilde{t}_1)-m(\tilde{\chi}_1^0)<85 \text{ GeV}$	1407.0608
	$\tilde{t}_1\tilde{t}_1$ (natural GMSB)	2 $e, \mu$ (Z)	1 $b$	Yes	20.3	$\tilde{t}_1$	150-580 GeV		$m(\tilde{\chi}_1^0)>150 \text{ GeV}$	1403.5222
	$\tilde{t}_2\tilde{t}_2, \tilde{t}_2 \rightarrow \tilde{t}_1 + Z$	3 $e, \mu$ (Z)	1 $b$	Yes	20.3	$\tilde{t}_2$	290-600 GeV		$m(\tilde{\chi}_1^0)<200 \text{ GeV}$	1403.5222
EW direct	$\tilde{\ell}_{L,R}\tilde{\ell}_{L,R}, \tilde{\ell} \rightarrow \ell\tilde{\chi}_1^0$	2 $e, \mu$	0	Yes	20.3	$\tilde{\ell}$	90-325 GeV		$m(\tilde{\chi}_1^0)=0 \text{ GeV}$	1403.5294
	$\tilde{\chi}_1^\pm\tilde{\chi}_1^\pm, \tilde{\chi}_1^\pm \rightarrow \tilde{\ell}\nu(\ell\bar{\nu})$	2 $e, \mu$	0	Yes	20.3	$\tilde{\chi}_1^\pm$	140-465 GeV		$m(\tilde{\chi}_1^0)=0 \text{ GeV}, m(\tilde{\ell}, \bar{\nu})=0.5(m(\tilde{\chi}_1^\pm)+m(\tilde{\chi}_1^0))$	1403.5294
	$\tilde{\chi}_1^\pm\tilde{\chi}_1^\pm, \tilde{\chi}_1^\pm \rightarrow \tilde{\tau}\nu(\tau\bar{\nu})$	2 $\tau$	-	Yes	20.3	$\tilde{\chi}_1^\pm$	100-350 GeV		$m(\tilde{\chi}_1^0)=0 \text{ GeV}, m(\tilde{\ell}, \bar{\nu})=0.5(m(\tilde{\chi}_1^\pm)+m(\tilde{\chi}_1^0))$	1407.0350
	$\tilde{\chi}_1^\pm\tilde{\chi}_2^0 \rightarrow \tilde{\ell}_L\nu\tilde{\ell}_L\ell(\bar{\nu}\nu), \ell\bar{\nu}\tilde{\ell}_L\ell(\bar{\nu}\nu)$	3 $e, \mu$	0	Yes	20.3	$\tilde{\chi}_1^\pm, \tilde{\chi}_2^0$	700 GeV		$m(\tilde{\chi}_1^\pm)=m(\tilde{\chi}_2^0), m(\tilde{\chi}_1^0)=0, m(\tilde{\ell}, \bar{\nu})=0.5(m(\tilde{\chi}_1^\pm)+m(\tilde{\chi}_1^0))$	1402.7029
	$\tilde{\chi}_1^\pm\tilde{\chi}_2^0 \rightarrow W\tilde{\chi}_1^0 Z\tilde{\chi}_1^0$	2-3 $e, \mu$	0-2 jets	Yes	20.3	$\tilde{\chi}_1^\pm, \tilde{\chi}_2^0$	420 GeV		$m(\tilde{\chi}_1^\pm)=m(\tilde{\chi}_2^0), m(\tilde{\chi}_1^0)=0$ , sleptons decoupled	1403.5294, 1402.7029
	$\tilde{\chi}_1^\pm\tilde{\chi}_2^0 \rightarrow W\tilde{\chi}_1^0 h\tilde{\chi}_1^0, h \rightarrow b\tilde{b}/WW/\tau\tau/\gamma\gamma$	$e, \mu, \gamma$	0-2 $b$	Yes	20.3	$\tilde{\chi}_1^\pm, \tilde{\chi}_2^0$	250 GeV		$m(\tilde{\chi}_1^\pm)=m(\tilde{\chi}_2^0), m(\tilde{\chi}_1^0)=0$ , sleptons decoupled	1501.07110
	$\tilde{\chi}_2^0\tilde{\chi}_3^0 \rightarrow \tilde{\ell}_R\ell$	4 $e, \mu$	0	Yes	20.3	$\tilde{\chi}_2^0, \tilde{\chi}_3^0$	620 GeV		$m(\tilde{\chi}_2^0)=m(\tilde{\chi}_3^0), m(\tilde{\chi}_1^0)=0, m(\tilde{\ell}, \bar{\nu})=0.5(m(\tilde{\chi}_2^0)+m(\tilde{\chi}_1^0))$	1405.5086
	GGM (wino NLSP) weak prod.	1 $e, \mu + \gamma$	-	Yes	20.3	$\tilde{W}$	124-361 GeV		$c\tau<1 \text{ mm}$	1507.05493
Long-lived particles	Direct $\tilde{\chi}_1^\pm\tilde{\chi}_1^\pm$ prod., long-lived $\tilde{\chi}_1^\pm$	Disapp. trk	1 jet	Yes	20.3	$\tilde{\chi}_1^\pm$	270 GeV		$m(\tilde{\chi}_1^\pm)-m(\tilde{\chi}_1^0)\sim 160 \text{ MeV}, \tau(\tilde{\chi}_1^\pm)=0.2 \text{ ns}$	1310.3675
	Direct $\tilde{\chi}_1^\pm\tilde{\chi}_1^\pm$ prod., long-lived $\tilde{\chi}_1^\pm$	dE/dx trk	-	Yes	18.4	$\tilde{\chi}_1^\pm$	482 GeV		$m(\tilde{\chi}_1^\pm)-m(\tilde{\chi}_1^0)\sim 160 \text{ MeV}, \tau(\tilde{\chi}_1^\pm)<15 \text{ ns}$	1506.05332
	Stable, stopped $\tilde{g}$ R-hadron	0	1-5 jets	Yes	27.9	$\tilde{g}$	832 GeV		$m(\tilde{\chi}_1^0)=100 \text{ GeV}, 10 \mu\text{s}<\tau(\tilde{g})<1000 \text{ s}$	1310.6584
	Stable $\tilde{g}$ R-hadron	trk	-	-	19.1	$\tilde{g}$	1.27 TeV			1411.6795
	GMSB, stable $\tilde{\tau}, \tilde{\chi}_1^0 \rightarrow \tilde{\tau}(\tilde{e}, \tilde{\mu})+\tau(e, \mu)$	1-2 $\mu$	-	-	19.1	$\tilde{\chi}_1^0$	537 GeV		$10<\tan\beta<50$	1411.6795
	GMSB, $\tilde{\chi}_1^0 \rightarrow \gamma\tilde{G}$ , long-lived $\tilde{\chi}_1^0$	2 $\gamma$	-	Yes	20.3	$\tilde{\chi}_1^0$	435 GeV		$2<\tau(\tilde{\chi}_1^0)<3 \text{ ns}, \text{SPS8 model}$	1409.5542
	$\tilde{g}\tilde{g}, \tilde{\chi}_1^0 \rightarrow e\bar{e}\nu/\mu\bar{\mu}\nu$	displ. $e\bar{e}/\mu\bar{\mu}$	-	-	20.3	$\tilde{\chi}_1^0$	1.0 TeV		$7<c\tau(\tilde{\chi}_1^0)<740 \text{ mm}, m(\tilde{g})=1.3 \text{ TeV}$	1504.05162
	GGM $\tilde{g}\tilde{g}, \tilde{\chi}_1^0 \rightarrow Z\tilde{G}$	displ. vtx + jets	-	-	20.3	$\tilde{\chi}_1^0$	1.0 TeV		$6<c\tau(\tilde{\chi}_1^0)<480 \text{ mm}, m(\tilde{g})=1.1 \text{ TeV}$	1504.05162
RPV	LFV $pp \rightarrow \tilde{\nu}_\tau + X, \tilde{\nu}_\tau \rightarrow e\mu/\tau\mu/\mu\tau$	$e\mu, e\tau, \mu\tau$	-	-	20.3	$\tilde{\nu}_\tau$	1.7 TeV		$\lambda'_{311}=0.11, \lambda'_{132/133/233}=0.07$	1503.04430
	Bilinear RPV CMSSM	2 $e, \mu$ (SS)	0-3 $b$	Yes	20.3	$\tilde{q}, \tilde{g}$	1.35 TeV		$m(\tilde{q})=m(\tilde{g}), c\tau_{\text{LSP}}<1 \text{ mm}$	1404.2500
	$\tilde{\chi}_1^\pm\tilde{\chi}_1^\pm, \tilde{\chi}_1^\pm \rightarrow W\tilde{\chi}_1^0, \tilde{\chi}_1^0 \rightarrow ee\tilde{\nu}_\mu, e\mu\tilde{\nu}_e$	4 $e, \mu$	-	Yes	20.3	$\tilde{\chi}_1^\pm$	750 GeV		$m(\tilde{\chi}_1^0)>0.2 \times m(\tilde{\chi}_1^\pm), \lambda_{121} \neq 0$	1405.5086
	$\tilde{\chi}_1^\pm\tilde{\chi}_1^\pm, \tilde{\chi}_1^\pm \rightarrow W\tilde{\chi}_1^0, \tilde{\chi}_1^0 \rightarrow \tau\tau\tilde{\nu}_e, e\tau\tilde{\nu}_\tau$	3 $e, \mu + \tau$	-	Yes	20.3	$\tilde{\chi}_1^\pm$	450 GeV		$m(\tilde{\chi}_1^0)>0.2 \times m(\tilde{\chi}_1^\pm), \lambda_{133} \neq 0$	1405.5086
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow qqq$	0	6-7 jets	-	20.3	$\tilde{g}$	917 GeV		$\text{BR}(t)=\text{BR}(b)=\text{BR}(c)=0\%$	1502.05686
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{\chi}_1^0, \tilde{\chi}_1^0 \rightarrow qqq$	0	6-7 jets	-	20.3	$\tilde{g}$	870 GeV		$m(\tilde{\chi}_1^0)=600 \text{ GeV}$	1502.05686
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow \tilde{t}_1 t, \tilde{t}_1 \rightarrow bs$	2 $e, \mu$ (SS)	0-3 $b$	Yes	20.3	$\tilde{g}$	850 GeV			1404.250
	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow bs$	0	2 jets + 2 $b$	-	20.3	$\tilde{t}_1$	100-308 GeV			ATLAS-CONF-2015-026
	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow b\ell$	2 $e, \mu$	2 $b$	-	20.3	$\tilde{t}_1$	0.4-1.0 TeV		$\text{BR}(\tilde{t}_1 \rightarrow b\ell/\mu)>20\%$	ATLAS-CONF-2015-015
Other	Scalar charm, $\tilde{c} \rightarrow c\tilde{\chi}_1^0$	0	2 $c$	Yes	20.3	$\tilde{c}$	490 GeV		$m(\tilde{\chi}_1^0)<200 \text{ GeV}$	1501.01325

$10^{-1}$

$1$

Mass scale [TeV]

 $10^{-1}$ 

1

Mass scale [TeV]

\*Only a selection of the available mass limits on new states or phenomena is shown. All limits quoted are observed minus  $1\sigma$  theoretical signal cross section uncertainty.

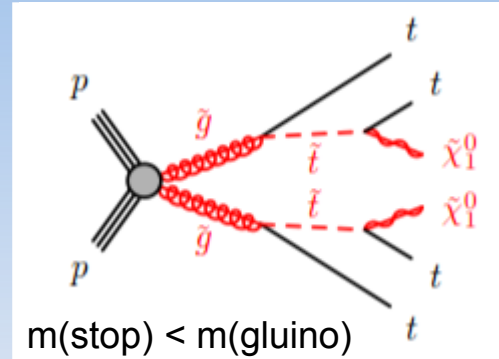
# An example of multi-object SUSY searches



## SUSY scenarios with the top squark as the lightest quark partner

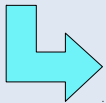
**SMS:** simplified model spectra →

Gluino pair production, with subsequent decays to two top quarks and the LSP through either an on-shell (the upper diagram) or a virtual (the lower) top squark.



Summary of the searches for squarks and gluinos using  $\sqrt{s} = 8$  TeV  $pp$  collisions with the ATLAS experiment at the LHC

arXiv:1507.05525



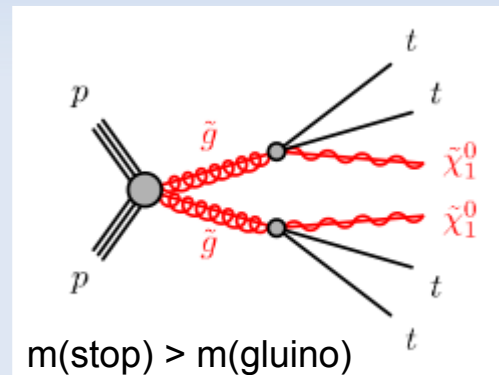
0/1-lepton + 3 b-jets +  $E_T^{\text{miss}}$

arXiv:1407.0600; JHEP 1410 (2014) 24

Search for supersymmetry in  $pp$  collisions at  $\sqrt{s} = 8$  TeV in events with a single lepton, large jet multiplicity, and multiple b jets

CMS-SUS-13-007

arXiv:1311.4937  
PLB 733 328 (2014)



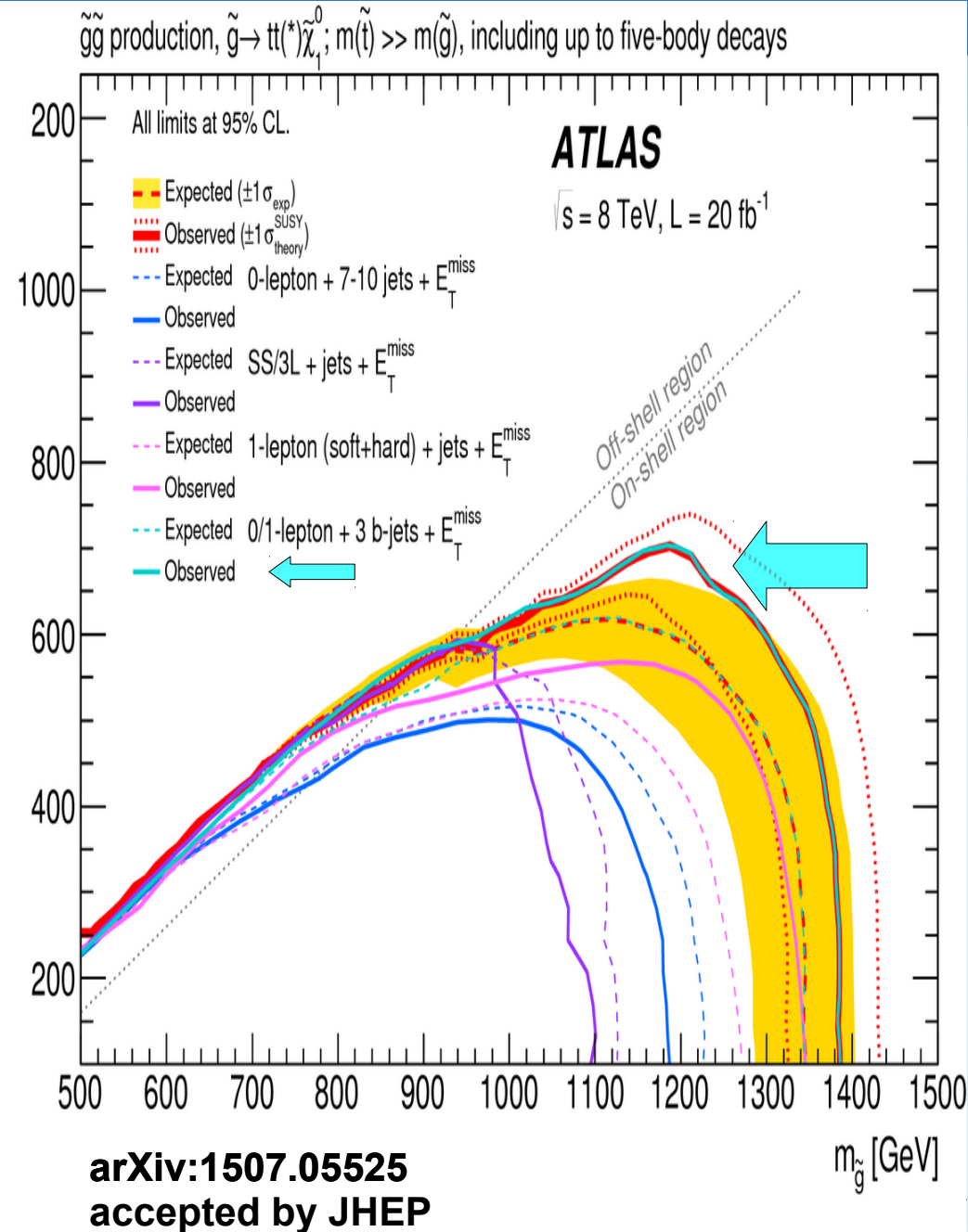
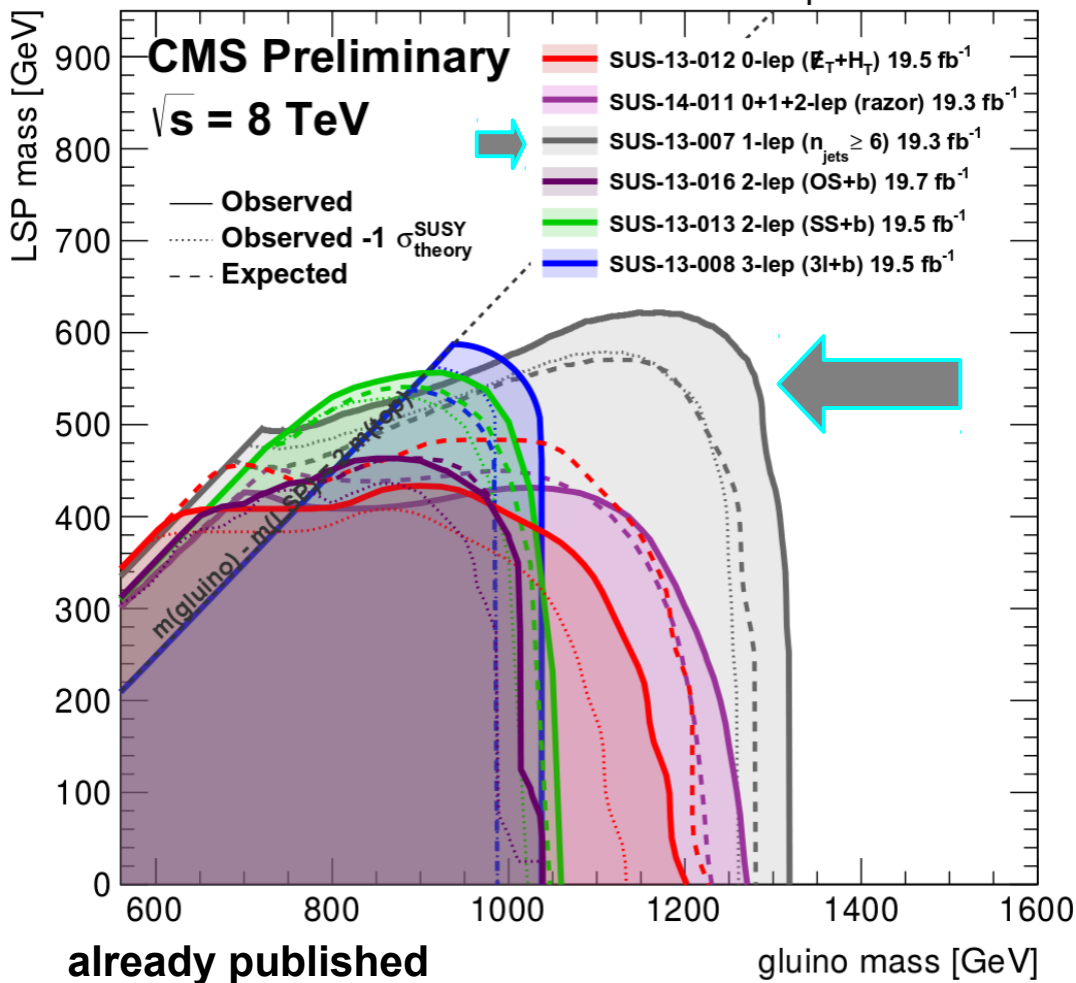
**Signature: 1 lepton (also 0 for ATLAS),  $E_T^{\text{miss}}$ , many jets of which multiple are b-tagged**

Dozens of different analyses performed by ATLAS and CMS to search for squarks and gluino production. Interpretation in several models including phenomenological pMSSM framework and SMS.

# An example of multi-object SUSY searches

Results of analyses **highlighted** in the previous slide are depicted with **arrows**.

$\tilde{g}\tilde{g}$  production,  $\tilde{g} \rightarrow t \bar{t} \tilde{\chi}_1^0$





## 2<sup>nd</sup> example: direct electroweakinos

Direct electroweak production of charginos and neutralinos (or sleptons) may dominate at the LHC if the strongly interacting SUSY particles are heavy.

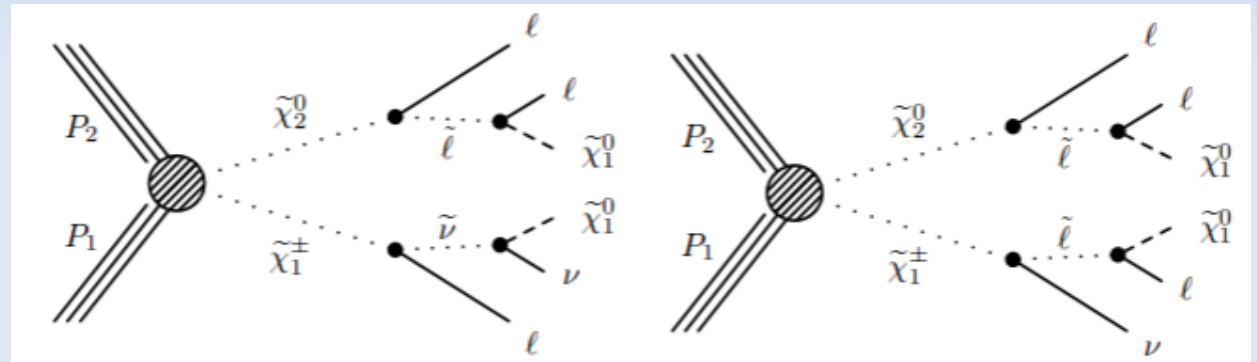
The corresponding final states do not necessarily contain much hadronic activity and thus may have eluded detection.

X-sections are much smaller than in the case of production *via* strong interaction.

The **CMS** Collaboration;

*Searches for electroweak production of charginos, neutralinos, and sleptons decaying to leptons and W, Z, and Higgs bosons in pp collisions at 8 TeV;*

arXiv:1405.7570, CMS-SUS-13-006  
Eur. Phys. J. C (2014) 74:3036



The **ATLAS** Collaboration;

*Search for direct production of charginos and neutralinos in events with three leptons and missing transverse momentum in  $\sqrt{s}=8\text{TeV}$  pp collisions with the ATLAS detector;*

arXiv:1402.7029;  
(so called, 3L analysis);  
**JHEP 04 (2014) 169**

$\tilde{\ell}_L$ -mediated: the  $\tilde{\chi}_1^\pm$  and  $\tilde{\chi}_2^0$  decay with a branching fraction of 1/6 via  $\tilde{e}_L$ ,  $\tilde{\mu}_L$ ,  $\tilde{\tau}_L$ ,  $\tilde{\nu}_e$ ,  $\tilde{\nu}_\mu$ , or  $\tilde{\nu}_\tau$  with masses  $m_{\tilde{\nu}} = m_{\tilde{\ell}_L} = (m_{\tilde{\chi}_1^0} + m_{\tilde{\chi}_1^\pm})/2$ ,

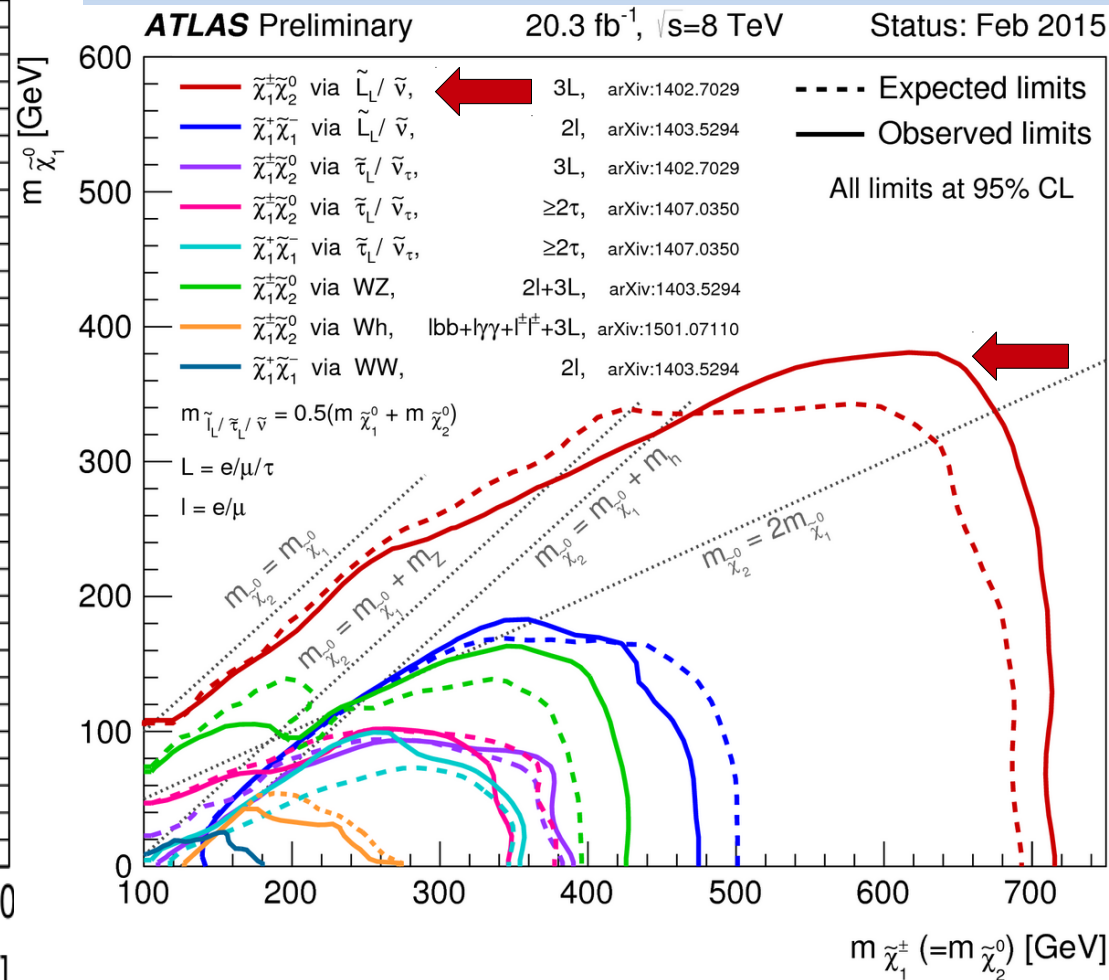
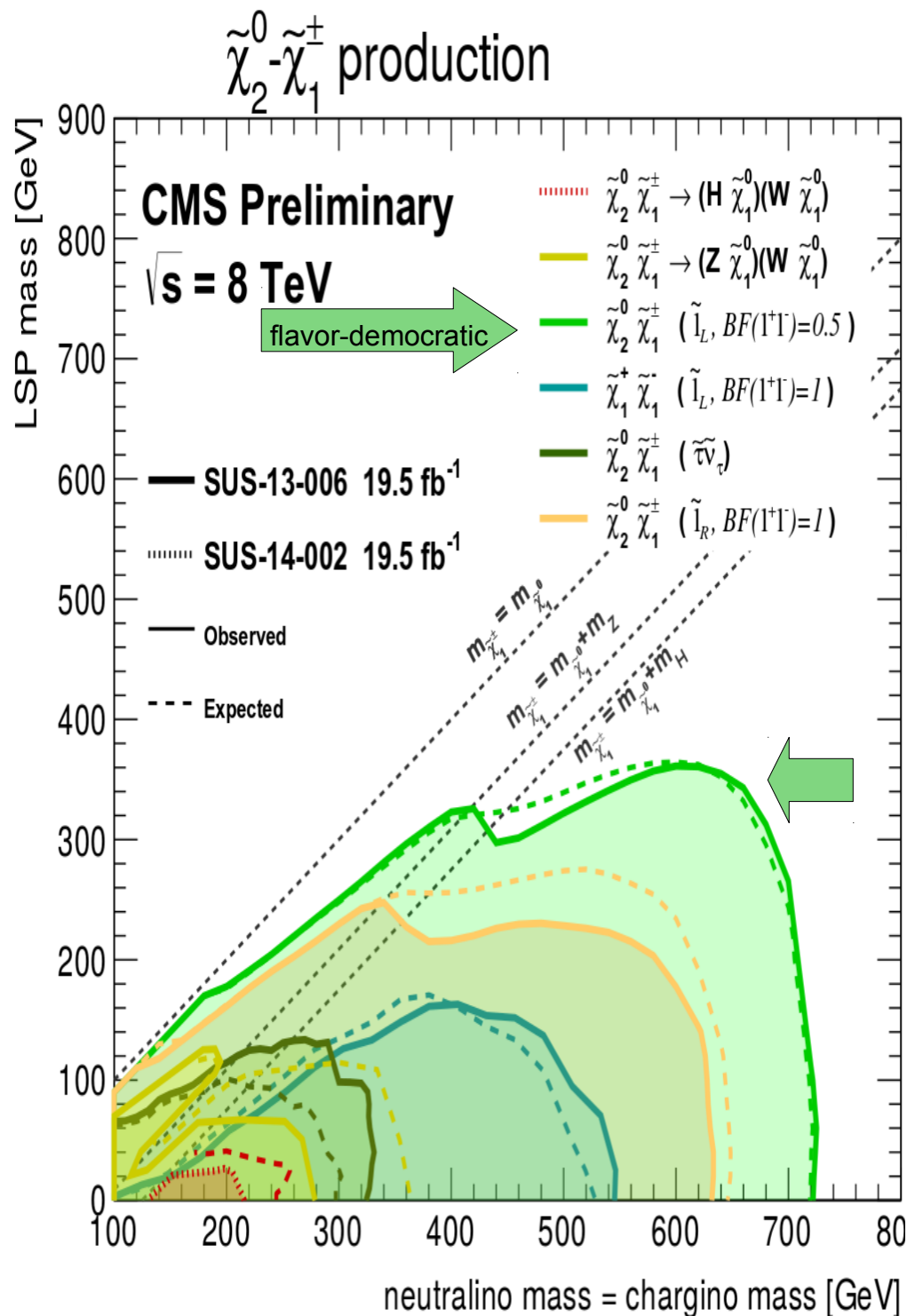
$\tilde{\tau}_L$ -mediated: the first- and second-generation sleptons and sneutrinos are assumed to be heavy, so that the  $\tilde{\chi}_1^\pm$  and  $\tilde{\chi}_2^0$  decay with a branching fraction of 1/2 via  $\tilde{\tau}$  or  $\tilde{\nu}_\tau$  with masses  $m_{\tilde{\nu}_\tau} = m_{\tilde{\tau}} = (m_{\tilde{\chi}_1^0} + m_{\tilde{\chi}_2^0})/2$ ,

**Many more scenarios taken into account by ATLAS and CMS**

# direct electroweakinos



different lines show  
constraints on  
different scenarios



All BSM searches not classified as SUSY falls into Exotic(a) (or B2G in the CMS).

Topological searches:

- Heavy narrow resonances (dijets, dileptons, diphotons)  
→  $Z'$ , RS gravitons, compositeness ...
- Single lepton +  $E_t^{\text{miss}}$  → leptonic decays of  $W'$
- Leptons and jets → leptoquarks
- Multiobject topologies → microscopic black holes
- Mono - "something visible" → dark matter
- ...

At the SUSY – Exotica border

- RPV SUSY (more on the SUSY side)
- Long-Lived Particles (LLP)
- Many topologies with jets and leptons are also common for SUSY and Exotica
- ...

EXOTICA *par excellence*

- Monopoles
- Multi-charged or fractionally charged particles
- Jet extinction scale
- ...

All BSM searches not classified as SUSY falls into Exotic(a) (or B2G in the CMS).

Topological searches:

- Heavy narrow resonances (dijets, dileptons, diphotons)  
→  $Z'$ , RS gravitons, compositeness ...
- Single lepton +  $E_t^{\text{miss}}$  → leptonic decays of  $W'$
- Leptons and jets → leptoquarks
- Multiobject topologies → microscopic black holes
- Mono - "something visible" → dark matter
- ...

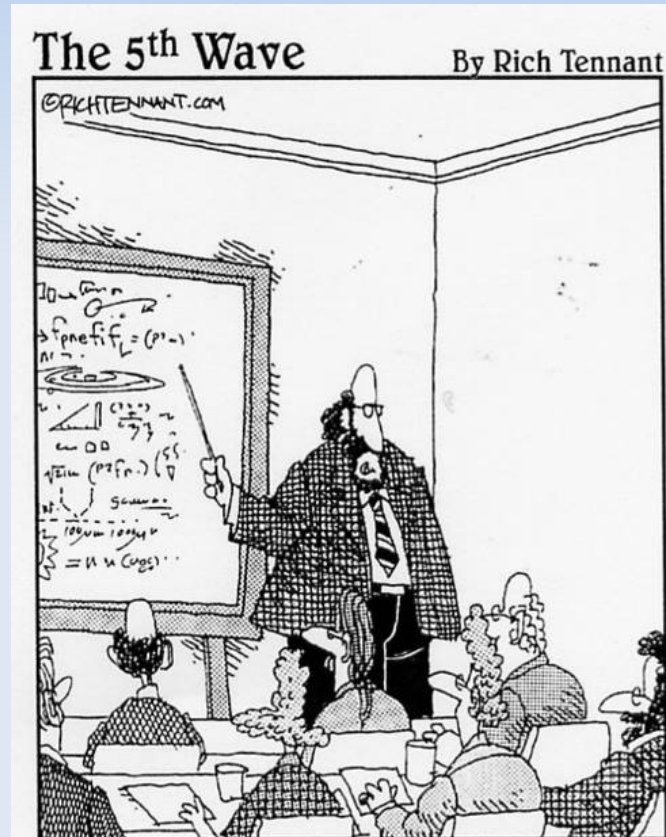
At the SUSY – Exotica border

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EXOTICA *par excellence*

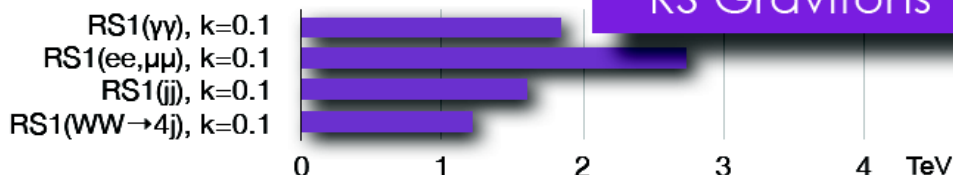
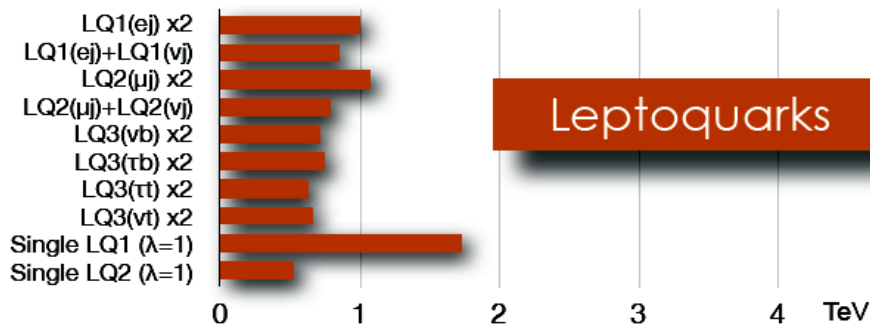
- Monopoles
- Multi-charged and fractionally charged particles
- Jet extinction scale
- ...

In short: we search for everything except 'doesn't matter' particles.

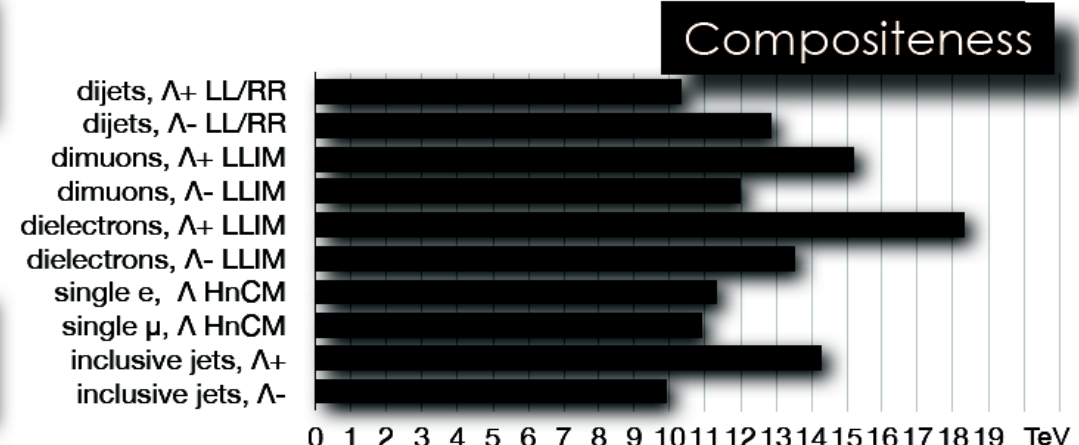
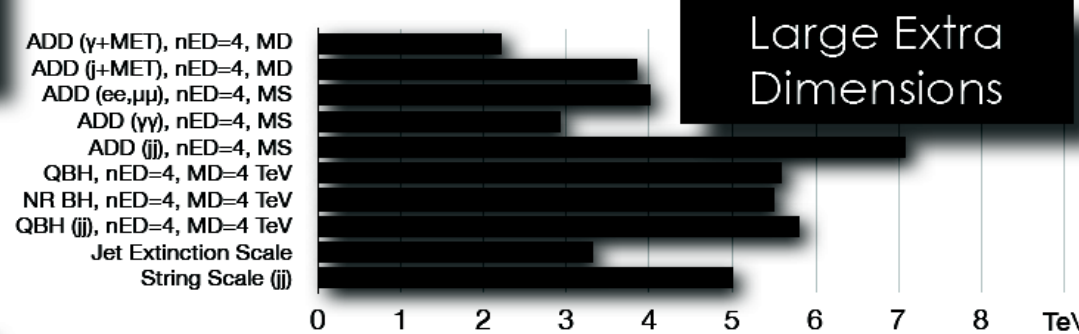
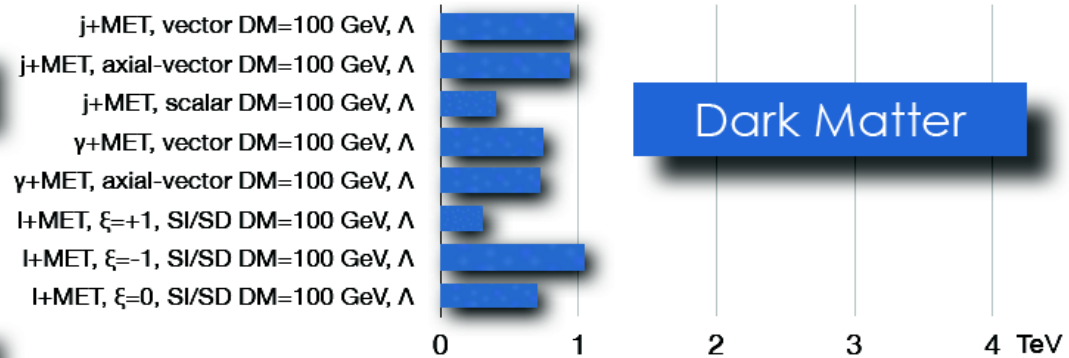
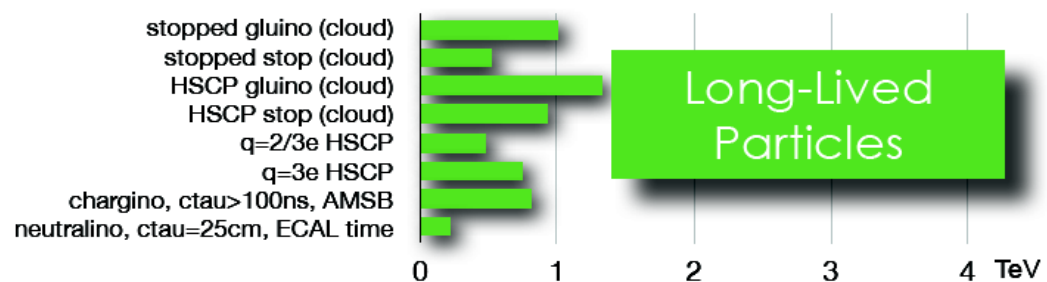
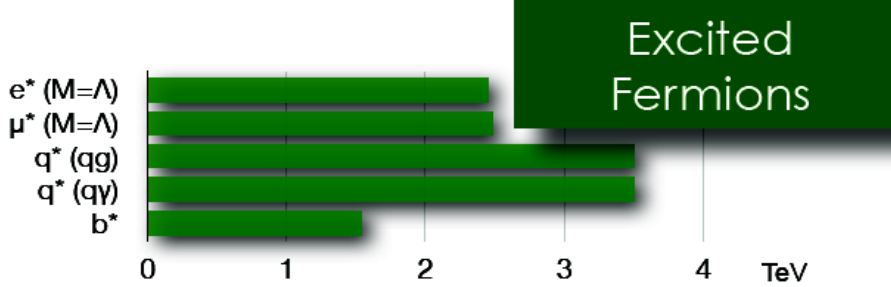
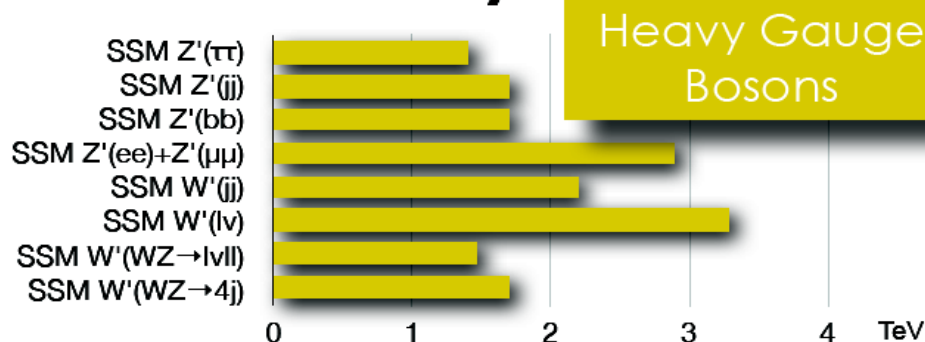


"After the discovery of 'antimatter' and 'dark matter', we have just confirmed the existence of 'doesn't matter', which does not have any influence on the Universe whatsoever."



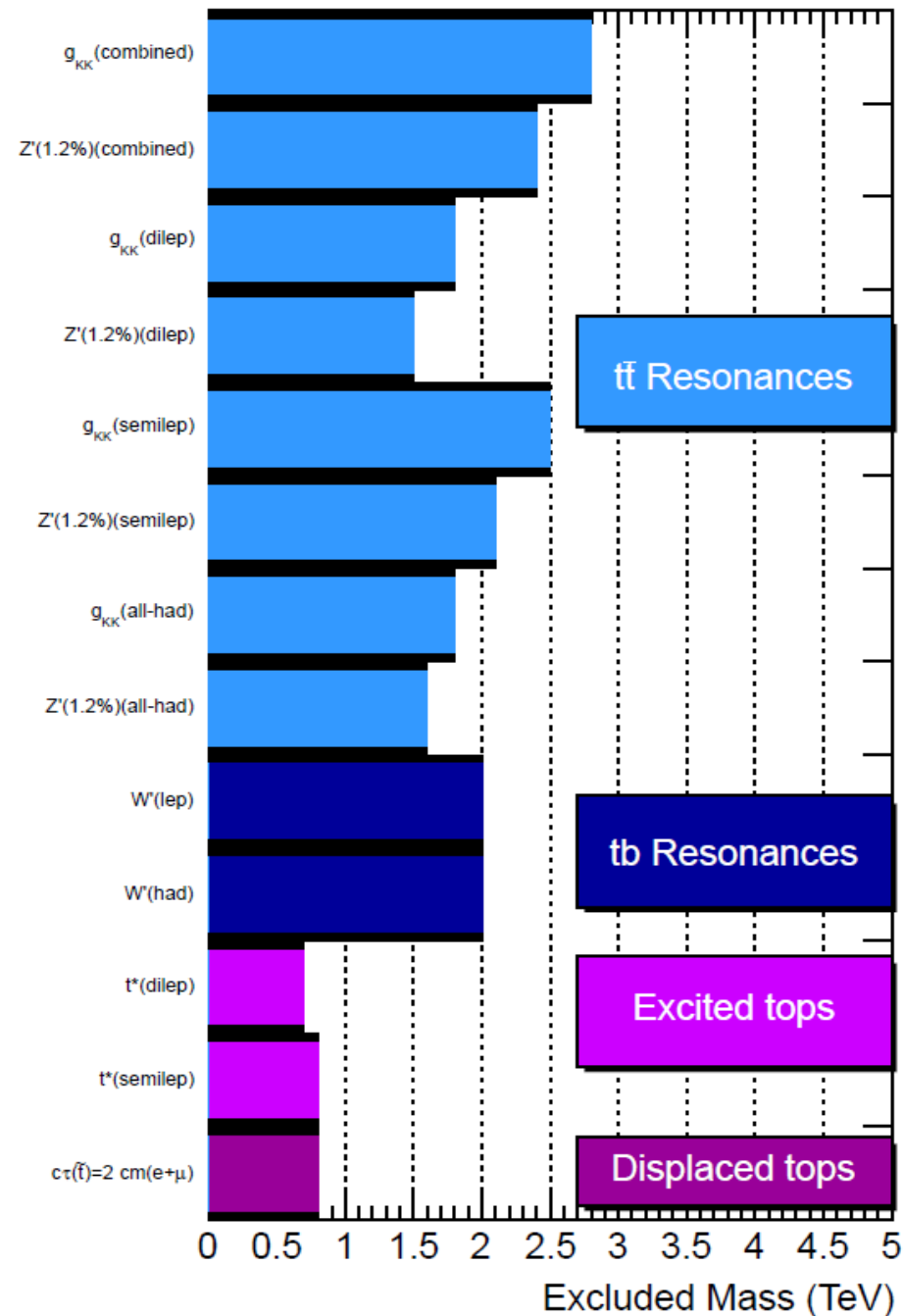
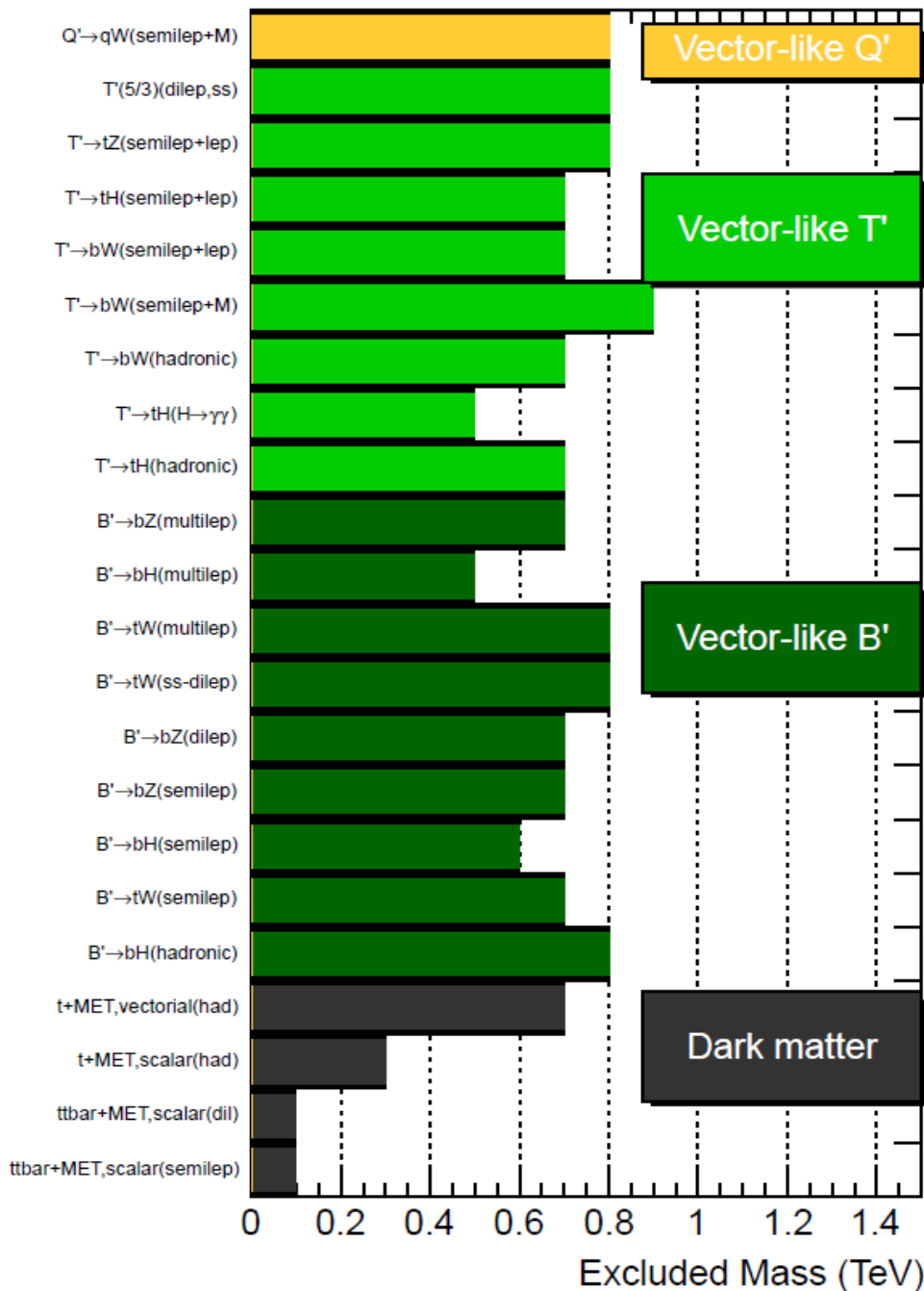


**CMS Preliminary**



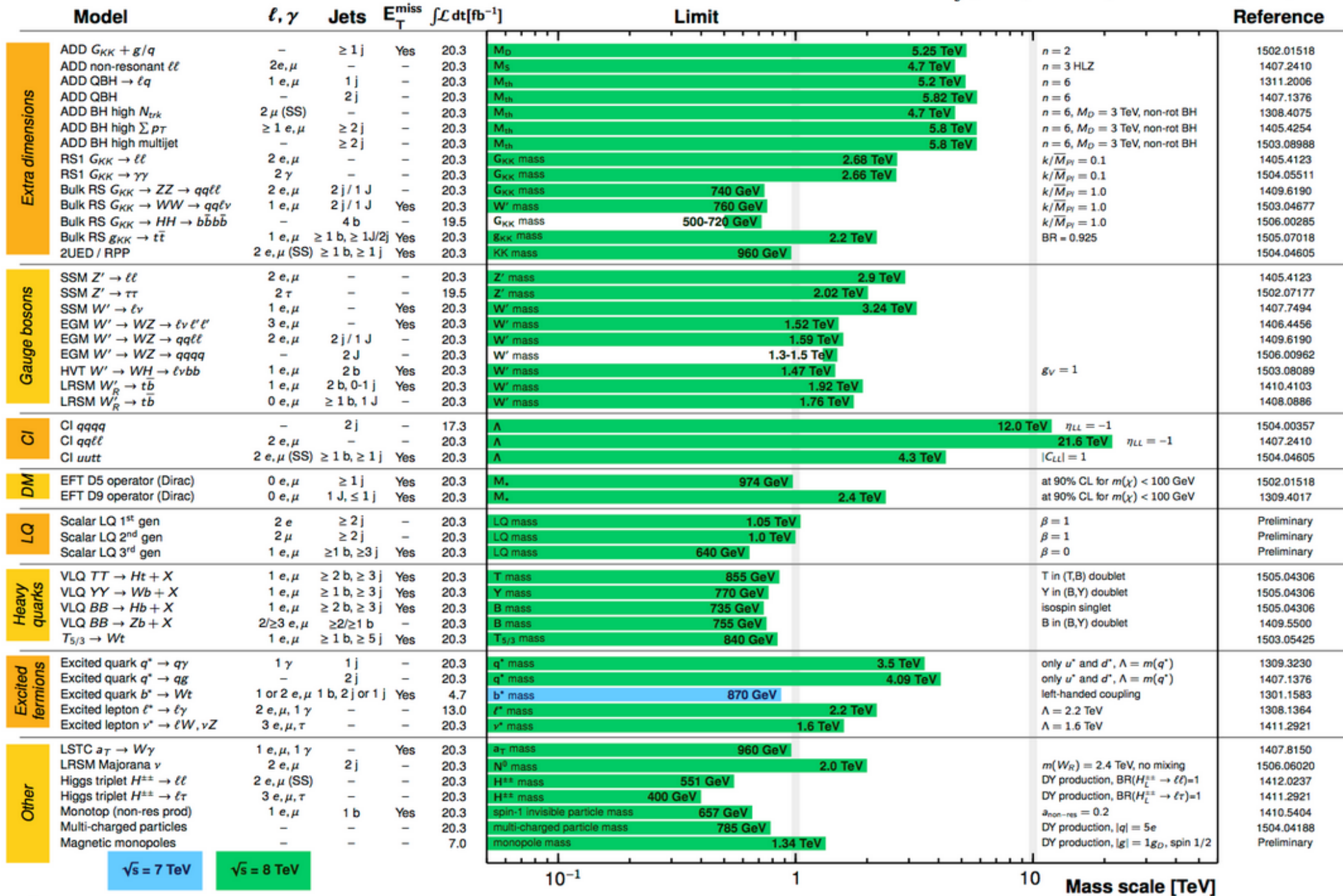
# CMS Searches for New Physics Beyond Two Generations (B2G)

95% CL Exclusions (TeV)



$$\int \mathcal{L} dt = (4.7 - 20.3) \text{ fb}^{-1}$$

$$\sqrt{s} = 7, 8 \text{ TeV}$$



\*Only a selection of the available mass limits on new states or phenomena is shown.

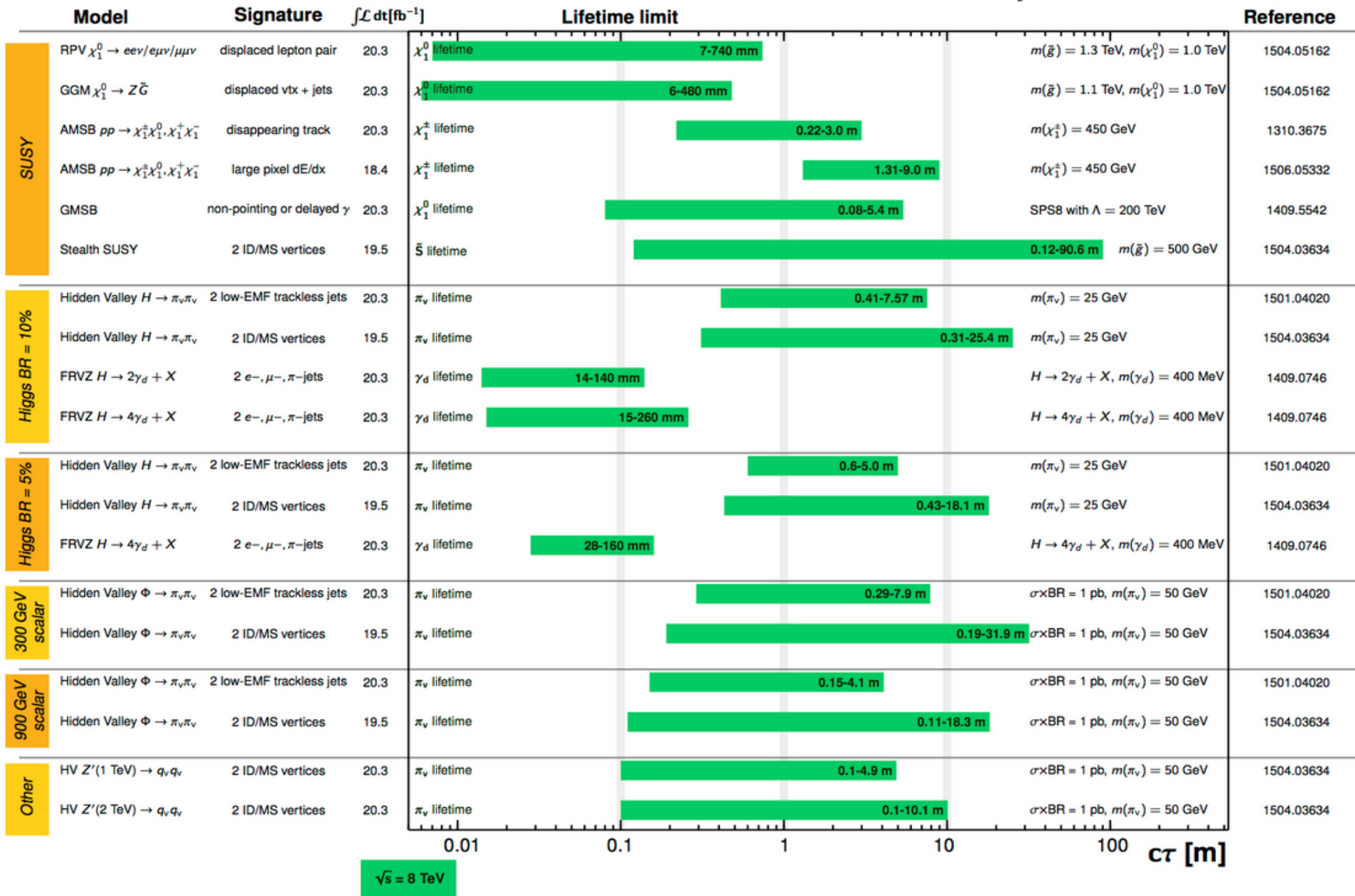


# ATLAS Long-lived Particle Searches\* - 95% CL Exclusion

Status: July 2015

ATLAS Preliminary

$$\int \mathcal{L} dt = (18.4 - 20.3) \text{ fb}^{-1} \quad \sqrt{s} = 8 \text{ TeV}$$



\*Only a selection of the available lifetime limits on new states is shown.



arXiv:1502.02522; CMS-EXO-13-006; Eur.Phys.J. C75 (2015) 7, 325

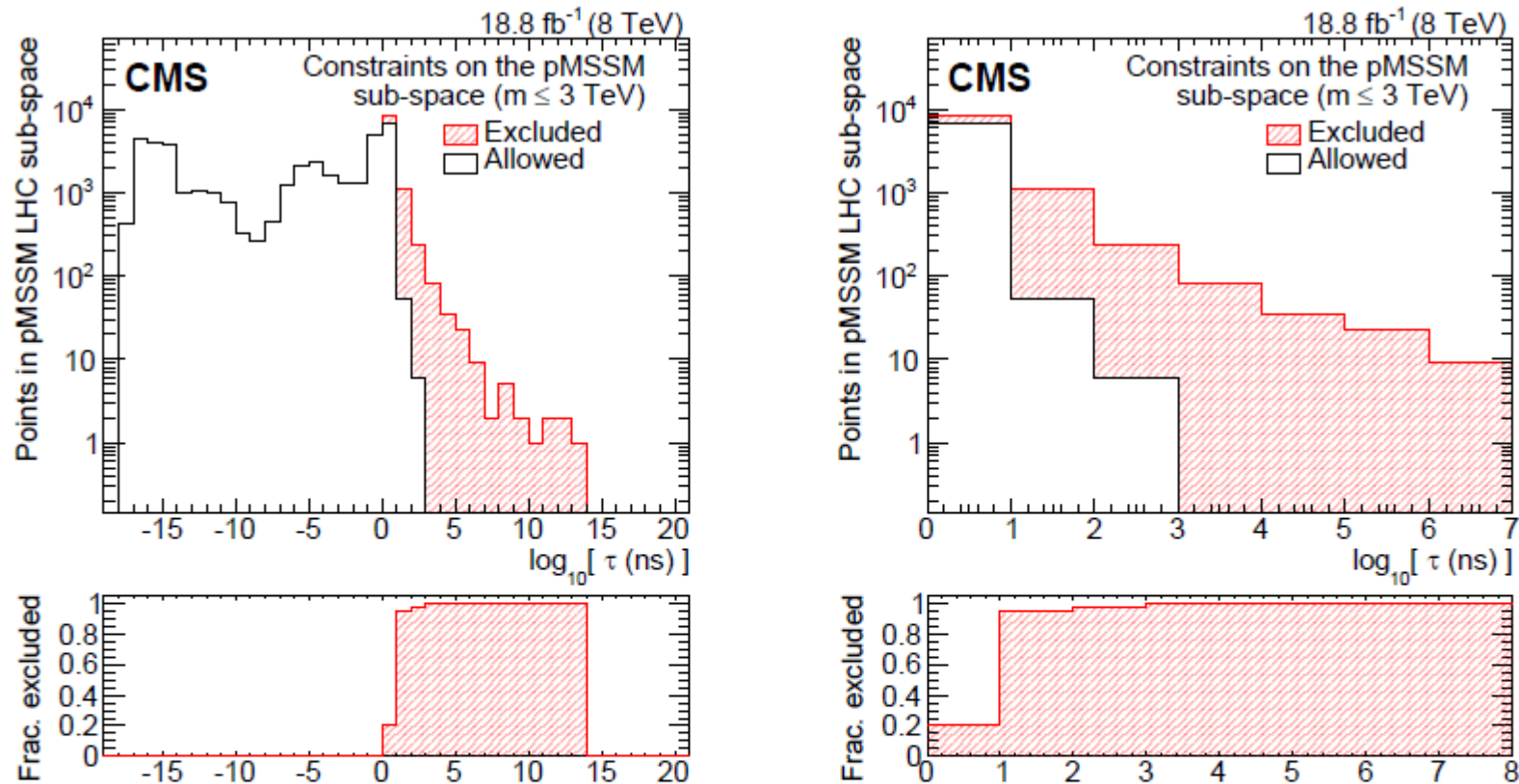


Figure 3: (left) Number of pMSSM points, in the sub-space covering sparticle masses up to about 3 TeV, that are excluded at a 95% CL (hatched red) or allowed (white) as a function of the chargino lifetime. (right) Enlargement of the long-lived region. The bottom panel shows the fraction of pMSSM points excluded by the analysis based on the results from the HSCP search [arXiv:1305.0491; CMS-EXO-12-026; JHEP07(2013)122].

HSCP → Heavy Stable Charged Particles

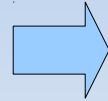
# LLP @ ATLAS

an example of complementarity:  
3 different analyses to cover  
wide lifetime range



LLP → Long-Lived Particles  
SMP → Stable Massive Particles

Constraints on the chargino  
mass-vs-lifetime plane  
for an AMSB model with  
 $\tan(\beta)=5$  and  $\mu>0$ .



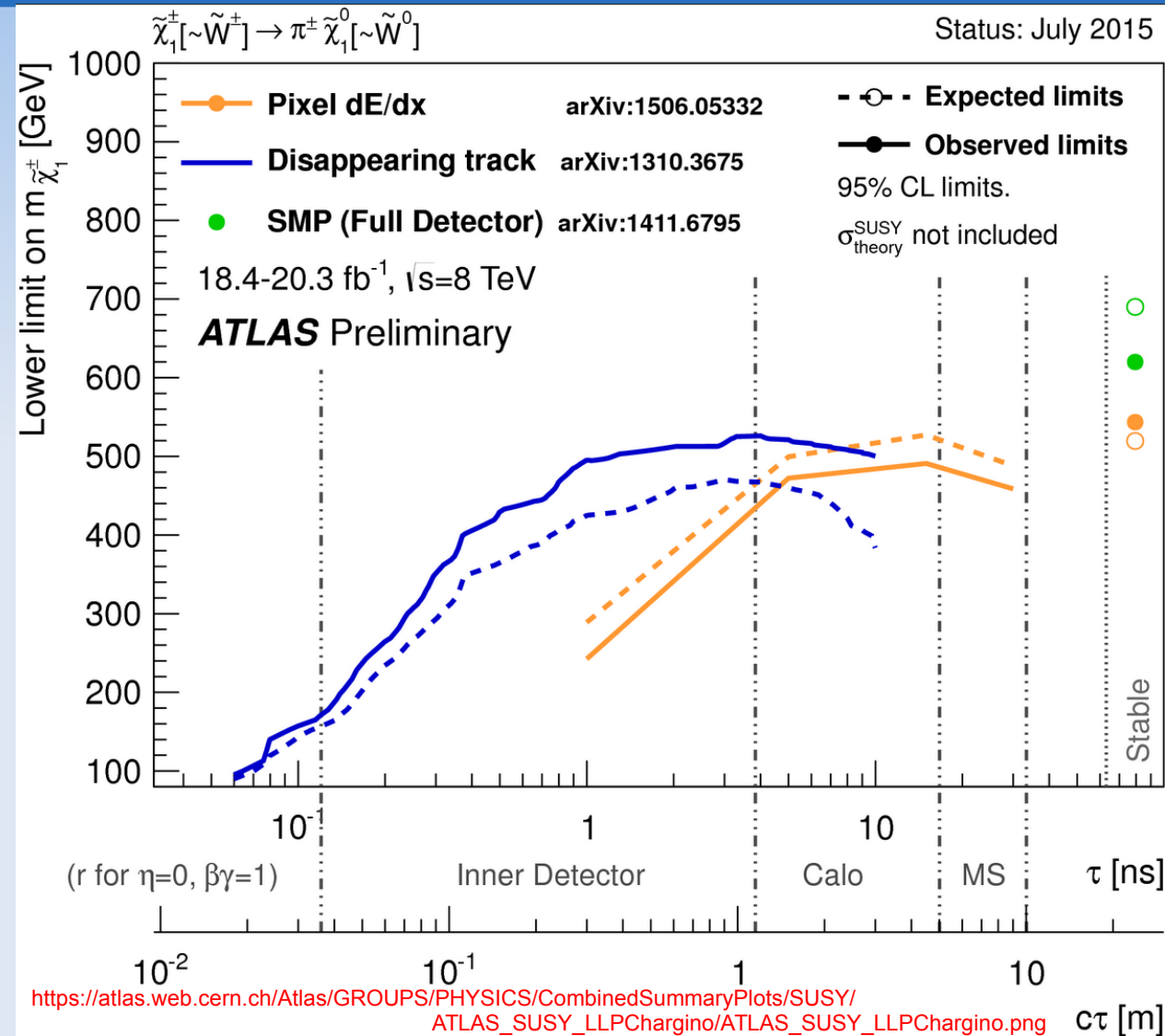
The wino-like chargino is pair-produced  
and decays to the wino-like neutralino  
and a very soft charged pion.

The solid lines indicate the observed  
limits, while the dashed lines indicate  
the expected limits.

The area below the curves is excluded.

The analyses have sensitivity  
at lifetimes other than those shown,  
but only the limits at tested lifetimes  
are shown.

The dots represent results for which the particle is assumed  
to be stable. In this context, stable means escaping the detector.

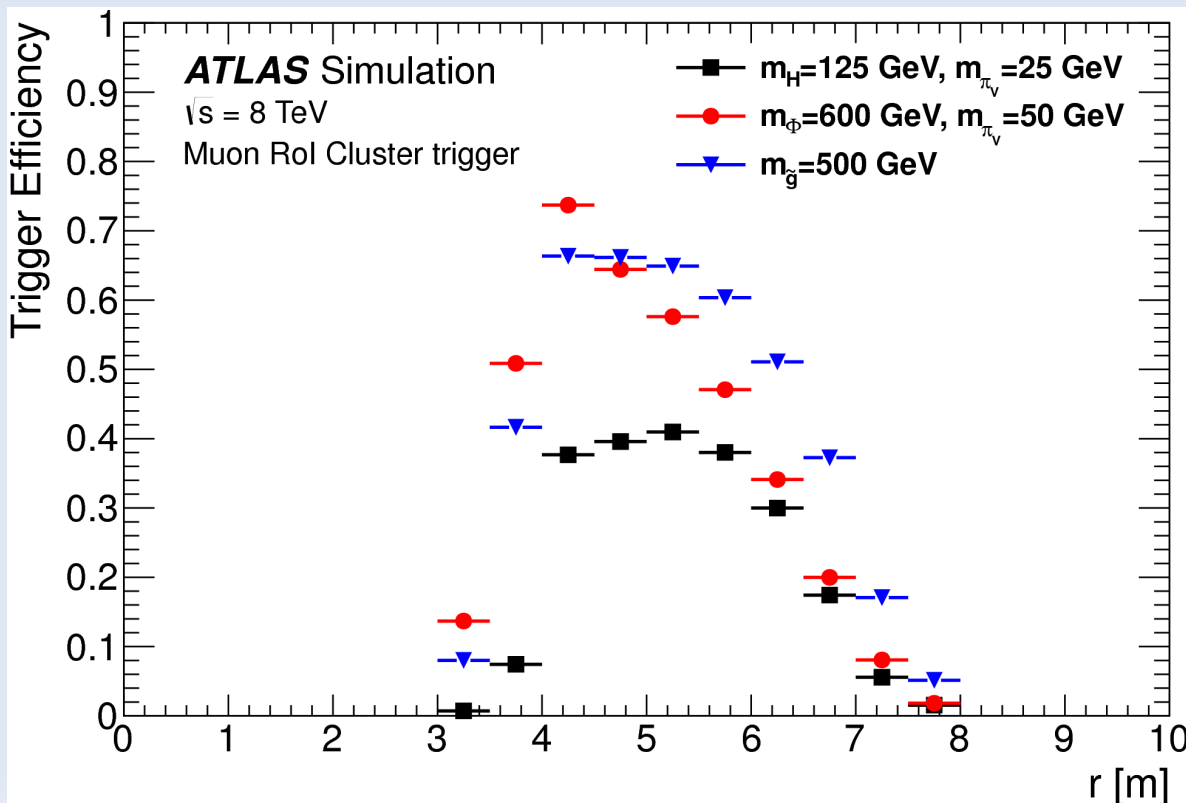
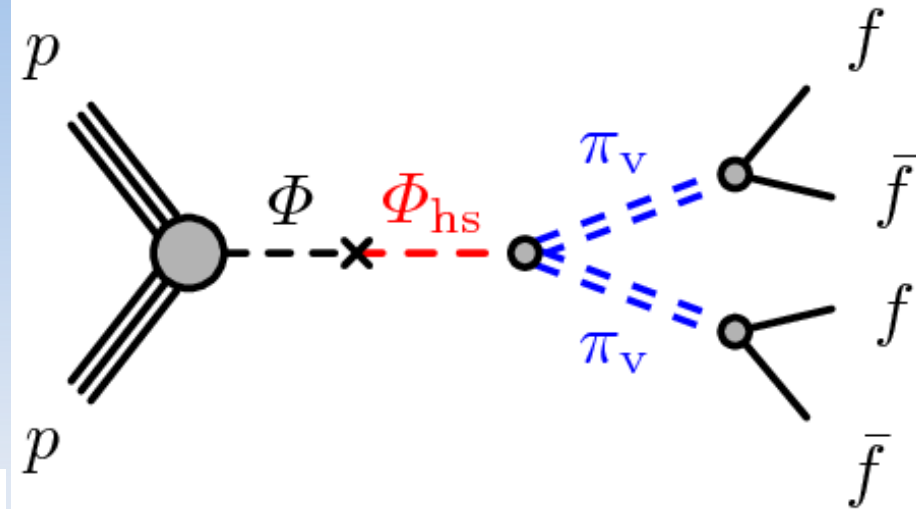


arXiv:1506.05332; accepted by EPJC  
arXiv:1411.6795; JHEP 1501 (2015) 068  
arXiv:1310.3675; Phys.Rev. D88 (2013) 11, 112006

# LLP: displaced jets

Search for long-lived, weakly-interacting particles that decay to displaced hadronic jets in proton-proton collisions at  $\sqrt{s}=8\text{TeV}$  with the **ATLAS** detector

arXiv:1504.03634,  
Phys. Rev. D 92, 012010 (2015)



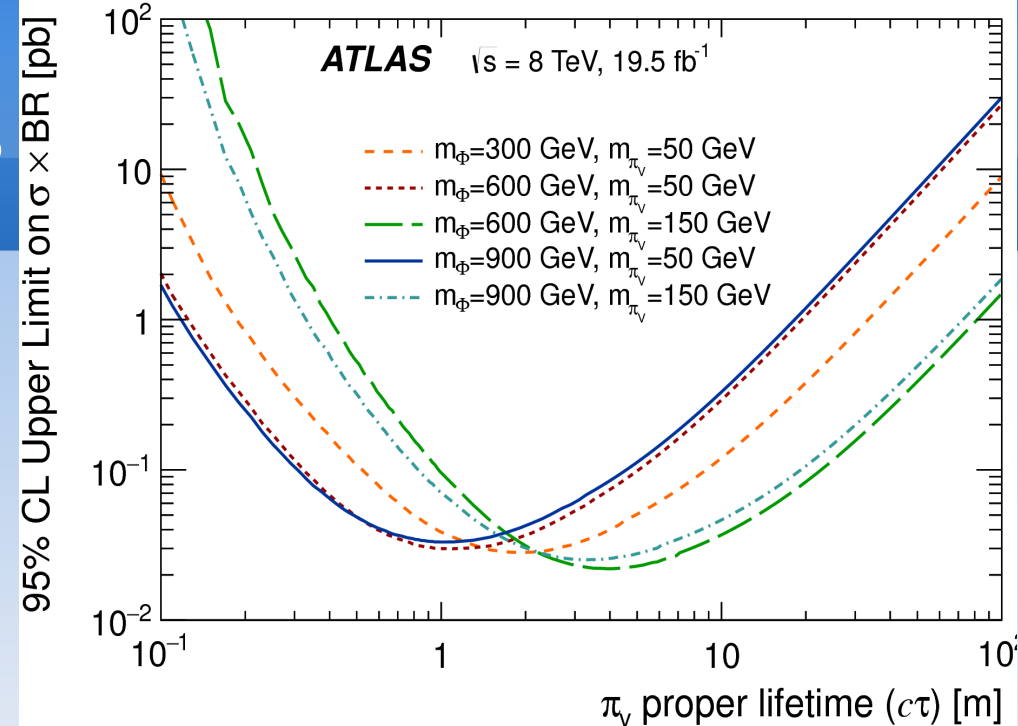
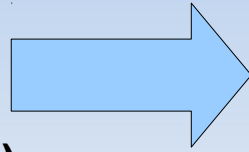
Custom Muon RoI Cluster trigger: it is a signature-driven trigger that selects decays of neutral particles in the muon system (MS).

It is efficient for hadronic decays occurring in the region from the outer radius of the HCal to the middle of the MS.

# LLP: displaced jets

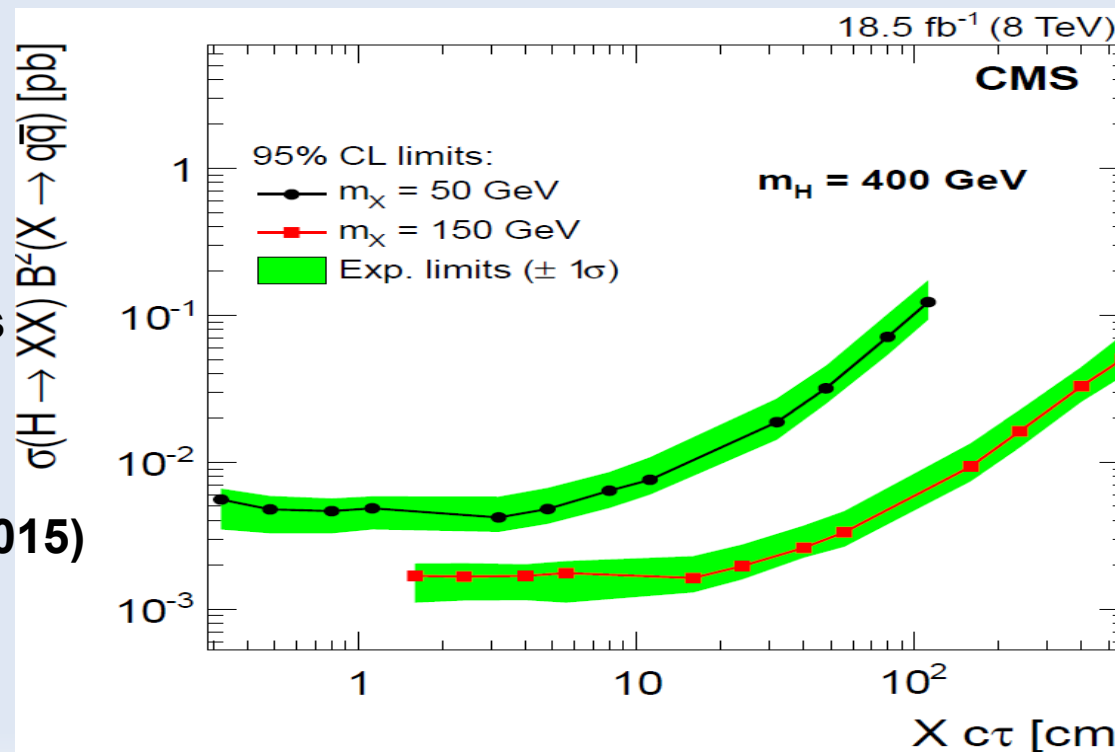
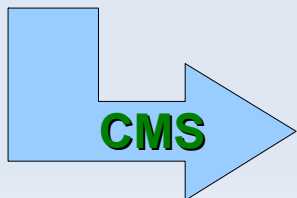
Search for long-lived, weakly-interacting particles that decay to displaced hadronic jets in proton-proton collisions at  $\sqrt{s}=8\text{TeV}$  with the **ATLAS** detector

arXiv:1504.03634,  
Phys. Rev. D 92, 012010 (2015)



Search for long-lived neutral particles decaying to quark-antiquark pairs in proton-proton collisions at  $\sqrt{s} = 8 \text{ TeV}$

arXiv:1411.6530,  
Phys. Rev. D 91, 012007 (2015)



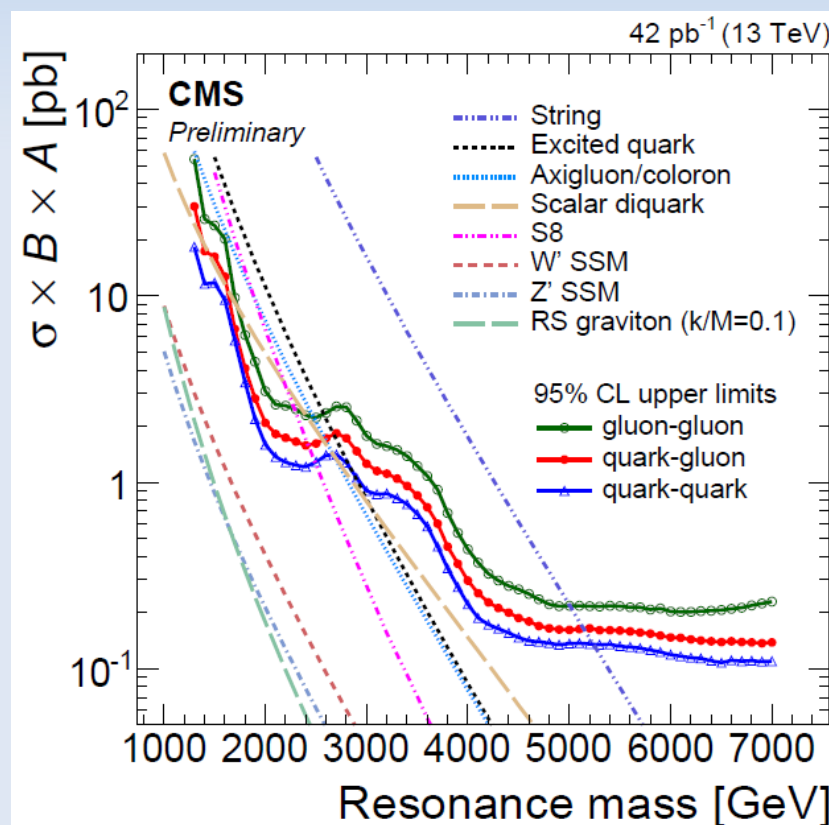
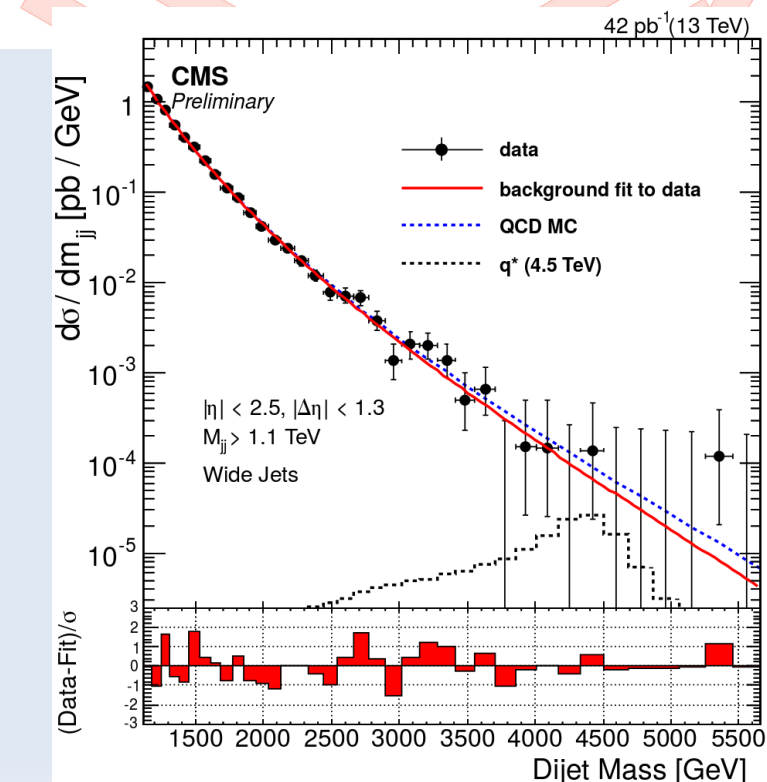


# CMS @ 13 TeV

Search for narrow resonances using the dijet mass spectrum with  $42 \text{ pb}^{-1}$  of pp collisions at  $\sqrt{s} = 13 \text{ TeV}$

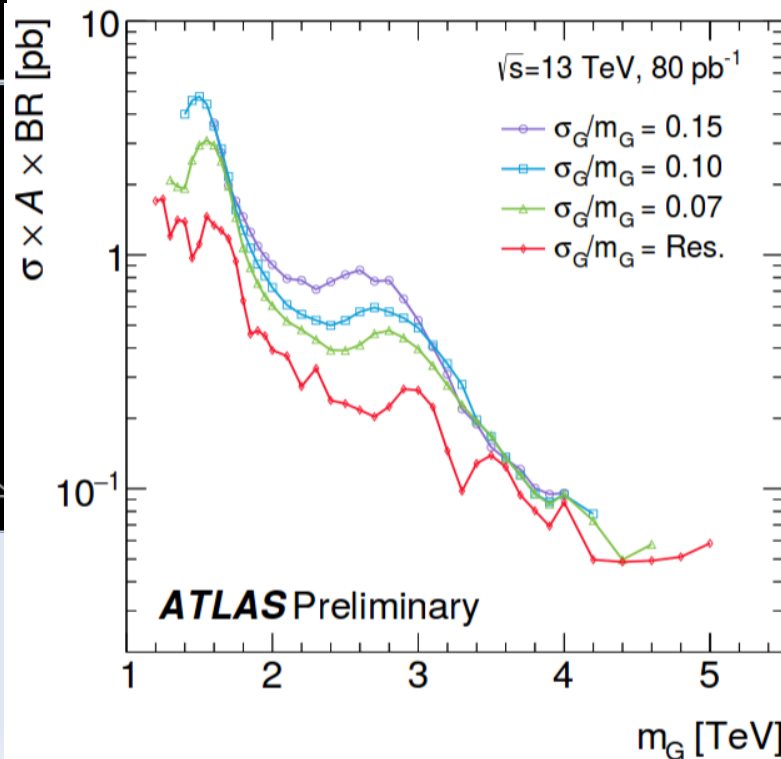
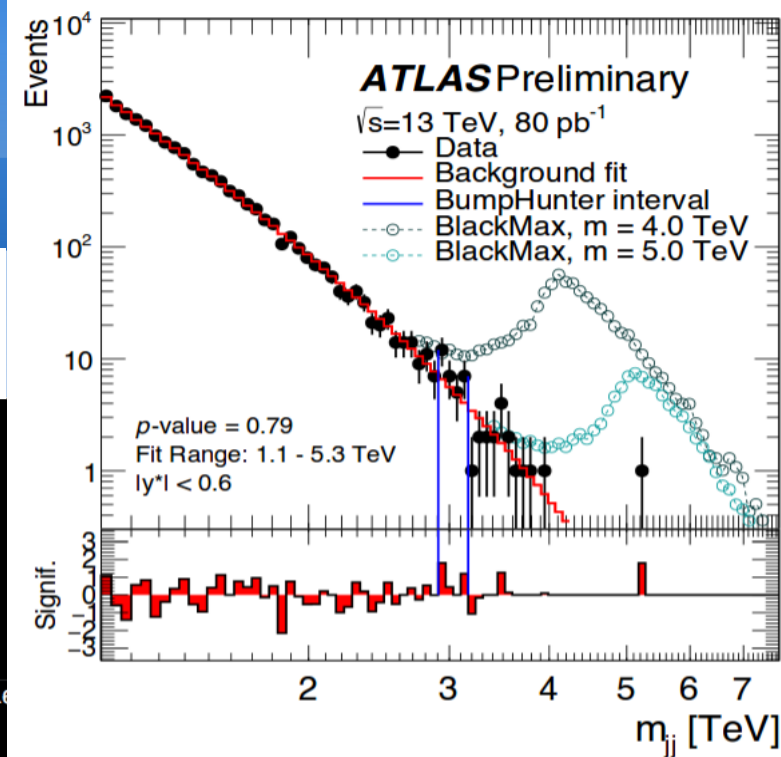
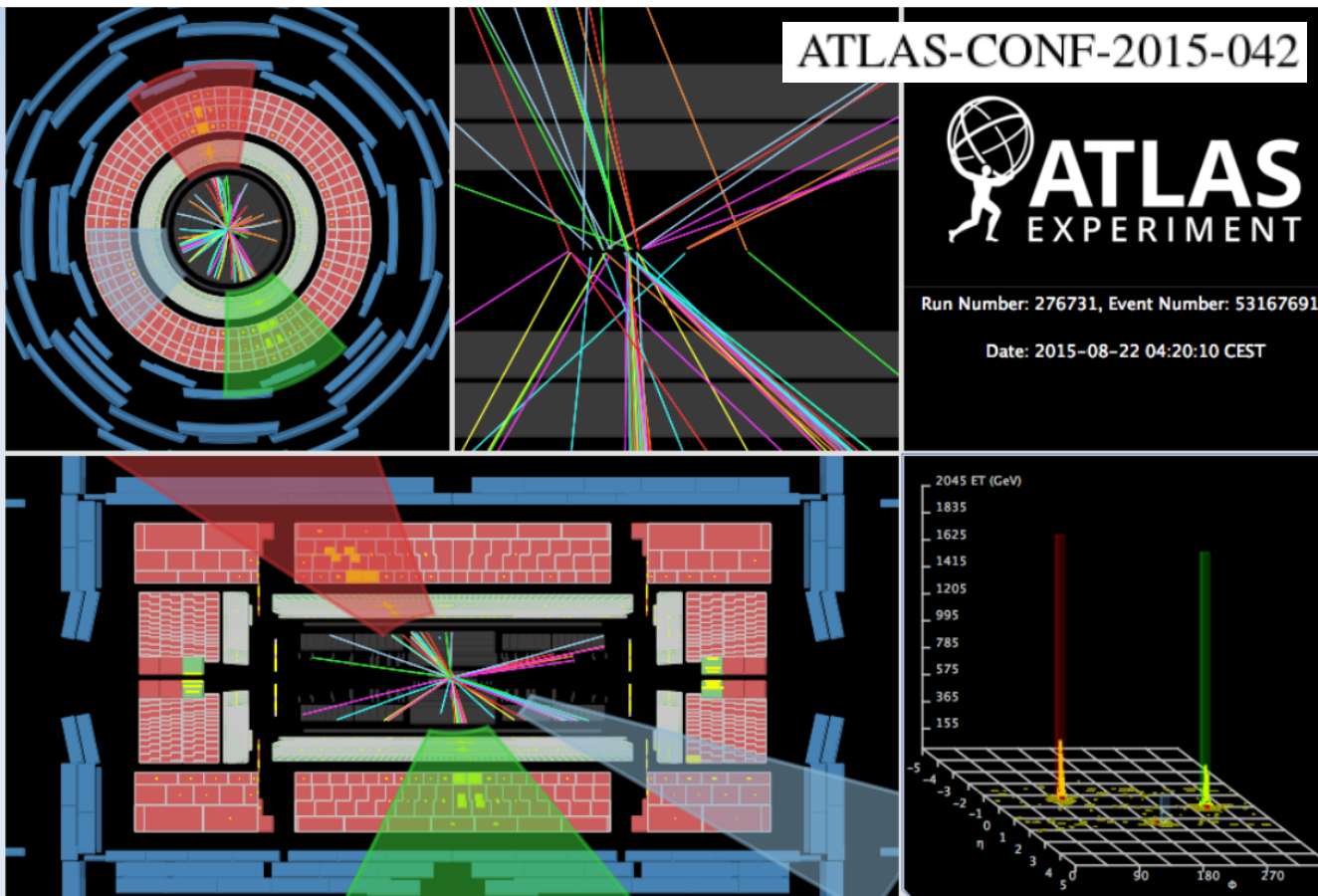
CMS PAS EXO-15-001

CMS Experiment at LHC, CERN  
Data recorded: Sun Jul 12 01:52:51 2015 CDT  
Run/Event: 251562 / 310157776  
Lumi section: 347  
Dijet Mass 5.4 TeV



# ATLAS @ 13 TeV

## Search for New Phenomena in Dijet Mass and Angular Distributions with the ATLAS Detector at $\sqrt{s} = 13$ TeV



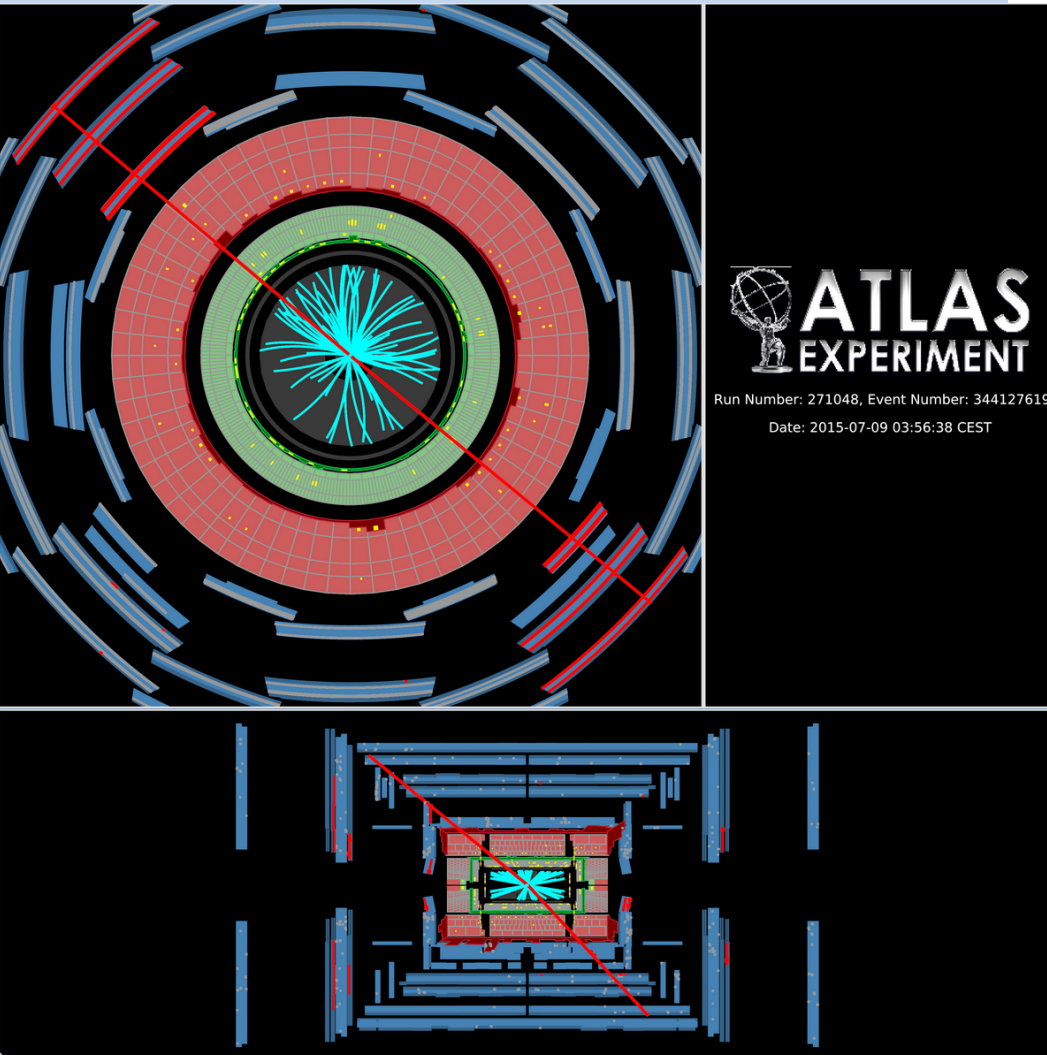
Other ATLAS @ 13 TeV Exotica:

"Search for evidence for strong gravity in jet final states ...",  
ATLAS-CONF-2015-043

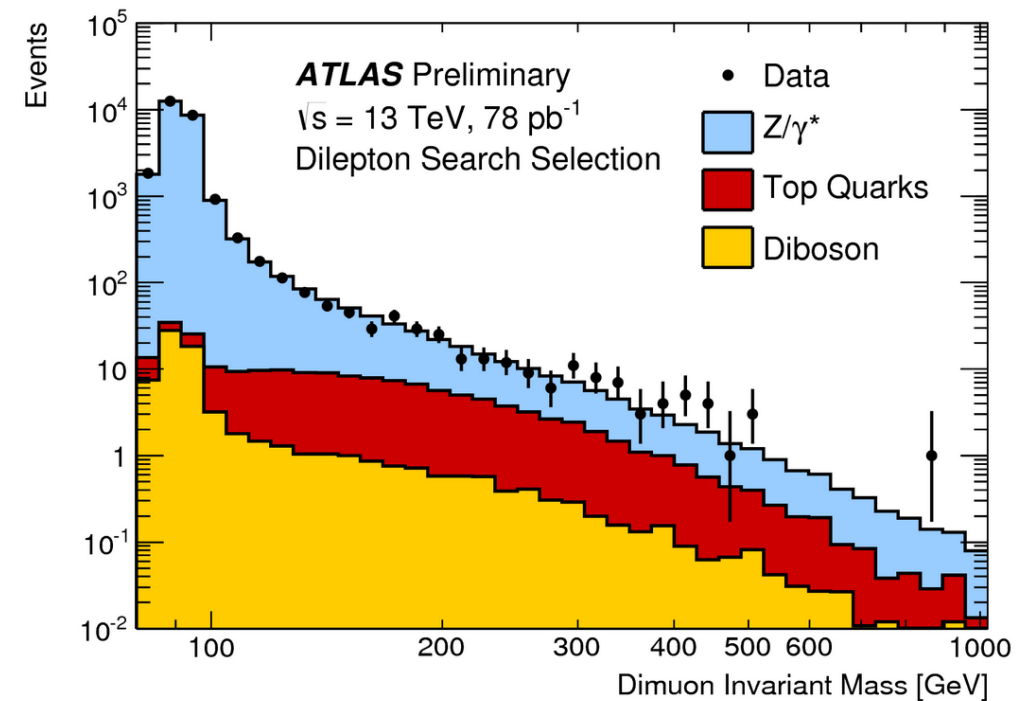
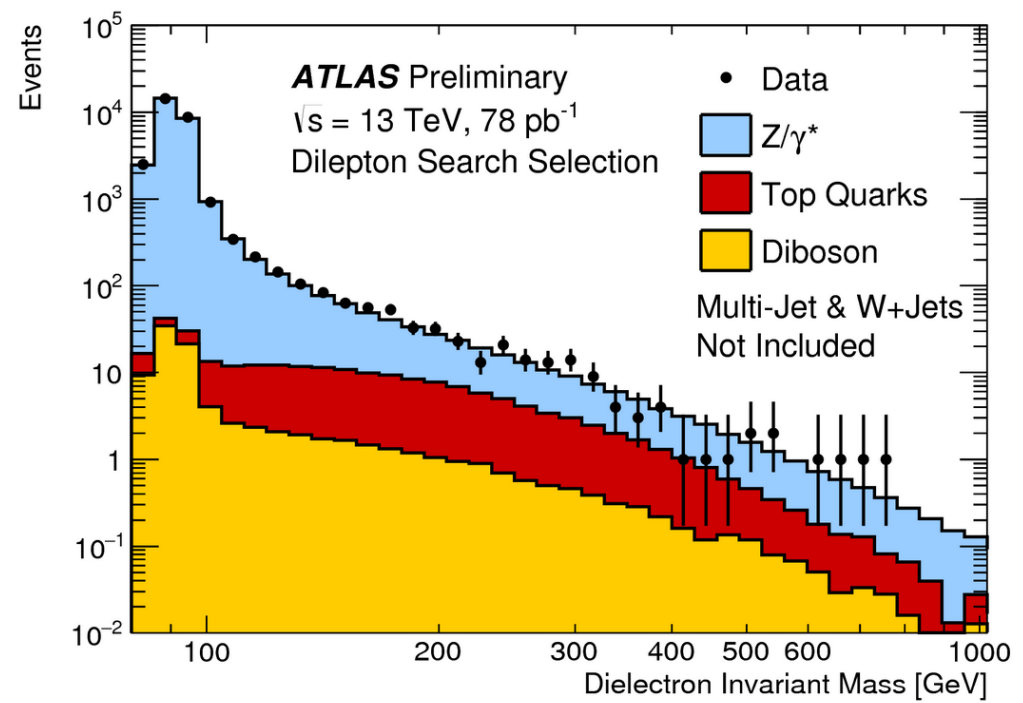
"Search for TeV-scale gravity signatures in high-mass final states with  
leptons and jets ...", ATLAS-CONF-2015-046

# ATLAS @ 13 TeV

dilepton mass spectra →



**ATLAS  
EXPERIMENT**  
Run Number: 271048, Event Number: 344127619  
Date: 2015-07-09 03:56:38 CEST





# Event Display of a Candidate Electron-Positron Pair with an Invariant Mass of 2.9 TeV

## CMS @ 13 TeV

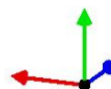


**CMS DP -2015/039**

This performance note shows the event display together with some kinematic quantities for a candidate electron-positron pair with an invariant mass of 2.9 TeV. The background expected from the SM above  $m(ee) = 1$  TeV, 2 TeV and 2.5 TeV for an integrated luminosity of 65 pb<sup>-1</sup> is also stated.

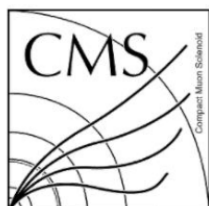
mass range	SM Bkg. Expectation
>1 TeV	0.21
> 2 TeV	0.007
> 2.5 TeV	0.002

CMS Experiment at LHC, CERN  
Data recorded: Sat Aug 22 04:13:48 2015 CEST  
Run/Event: 254833 / 1268846022  
Lumi section: 846

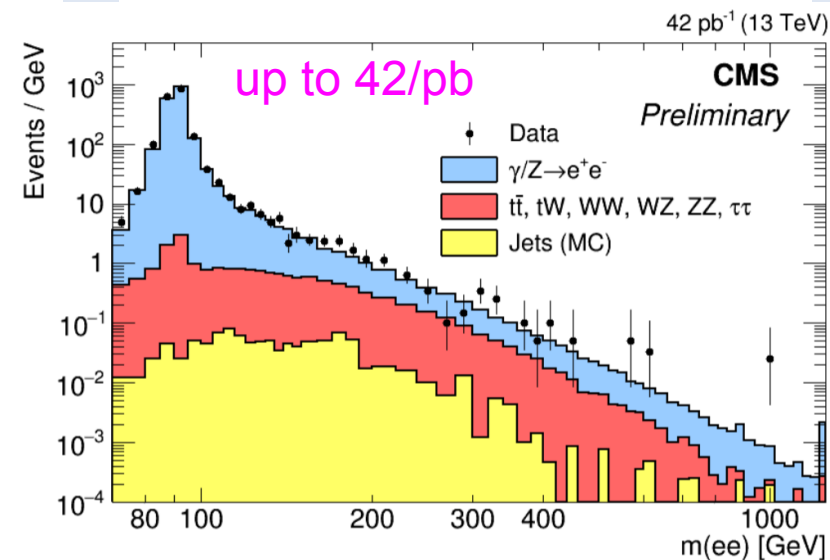


Electron 1,  
pt = 1278.63  
eta = -1.312  
phi = 0.420

Electron 0,  
pt = 1256.20  
eta = -0.239  
phi = -2.741



This event was taken after 42/pb





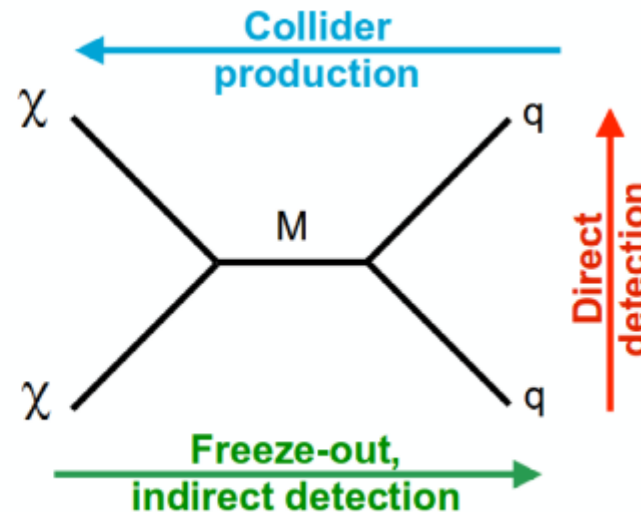
# Dark matter @ LHC

♦ There are three main approaches to detect DM:

- ◉ DM-nucleon scattering (direct detection)
- ◉ Indirect detection (coannihilation)
- ◉ Pair production at colliders

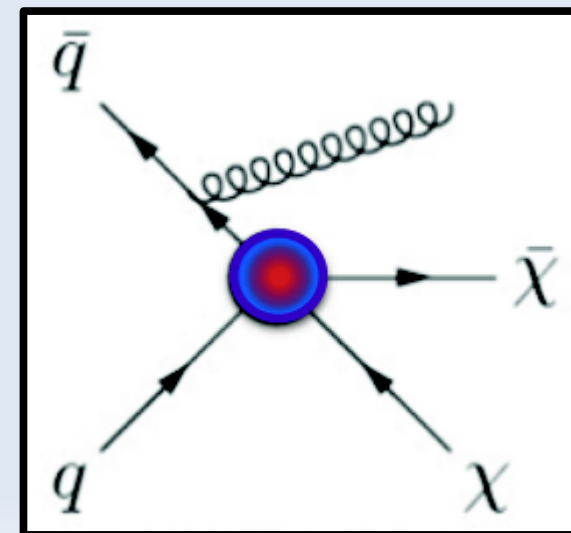
♦ All three processes are nothing but topological permutations of one and the same Feynman diagram:

- ◉ But: how to trigger on a pair of DM particles at colliders?



Majority of BSM searches @ LHC could be interpreted as searches for dark matter because almost all BSM models contain some DM candidate.

mono –  
something detectable & triggerable

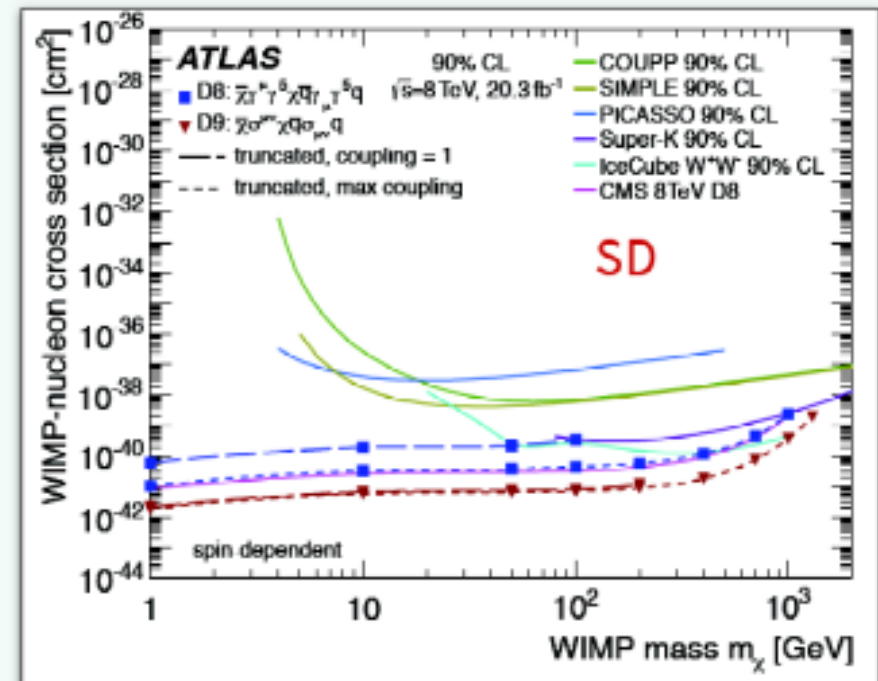
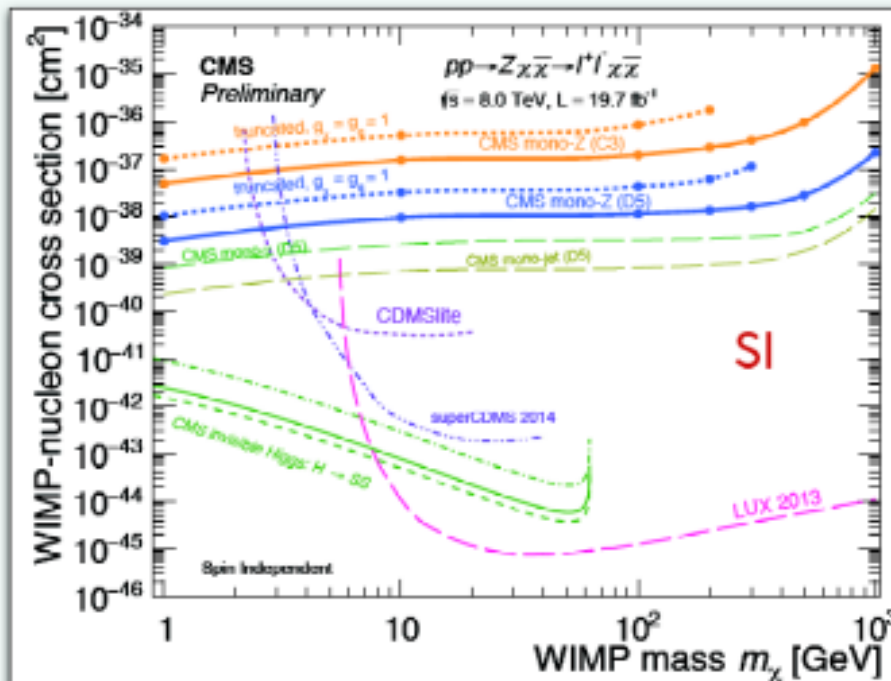


# Dark matter @ LHC

## Translate collider limits to bounds on $\sigma_{\chi-N}$

- ▶ Lower limit on  $M^*$  translated to limits on  $\sigma_{\chi-N}$ 
  - Phys.Rev.D82:116010,2010
- ▶ Collider searches more sensitive at low  $M_\chi$ 
  - And up to medium  $M_\chi$  for spin-dependent interactions
- ▶ Complementarity to direct searches

But if **DM** particles are superweakly interacting eg. **gravitino** than the LHC is the only device that could provide a hint for DM in near future!



# Conclusions

- A detailed search for almost every imaginable symptom of BSM physics was performed on 7 & 8 TeV data by ATLAS and CMS.
- No signal of BSM phenomenon of any kind has been found yet.
- LHC run 2 at 13 TeV has just started.
- The amount of integrated luminosity at 13 TeV is limited but we expect an order of magnitude more up to the end of the year.
- Physics objects are being validated for 13 TeV run.
- Cross sections for BSM physics significantly grow with pp energy.
- More search results are to be expected already this year.
- We have found nothing at 8 TeV, but we will keep searching at 13 TeV!
- LHC is a unique tool to search for DM.

<https://twiki.cern.ch/twiki/bin/view/AtlasPublic>

<https://twiki.cern.ch/twiki/bin/view/CMSPublic/PhysicsResults>

