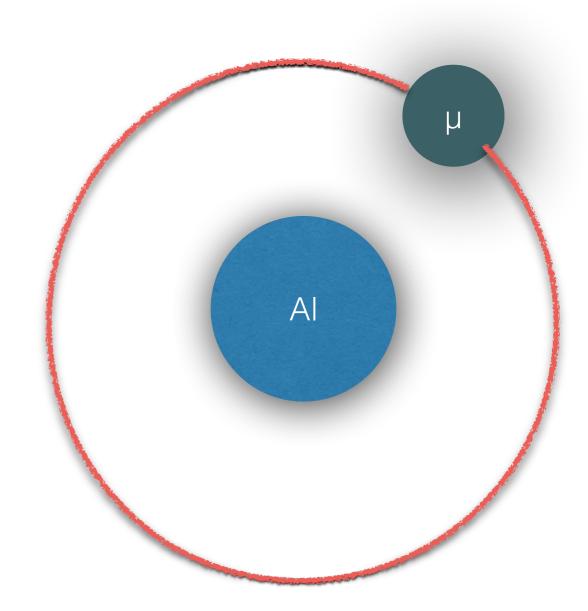
Bound muon decay and its role in New Physics searches



Matter To The Deepest September 16, 2015

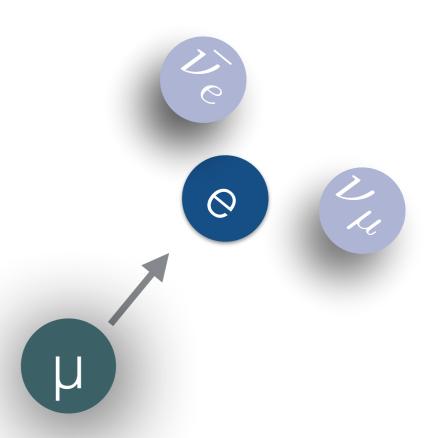
Outline

- Decay in orbit spectrum:
 - central region
 - endpoint region
- Summary



Free muon decay

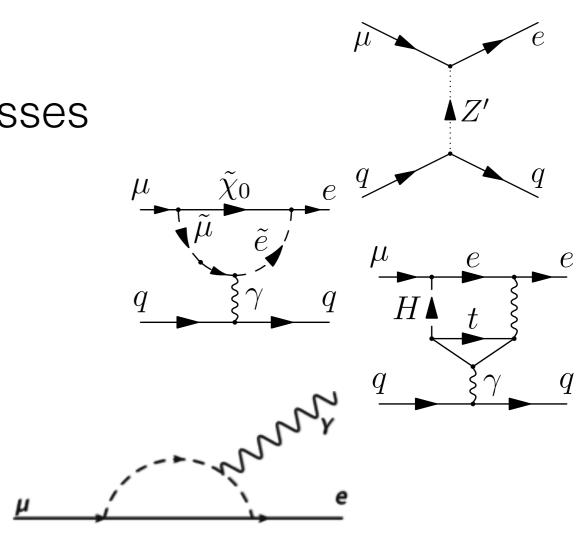
- Well known SM process
- Source of New Physics constraints
- NLO corrections calculated in 1950s
- NNLO corrections are also known
- Only lepton flavour conserving decay modes have been observed
- Anomalous magnetic moment may indicate a need for a NP contributions



Off—diagonal dipole moments

- Similar type of operators may contribute to g-2 and Charged Lepton Flavour Violation (CLFV)
- CLFV is suppressed in SM
- Three interesting CLFV processes
 - $\mu \to e \gamma$
 - muon electron conversion

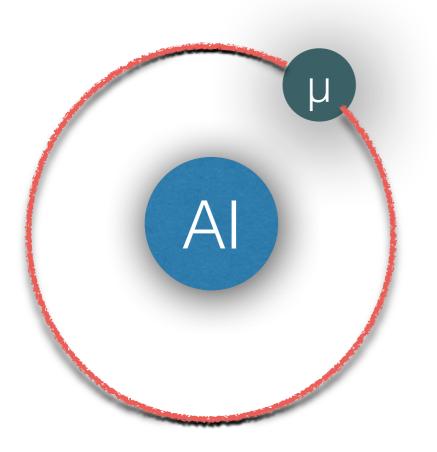
• $\mu \rightarrow eee$

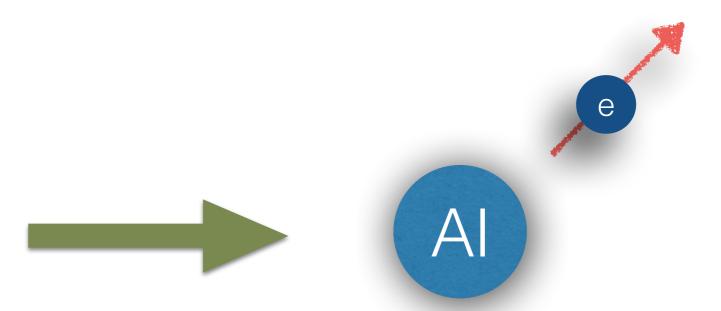


Three processes with bound muons

| Proces | | SM rate | Why important? |
|-----------------------|---|---|---|
| Conversion | $(\mu^- N) \to N + e^-$ | Negligible | Observation indicates New Physics |
| Decay in Orbit DIO | $(\mu^- N) \rightarrow$ $N + e^- + \overline{\nu_e} + \nu_\mu$ | Approximately equal to free muon decay rate | Background to conversion |
| Capture | $(\mu^- N) \to N' + \nu_\mu$ | Depends on Z | Normalization factor for conversion |

Muon electron conversion

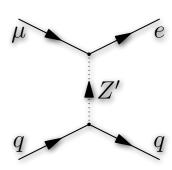




Muon converts to electron without emitting neutrinos Lepton family number not conserved

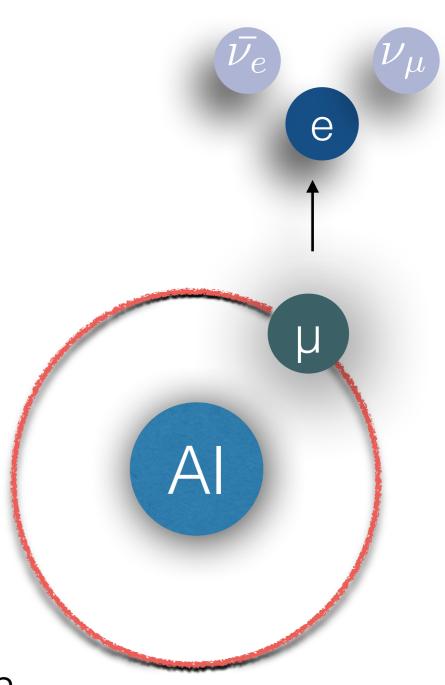
Muon electron conversion

- Clean experimental signature mono-energetic electron
- Current limit on the ratio R of the conversion to the capture $R < 7 \times 10^{-13}$
- Planned experiments expect to improve R by ~4 orders of magnitude, equivalent to probing New Physics scale up to 10 000 TeV!
- Conversion can probe larger class of operators than $\mu \to e \gamma$

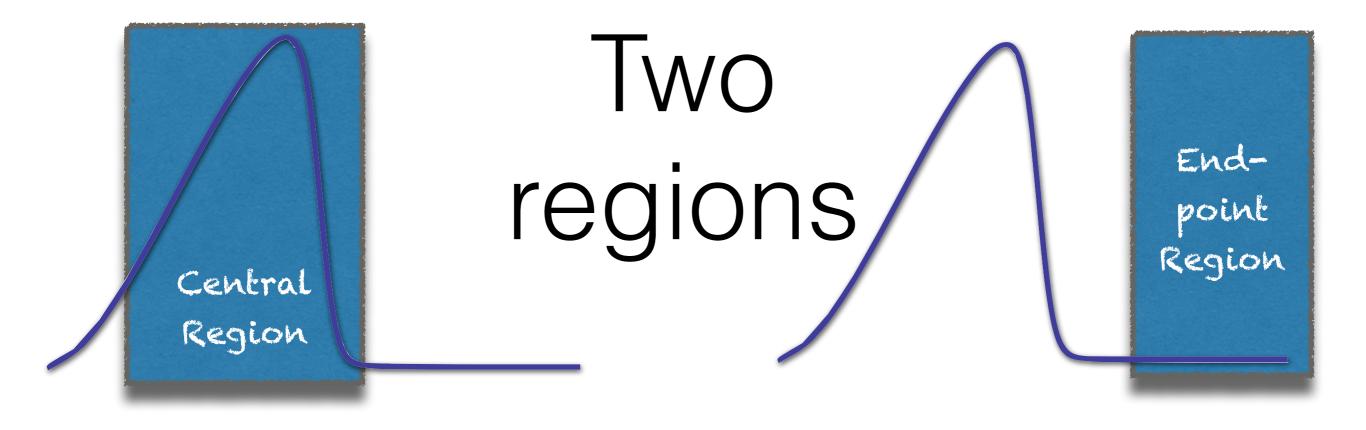


Bound muon decay

- Muon DIO: standard muon decay into an electron and two neutrinos, with the muon and a nucleus forming bound state
- For a free muon, energy and momentum conservation restricts electron spectrum to $E_e < \frac{m_{\mu}}{2}$
- For DIO, momentum can be exchanged between the nucleus and both the muon and the electron



dΓ Free µ dE_e Conversion DIO signal COMET, TWIST, 2009 Mu2E Ee $\frac{1}{2}m_{\mu}$ m DIO Spectrum



Most important effect:

muon motion in an atom

exchange of a hard photon

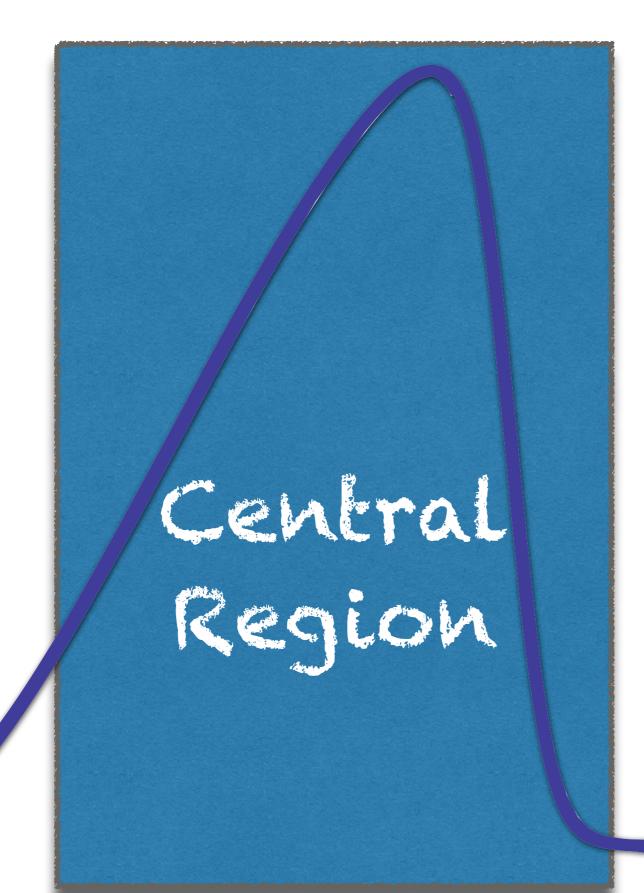
Corrections:

final state interaction

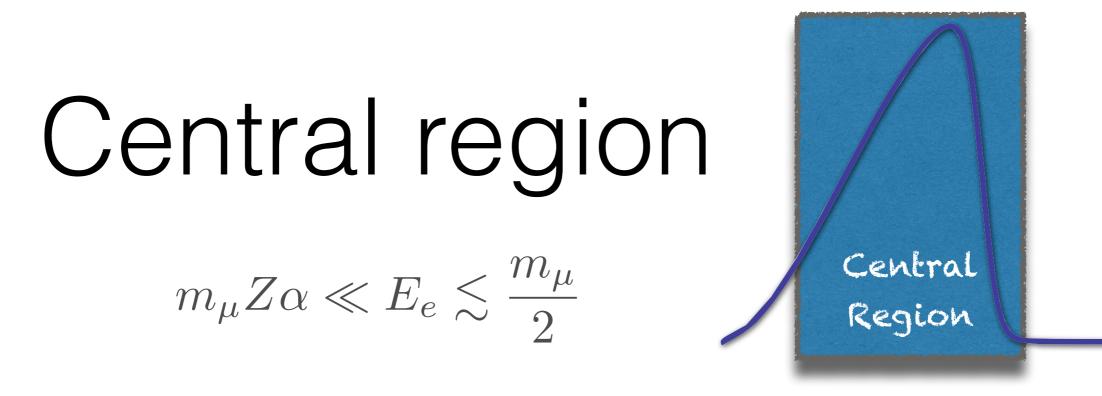
finite size of the nucleus

recoil effects

Radiative corrections!



Also known as Michel Region



- Typical momentum transfer between nucleus and muon is of the order of $m_\mu Z \alpha$
- Requires resummation
- Dominant effect muon motion in the initial state
- Similar problems decays of heavy quarks in mesons

QED shape function

Charged particle in the external field is almost on-shell

• We are interested only in the leading corrections

QED shape function

 Shape function is defined as an expectation value: Momentum

$$S(\lambda) = \int d^3x \psi^*(x) \delta(\lambda - n \cdot \pi) \psi(x) - \text{distribution}$$

• We work in light-cone gauge

 $n \cdot A = 0$

Final state interaction, required by gauge invariance

• Normalization:
$$\int_{-\infty}^{\infty} d\lambda S(\lambda) = 1$$

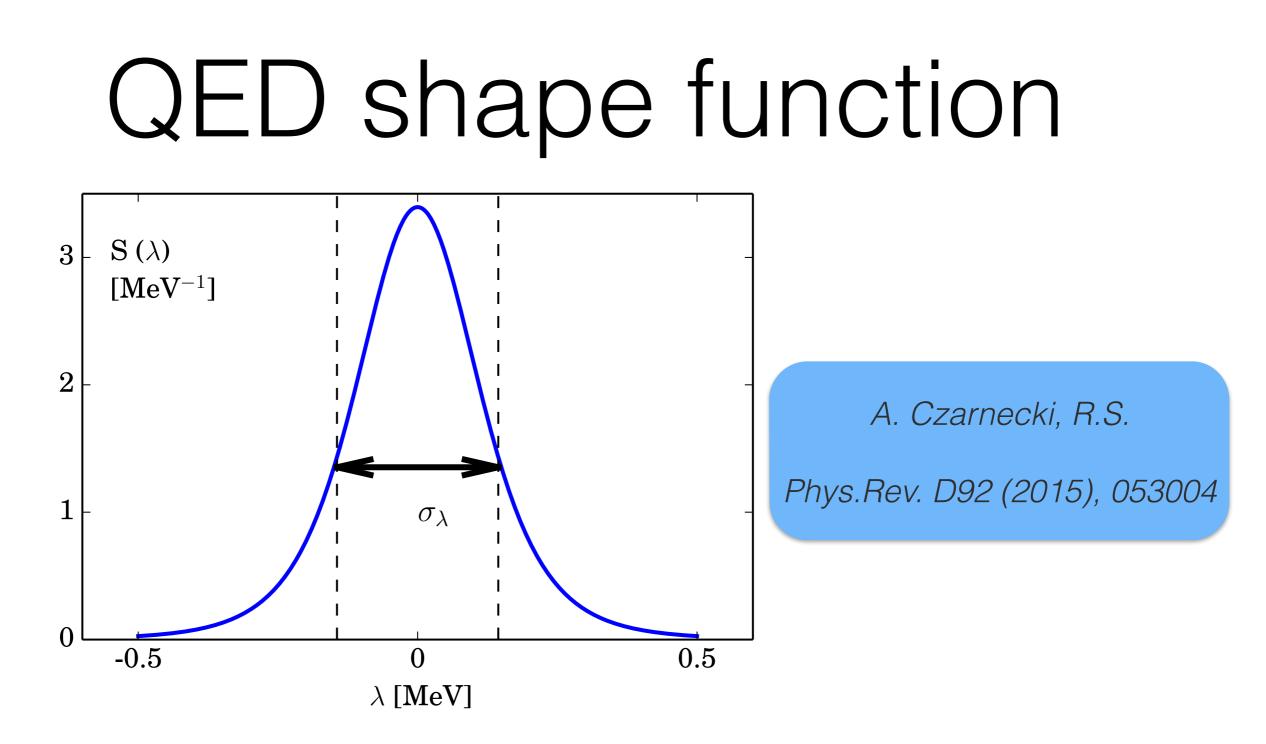
Power counting

•
$$\lambda \sim \frac{p_e^2}{2E_e} \sim m_\mu Z \alpha$$
 (muon momentum in an atom)

- Shape function behaves as $S(\lambda) \sim \frac{1}{Z\alpha}$
- First moment is zero in the leading order

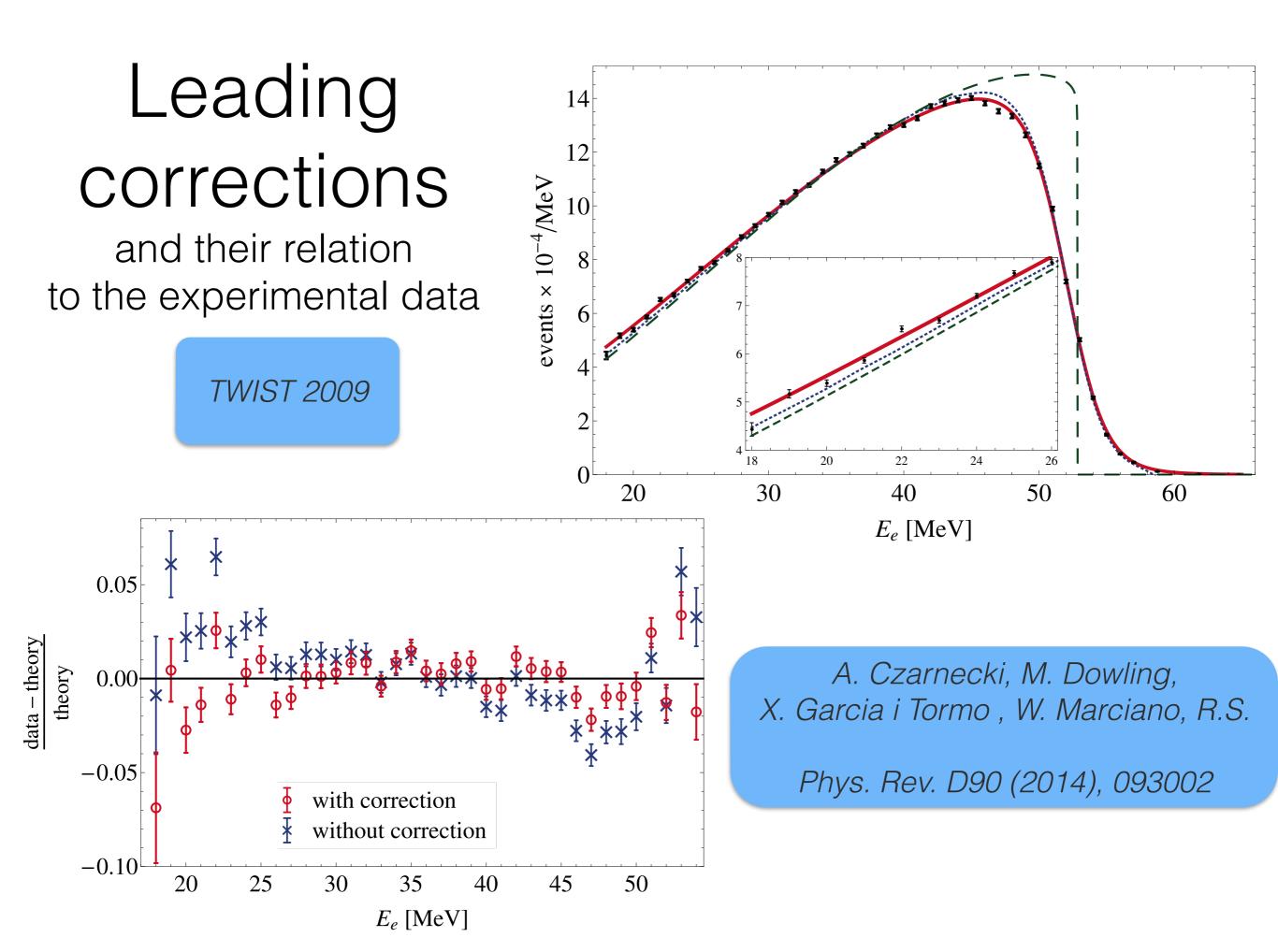
$$\int d\lambda \lambda S(\lambda) \sim (Z\alpha)^2$$

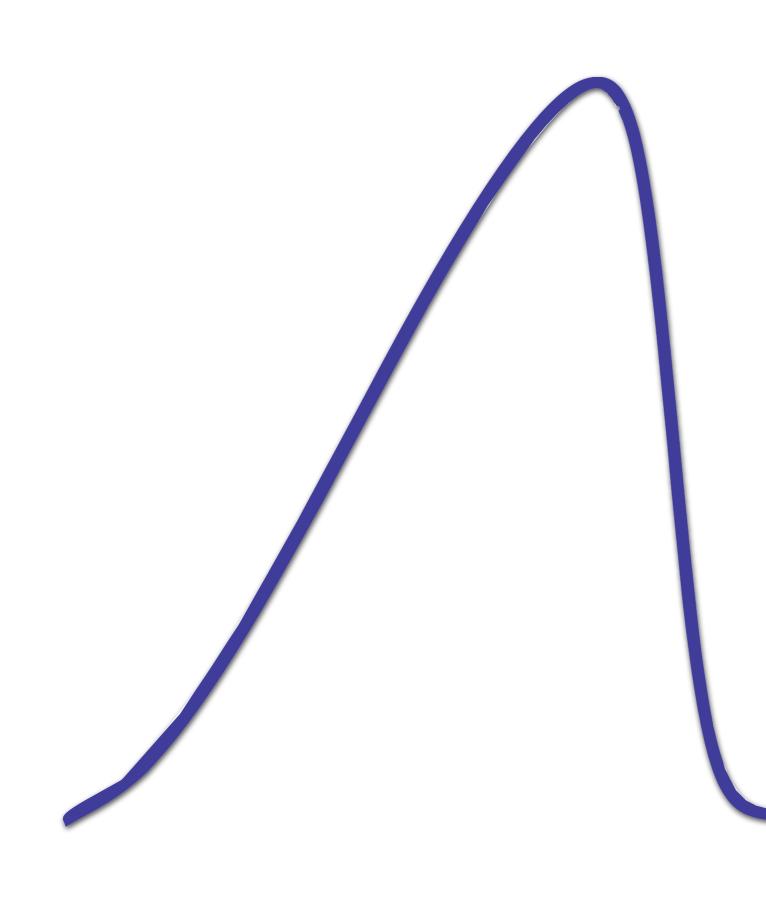
• Second moment $\int d\lambda \lambda^2 S(\lambda) = \frac{1}{3} (m_\mu Z \alpha)^2$

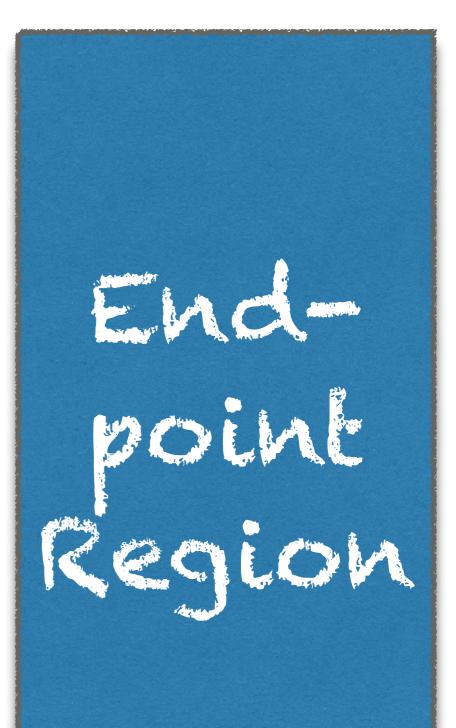


Spectrum can be calculated using factorization formula

$$\frac{d\Gamma_{DIO}}{dE_e} = \frac{d\Gamma_{Free}}{dE_e} * S$$







Endpoint region $E_e \sim m_\mu$

 Typical momentum transfer between the nucleus and the muon is of the order of the muon mass

End-

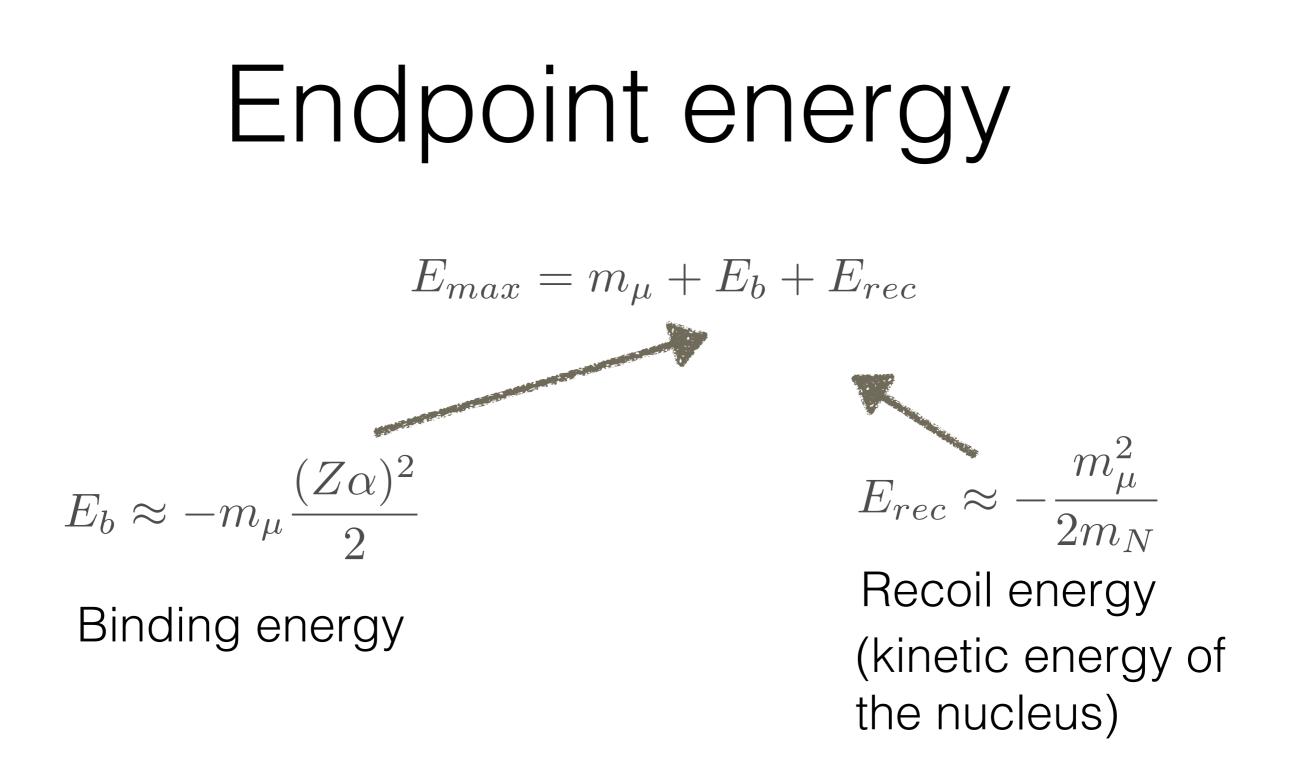
point

Region

 E_{e}

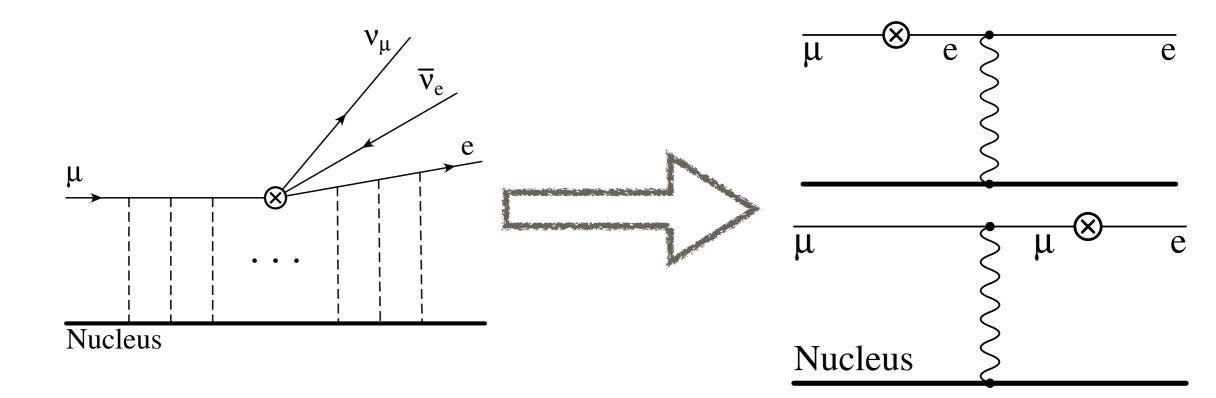
• Both wave functions and propagators can be expanded in powers of $Z\alpha$

$$\frac{m_{\mu}}{\Gamma_{Free}} \frac{d\Gamma}{dE_e} \approx \frac{1024}{5\pi} (Z\alpha)^5 \left(\frac{\Delta}{m_{\mu}}\right)^5 \qquad \Delta = E_{max} -$$



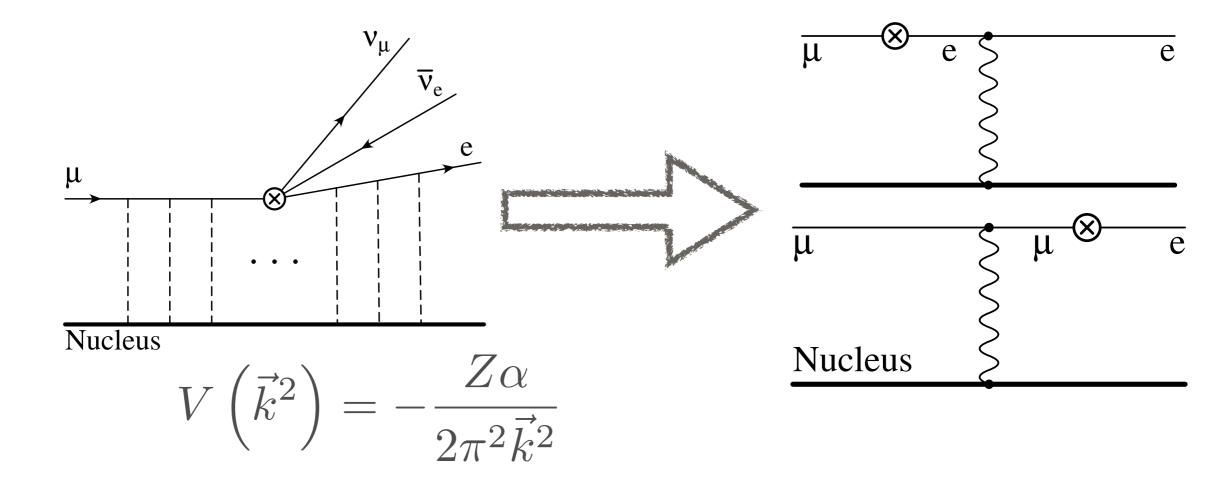
Both corrections decrease the endpoint energy

Endpoint expansion



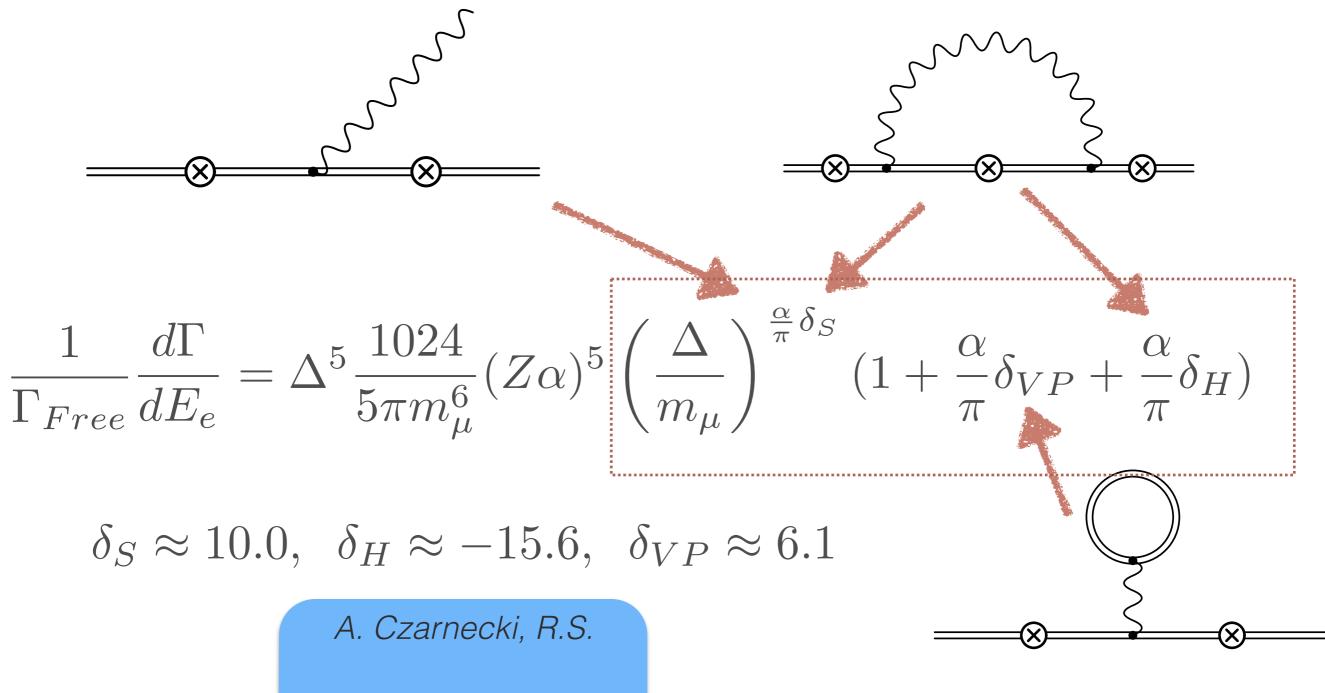
$$\int \frac{\mathrm{d}^{3}\nu}{\nu_{0}} \frac{\mathrm{d}^{3}\overline{\nu}_{0}}{\overline{\nu}_{0}} \delta\left(\Delta - \nu_{0} - \overline{\nu}_{0}\right) \dots \psi \dots \overline{\psi} \sim \Delta^{5}$$

Endpoint expansion



$$\mathcal{A} \sim \psi(0) \times V(m_{\mu}^2) \sim (Z\alpha)^{\frac{3}{2}} \times Z\alpha = (Z\alpha)^{\frac{5}{2}}$$

Radiative corrections



arXiv:1505.05237

Higher order terms

- Expansion parameter is $\pi Z \alpha$, again very similar to the calculations of photoelectric effect
- Higher order terms were calculated numerically; they give -21% correction for a point-like nucleus
- Finite-size nucleus corrections suppress the higher order terms
- Also higher orders in Δ may be required for precise determination of experimental background

Summary

- Searches for rare decays require accurate predictions for the SM background
- TWIST measurement of the DIO spectrum is sensitive to radiative corrections
- Muon DIO spectrum:
 - We have radiative corrections in regions relevant for experiment
 - Ultimate goal is a correction to the spectrum in the whole energy range